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# DISCO-2 – an ambitious earth observing student CubeSat for arctic climate research

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The severe impact of global warming, especially in the arctic region, have a multitude of consequences spanning from sea-level rises and freshening of the ocean, to significant changes to the animal life, biodiversity and species distribution. As the arctic regions are inherently remote and can be both hazardous and difficult to reach, research to improve our understanding of the climate change impact is often limited to short term field-campaigns. Here we present the Danish DISCO-2 student CubeSat mission, designed to meet the growing need for an Earth-observing platform. This mission leverages the rapid advancements in CubeSat technology over the past decades to

overcome the limitations of traditional fieldwork campaigns. DISCO-2 will assist on-going arctic climate research with a payload of optical and thermal cameras in combination with novel in-orbit data analysis capabilities. It will further be capable of performing photogrammetric observations to determine ice volumes from deteriorating glaciers and provide surface temperatures, enabling studies of heat transfer between glaciers and arctic fjords. As a student satellite, the payload capabilities will also be offered to novel student research ideas throughout the mission life time. The modularity and wide range of of-the-shelf-components for CubeSats has facilitated an immense opportunity to tailor this earth observing CubeSat to accommodate specific scientific goals and further provided students at the participating universities with an unparalleled possibility to go from an initial research idea to a running CubeSat mission.

KEYWORDS

CubeSat, arctic climate, remote sensing, machine learning, student involvement, photogrammetry

# 1 Introduction

The Arctic is one of the regions most severely impacted by climate change, with observed warming significantly faster than the global average Rantanen et al. (2022). Improving our understanding of the impact of these changes both locally and globally is a key factor in efforts to minimize and mitigate their influence. Studying the impact of climate changes in arctic regions has proven to be both costly and limited in relation to the physical extent of field-work campaigns, and arctic climate research therefore serves as a prime subject to benefit from remote sensing tools i.e., small satellites. Ever since the first suggestion of the CubeSat standard in the late 90 s and early 2000s to create low-cost space experimentation Heidt et al. (2000), the increased development of CubeSats has created a more affordable and faster way to launch increasingly complex instruments into orbit. This has allowed for numerous successful CubeSat missions from universities, research institutes, and companies. The low cost and relatively short development timescale for CubeSat missions have especially proven to be a benefit for universities to run student-driven missions Murphy et al. (2018); Jallad et al. (2023), as a typical timeline for these projects allows students to be involved in several if not all phases of a CubeSat mission Murphy et al. (2018). Danish universities have been key players in this development ever since the launch of the very first CubeSats in 2003, e.g., AAU CubeSat and DTUsat Alminde et al. (2004) and the continuous development of CubeSats has now enabled these nanosatellites to be used for scientific purposes or as cost-efficient proof of concept missions to test new technology for future missions Kramer and Cracknell (2008). DISCO-2 is a mission of such character and will be one of the most ambitious student CubeSats. The DISCO-2 mission builds on the experience gathered from the first CubeSat mission of Aarhus University, Delphini-1, the DISCO-1 mission, and by extending the mission with a scientific goal of contributing to arctic climate research the mission ensure a long-term application and commitment after the launch. DISCO-2 is the second CubeSat from the DanIsh Student CubeSat prOgramme DISCO, which is a collaboration between three Danish universities Aarhus University, IT University of Copenhagen, and University of Southern Denmark. DISCO-2 is a 3 unit + (3U+) CubeSat carrying a payload consisting of 2 optical cameras (1 "high-resolution" and 1 "wide-field"), a thermal camera for ground temperature measurements, and a System-on-Module (SoM) board enabling on-board data sorting/quality control using machine learning.

# 2 Mission objectives

The main objective of the DISCO program and hence also the DISCO-2 satellite is to give students at the participating universities hands-on experience with a wide variety of skills and know-how sought after within both the space sector and industry. A main priority for DISCO-2 is therefore to have student involvement in all aspects of the satellite; from the very early conception of mission objectives to daily operations and data users. This has led to several smaller groups of students running different parts of the mission, i.e., structure design, harnessing, printed circuit board (PCB) manufacturing, software development, integration, testing, operations, etc. By providing the opportunity to gain valuable practical experience with several aspects of satellite space technology and research, the DISCO program will make the participants more attractive to the high-tech and space industry. The massive student involvement is also an attempt to secure a steady pipeline of startups to the Danish ESA Business Incubation Centre. The early stages of the scientific motivation for DISCO-2 were also conceived by students and then further refined in collaboration with scientists at Arctic Research Centre and iCLIMATE at Aarhus University. The funding level for DISCO-2 is 400000 Euros (EUR).

## 2.1 Scientific motivation

DISCO-2 is devised as an Earth Observation (EO) platform to support ongoing arctic research conducted at participating universities. In particular, the scientific motivation is based on research conducted in northeast Greenland, where interactions between glaciers and fjord systems are already being studied using ground truth monitoring equipment Rysgaard et al. (2022) and data retrieved during fieldwork campaigns. Adding a dedicated EO platform in the form of a CubeSat will enable more frequent data acquisitions covering larger areas. To fully utilize the potential of the small CubeSat format it is therefore critical that the satellite will be able to monitor using both thermal and visible images.

#### 2.1.1 Glacier-ocean interactions

To utilize an EO CubeSat for supporting ongoing arctic research in northeast Greenland, a combination of visible and thermal imaging is preferred to investigate a variety of interactions between glaciers and fjords, i.e., temperature gradients, heat dissipation and transport concerning surface temperatures and glacial deterioration. All of these interactions are important aspects to improve our understanding of ice-ocean interactions and the large-scale effects of climate changes in northeast Greenland.

Thermal imaging from an EO satellite will enable large-scale observations of surface thermal gradients in fjords. These gradients will enable further analysis of the water flow of fjords and hence be beneficial for the investigation of water circulation and analysis of heat exchange in northeast Greenland fjord systems. Using daily and long-term tracking of glacier surface temperatures will also enable further analysis of glacial deterioration and the interactions and effects between glaciers and their surroundings. Additional visible observations can further improve the investigation of both ocean surface flow and the deterioration of glaciers, as high-resolution imaging of the fjords will enable the tracking of smaller icebergs and possibly correlate the drift of these with thermal imaging to investigate surface flow. A combination of thermal and visible imaging will also enable the investigation of the effects of melting icebergs on the surface temperature in fjord systems.

With an EO CubeSat with thermal and visible observation capabilities, we also intend to investigate subglacial discharge plumes to improve our understanding of the ice-ocean interface, as the ocean effect on glaciers and ice-ocean interactions are still uncertain Everett et al. (2021). As subglacial discharge plumes are occurring in hazardous and remote areas, they represent a prime example of an ice-ocean interaction that will benefit from remote sensing observations. Observations of plumes are critical to understanding heat transfer between ocean and ice while they also play a role in the freshwater addition to the arctic fjords and potentially general freshening of the ocean with the accelerated temperature trend in the arctic Hewitt (2020).

Utilizing high temporal observations with both visible and thermal imaging of glacier fronts will enable a better understanding of the surface water-mixing which happens when positive buoyant freshwater from the plumes moves towards the surface layers of the saline fjord water. In addition, visible imaging can provide a usable tool when investigating the source of plumes i.e., change in extent of supraglacial lakes.

Given the remote nature of northeast Greenland, a dedicated EO platform will enable a high frequency of observations, creating a higher temporal resolution of thermal and visual changes on the ground, which will facilitate investigations into seasonal changes in heat accumulation. Recent Landsat observations at twilight, i.e., wintertime in the Arctic, have also revealed the applicability of thermal imaging in low-light winter conditions, to map and monitor changes in sea ice during polar night Scambos et al. (2024).

## 2.1.2 Photogrammetry

A key feature of DISCO-2 will be to investigate the ability to regularly perform photogrammetric imaging of large glaciers by utilizing a high-resolution camera in coordination with the attitude determination and control system (ADCS) to create 3-dimensional imaging of large-scale structures, i.e., glacier fronts (Ghuffar, 2018). By taking advantage of the high frequency of passes and acquiring multiple images at each pass, the changes in the volume of a glacier front can be analyzed over time. This will provide a tool to examine the volumes of ice that deteriorate from glaciers on seasonal and large single events.

# 2.2 In-orbit analysis

During passes of arctic areas, DISCO-2 have the ability to capture a very large number of high resolution images. The relatively low data rate and transfer window of the downlink create a potential bottleneck for images to be sent to Earth. To mitigate this the DISCO-2 payload is equipped with an image processing unit (IPU) inspired by a predecessor system on DISCO-1 Bayer et al. (2024) that allows for on satellite image analysis. The IPU consists of a flexible software image processing pipeline, which runs on a CPU and hardware NPU machine learning accelerator. The IPU is capable of standard image processing, and also Machine Learning (ML) tasks such as image identification, segmentation and classification. This allows DISCO-2 to analyse images at the rate of aquisition, and transmit only a small subset that match a given criteria back to Earth. Analyses range from simple tasks such as selecting correctly exposed images to more complex ML tasks. Currently in development are a snow/cloud discriminator that uses ML to reject images with over 30% cloud cover and an High Dynamic Range (HDR+) image enhancement module. The image processing pipeline is modular and can be re-configured in orbit, so the platform allows for a wide range of future in orbit image processing projects depending on scientific need and student interest. The development of the IPU is itself an active research project in Embedded Machine Learning and edge computing.

# 2.3 Additional science

In addition to supporting ongoing Arctic research, DISCO-2 will also be available for general-purpose EO. As a student-driven CubeSat DISCO-2 will be applicable to perform observations for novel student research ideas and provide students with first-handexperience in conducting EO. As a side effect of the arctic focus, the constraints on the orbit result in a practical total earth coverage with an increasing frequency of close passes as the orbit reaches the poles. This will further enable relatively rapid response time observations, making DISCO-2 able to perform observations of sudden large-scale events, for example, large changes in the landscape, thermal imaging of volcanoes, etc. These additional science cases will support the more general aims of the DISCO programme to showcase STEM subjects to a general audience and enabling students to get experience with space research.

# 2.4 Student involvement

The development, production and operation of DISCO-2 is relying on continuous student involvement from Aarhus University, Southern Danish University and IT University of Copenhagen. Currently more than 50 students ranging from 1st year bachelor students to master students have been involved with DISCO-2, and more students are expected to join for the next steps of DISCO-2. The students come from various department in example physics, engineering, computer science and geo-science. The participating students are also exposed to a variety of options to participate in space related coursework outside of DISCO-2 and several have participated in external courses hosted by ESA. The students are arranged in work groups with different focus i.e., groundstation, operations, test and assembly, harness, structure, Public Relations, etc. Together with staff members each work group identifies and resolves relevant tasks and have regular meetings both within the group and across the different work groups to both get a progress overview and identify tasks that needs input from multiple work groups. Staff members and some work group leaders are also appointed as section leaders and have status meetings every fortnight where larger issues and impediments can be discussed and mitigated/resolved.

# **3** Scientific requirements

To meet the scientific goals of DISCO-2, the satellite, particularly for the payload, will have to meet a set of requirements specifying the optics, hardware, and orbit of the satellite. In this section, we focus on describing some of the key requirements constrained by the scientific aims of DISCO-2.

## 3.1 Resolution

To obtain high-resolution images with high revisit frequency of the northeast Greenland fjords and glaciers, the cameras will require the ability to resolve typical northeast Greenland fjords and cover large sections of the fjords in a single frame. The most limiting feature the camera systems will need to meet is the ability to resolve surface sediment pools, caused by subglacial discharge plumes with positive buoyancy reaching the surface of the fjords and the glacier fronts. The plume pools are typically on the scale of a few hundred meters wide Mankoff et al. (2016); Jackson et al. (2017), while the ice cliff height above sea level for grounded ocean terminating glaciers can reach approximately 100 m Parizek et al. (2019). The sizes of the fjords vary significantly, with a width on the scale of a few kilometers at the head of the fjords, while the length of the fjords varies on the scale from tens to hundreds of kilometers Morlighem et al. (2017). This provides a requirement for the cameras to have a Ground Sampling Distance (GSD) on a sub-hundred-meter scale to resolve the glacier fronts and surface sediment pools and have entire fjords within the Field of View (FoV).

## 3.2 Thermal imaging

The thermal observations must be able to cover northeast Greenland fjords on the scale of kilometers to hundreds of kilometers to observe surface temperature gradients along and across the fjords. Similar to the optical camera requirement, the spatial resolution of the thermal camera should allow for tracking changes in the surface temperature of the sub-glacial discharge surface sediment pools. Given the longer wavelengths required to observe in the thermal range, the ability to sharply resolve the extent of the pools is not required. A spatial resolution requirement of 200 m per pixel will therefore satisfy the ability to observe temperature changes from surface sediment pools. In addition the thermal camera must be sensitive enough to observe changes in surface temperature with sufficient detail, as surface temperatures in Greenlandic fjords can experience significant changes on short temporal scale and physical extend Mortensen et al. (2014). Therefore the thermal camera must be able to resolve temperature gradients across an image with a precision better than a single degree Celsius.

# 3.3 Pointing

To optimize the satellite's ability to perform photogrammetric observations of single specified targets during passes, the pointing of the satellite must be accurate enough to keep a target (glacial front, large iceberg, etc.) within the FoV of the cameras during a pass. This can translate to a pointing requirement for the ADCS to have the ability to control the orientation of the satellite during a pass to enable observations of targets at an array of different angles, which benefits the photogrammetric analysis. This means we need sufficient pointing accuracy to acquire a minimum of 5 observations of a specific target during a single pass. The targets are not required to be located on the same pixels for these 5 images. In order to maximize the amount of data for each target, the individual targets will be observed for multiple days. The ADCS should together with data reduction algorithms be capable of tagging the images with a set of rational polynomial coefficients (RPCs) that are needed for the photogrammetric analysis. RPCs relates ground coordinates to the corresponding image pixel coordinates using a set of polynomial (Sohn et al., 2003).

## 3.4 Orbital requirements

The geographic location of the main scientific research area places a requirement on the inclination of the satellite orbit. To optimize the temporal resolution (i.e., frequency in passes over northeast Greenland) the orbit of DISCO-2 should be a polar orbit. Adding a soft requirement of the orbit being sunsynchronous provides a benefit of revisits at specific points the same local time each pass, enabling a better comparison of images acquired at different days. The altitude of the orbit must also be able to keep DISCO-2 in orbit for two full seasons to cover seasonal variability. Between the lowest and highest latitudes of Greenland, the DISCO-2 satellite will have between 2 and 5 passes pr. day



FIGURE 1

Ground track for near-polar orbiting satellite with altitude of 510 km and inclination of 97.4° for a 24 h duration. The ground tracks were made using the Skyfield python package Rhodes (2019).



#### FIGURE 2

DISCO-2 in "flight mode" with fine sun sensors, solar panels and UHF antennas deployed. The backside of DISCO is covered with solar panels similar to the panels shown on the deployed solar panels. DISCO-2 is a 3U+ CubeSat, with outer dimensions  $10 \times 10 \times 37.65$  cm (prior to deployment of antennas, sun sensors and

solar panels). The outer dimensions of the box shape is  $10 \times 10 \times 34.05$  cm and the UHF antenna is placed in the "can" shape on the bottom of the satellite, with radius 3.2 cm and height 3.6 cm.

assuming a viewing zenith angle within 45°, with the number of good passes increasing with increasing latitude. The calculations are based on an orbital altitude of 510 km. The ground tracks for a near-polar orbit at an 97.4° inclination and 510km altitude is shown in Figure 1. Here its clear that the ground tracks closes in at the poles which results in the increase in potential daily passes with increasing latitude. During the mission DISCO-2 will observe the arctic during both polar day and night. During the months of polar night DISCO-2 will be able to benefit from thermal observations, as demonstrated with recent Landsat special request data collection program Scambos et al. (2024).

# 4 Mission design

In collaboration with industry, the students designed the DISCO-2 CubeSat to match the standard Umax format (outer dim.) with flight-proven subsystems. The satellite bus parts are provided by Space Inventor, while the structure, payload, and the associated payload bracket were designed and selected by the students and staff of the DISCO-2 team. A 3D Computer-aided design (CAD) drawing of the DISCO-2 satellite is shown in Figure 2. DISCO-2 has been designed to have the cameras and directional S-band antennas on the same side to allow for both observational and high-rate data-downlink capability simultaneously. All main avionics are connected through a redundant control area network (CAN) bus applying CubeSat Space Protocol (CSP 2.01) for both internal communication between the subsystems and space-to-ground communication.

<sup>1</sup> https://github.com/libcsp/libcsp/tree/libcsp-2-0

TABLE 1 Key parameters for the cameras DISCO-2. The Field of View (FoV) and Ground Sampling Distance (GSD) is calculated theoretically (Theo) based on an orbital altitude of 510 km. Additionally the GSD for the optical cameras are also found through empiric measurements (Emp). The optical cameras are from Alvium, and the thermal from FLIR.

Model	Sensor size [px]	Pixel size [µm]	Focal l. [mm]	FoV [km]	GSD Theo   Emp [m/px]	Min exposure [µs]	Max frame rate [fps]
1800 U-811	$2848 \times 2848$	2.74	70	57 × 57	40 — 54.38	45.131	51
1800 U-507	$2464 \times 2056$	3.45	8	542 × 452	440 — 408.18	50.667	34
Boson	640 × 512	12.00	36	109 × 87	170 — -	-	(60 Hz)







#### FIGURE 4

Test image from thermal camera. Left image shows a full frame test image of the payload PCB ( $\approx 10 \times 10$ ). Right side shows the temperature difference across a single row of pixels marked with the blue line on the thermal image. The zoomed in part shows the ability of the thermal camera to resolve thermal differences within 1°C.

# 4.1 Payload

DISCO-2 will carry a payload consisting of 3 cameras and an IPU comprising 2 redundant System-on-modules (SoM) board for in-orbit data analysis and machine learning applications. The three cameras consist of two visible light cameras with different resolutions and FoV and a thermal camera. The key parameters for each camera are listed in Table 1. Test images from the low resolution optical camera and the thermal camera are shown in Figures 3, 4. Both visible cameras have global shutter. Focus for the



cameras will be set to "infinity", which in practice is a point beyond their hyperfocal distance, which are  $\approx 115$  m for the high resolution optical camera, and  $\approx 2.5$  m for the low res optical camera. The camera control software embedded in the payload will manage the cameras and facilitate a variety of functions. This enables adjustment of the cameras exposure time, gain, number of images and image interval depending on observation requirements for the specific task. The optical cameras sends 12 bit raw unprocessed images to the image processing pipeline which can then run any additional further configurable processing via modules, for instance removing lens distortion, compression, greyscale setting gamma, crop, tile, ML analysis, etc. The control software will also enable advanced postprocessing like sub-pixel shifting to improve image quality by resolution enhancing through software Stankevich et al. (2020). The two visible cameras and lens combinations are chosen to accommodate the scientific requirements, and to optimize the duty cycle for the payload concerning power consumption and required computational power. The payload selection process and general design of the payload area are also heavily constrained by the size limitation of the 3U+ CubeSat stated in the design specification Johnstone (2022).

To be able to keep the cameras constrained to a volume of  $10 \times 10 \times 10$  cm, a team of students designed a payload bracket by utilizing the topology optimization generative design feature in Autodesk Fusion 360. The fixed point and fixed geometries are defined as the structure fixtures by bolts, the payload fixtures and masses (3 cameras, SoM module and magnetorquer), to make room for cabling and heat management. The design is optimized for 3 times max vibrational load (75 g) in all 3 degrees of translational directions, as the orientation in the rocket is unknown. The material used is aluminium due to good weight-to-strength ratio combined

with very good heat transfer properties. The resulting organic design is optimized for this load case, but cannot be manufactured using traditional manufacturing processes. Hence the bracket is 3D printed in aluminium using a Selective Laser Melting (SLM) printer. The designed payload-bracket with cameras is shown in Figure 5.

### 4.1.1 Payload software

The software used to control the payload runs on dedicated processors independently from the flight control onboard computer (OBC). The payload computer has two independent SoM's, each with three different types of processors; a 4-core A53 running Yocto-Linux, a low-power M7 core with FreeRTOS, and a neural processing unit (NPU) that can be accessed from the Linux system. The high-powered A53 cores are used for camera control and image processing and are kept suspended when not in use. The M7 core runs a scheduler with the responsibility for booting the Linux system when it needs to be accessed, sending commands to the payload, and powering it off afterward to minimize power usage. Setup and control of the scheduler is done using a domain-specific language (DSL) developed for it, called Proc. It uses the CSP and the parameter system, Param, provided by Space Inventor which is also used in the other components of the satellite. The DSL allows coding both simple and complex procedures with flow control and branching, based on parameters from various components. It makes it possible for example, to wake up the power-hungry A53 cores and start taking images when entering a geofenced area based on coordinate parameters from the GNSS module. To better utilize the downlink bandwitdh, the DISCO-2 IPU is capable of in-orbit image processing, for example, compression, image classification or discarding of bad images automatically. Depending on the project images are taken for, the requirements for onboard processing will differ. For this task, we introduce an image processing pipeline (DISCO Image Processing Pipeline, DIPP) with modularity and configurability as its key attributes. New modules can be uploaded to the satellite as compiled.so (shared object) files, along with configurations that chain modules together, where the commands for taking images include the configuration name, for identifying the pipeline, it should be passed through. The different repositories with DIPP and other modules are available on GitHub<sup>2</sup>.

#### 4.1.2 Global navigation satellite system (GNSS)

The payload area will include a Skytraq Orion B16 GNSS module with modified firmware to provide accurate positioning, which will be recorded along each image acquisition to obtain accurate time and position stamps. The Orion B16 GNSS module has a positioning accuracy of 2.0 m Circular Error Probability (CEP), velocity accuracy of 0.1 m  $\cdot$  s<sup>-1</sup> and a time accuracy of 5 ns. This GNSS module is specifically developed for CubeSats and has a good flight heritage. The payload can also be queried through CSP from the bus modules to provide location information from the GNSS module at any given time.

<sup>2</sup> https://github.com/discosat

# 4.2 Onboard computer (OBC)

The OBC handles the flight software on the satellite including the ADCS. The flight software and the ADCS are separated into two plates placed on each side of the module. The flight software platform consists of two independent ARM cortex-M7 modules, each with a separate power supply, interfacing, and storage. The dual architecture makes it a suitable choice for hot/cold redundancy solutions. The OBC further enables 64 GB of storage for user data and collects and stores housekeeping data.

## 4.3 Power system

The power system on DISCO-2 consists of body-mounted solar panels, 2 deployable solar panels, a maximum power point tracker (MPPT), a battery, and two power distribution units (PDU). Azure Space Cell Assembly 3G30A were selected. Among other features they include a 30% triple junction GaAs, to optimize power production at different wavelengths. The deployable solar panels will be released after DISCO-2 has been launched from the CubeSat pod using a flight-proven hold-down-release mechanism. The MPPT ensures optimal operating voltage is achieved for the individual solar cell arrays during different temperatures and irradiance levels. The MPPT is further supplied with a current and over-voltage protection for the battery output, to prevent overcharging and a pass-through mechanism allows for the satellite to pass solar output directly to the battery in case of battery power dissipation, potentially preventing complete loss of the satellite. The battery pack on DISCO-2 is an 8-cell Lithium-ion battery pack in a 4s2P configuration with a nominal capacity of 92 Wh. It has connectors for both soft and hard kill switches. It further carries an always-on-ultra-low-power real-time clock to provide timer continuity during satellite shutdown. Each PDU has 4 redundant channels (8 single) and dual redundant controllers to control all 8 switches.

## 4.4 Communications

To enable communication between DISCO-2 and the ground station, it is equipped with both low-rate UHF and high-rate S-band radio channels. The UHF radio system features dual hot redundancy and is tailored for nanosatellites with a Low Earth Orbit (LEO) lifespan of up to 5 years. The Telemetry, Tracking, and Command (TTC) transceiver connects to a pair of UHF monopole antennas, forming an omnidirectional antenna system with an approximate data rate of 9.6 kbps. These antennas deploy from the satellite after it is released from the pod. DISCO-2 is equipped with both receiving and transmitting S-band antenna patches on the Earth-facing front. The S-band radios operate in the frequency sub-bands of 2025-2210 MHz (uplink) and 2400 - 2450 MHz (downlink). It is developed to allow for high-rate communications for most satellite passes with a half-power beam width of 80°. The S-band downlink radio provides a high data rate channel for transferring the DISCO-2 payload data with an approximate downlink data rate at 1 Mbps. The S-band uplink radio offers a TTC backup channel and allows for experimental use.

Measurement	Value	Unit
Azimuth width (3dB)	2.788	deg
Elevation width (3dB)	2.076	deg
Antenna gain	37.3	dBi

#### 4.4.1 Ground segments

The ground segment for DISCO-2 consists of an S-band dish antenna for image payload data download located at SDU and UHF antennas at AU, ITU, and SDU for TTC. The antenna systems are mounted on an azimuth/elevation-controlled rotor platform to allow tracking of the CubeSat during its passes. The S-band antenna at SDU consists of a 3 m dish reflector with a circularly polarized feedhorn antenna in series with a Low-Noise Amplifier (LNA) and a Software-Defined Radio (SDR) receiver. The far-field antenna radiation pattern has been measured by the company QuadSAT concerning key performance parameters found in Table 2. The offset of the main lobe is less than 0.1° and negligible since the rotor resolution is about 0.1°. The receiver is an Ettus USRP B200 SDR. The S-band uplink channel will furthermore be equipped with a power amplifier and a switch to control the change of reception and transmission modes of operation. The UHF radio channel utilizes two cross-polarized Yagi antennas, one for uplink and one for downlink. The antennas are circularly polarized which results in an acceptable polarization loss of up to 3 dB. The UHF radio receivers are two HackRF One SDRs in series with two LNAs. The S-band dish antenna is mounted on azimuth/elevation rotors that makes it possible to follow the satellite path.

# 4.5 Attitude determination and control system (ADCS)

#### 4.5.1 Reaction wheels

The main actuators on DISCO-2 are the reaction wheels. The four reaction wheels are arranged in a pyramid configuration, overlapping cover the 3 degrees of rotational freedom as shown in Figure 6. This redundancy allows continuous attitude control in the case of one reaction wheel failure. Over time, the momentum of the reaction wheels will increase due to environmental torques, i.e., drag force, radiation pressure, and the interaction between spacecraft battery and geomagnetic field. Eventually, the reaction wheels will be saturated and not able to counteract the impacts of torque. To de-saturate the reaction wheels, magnetorquers are used Wafi et al. (2022).

#### 4.5.2 Magnetorquers

The magnetorquers are essentially electromagnetic coils that interact with the geomagnetic field when current runs through them. They provide an external torque, making it possible to dump momentum build up in the reaction wheels. They are not useful for precise pointing, but proves useful for the initial de-tumbling phase of DISCO-2 immediately after deployment from the CubeSat pod.



### 4.5.3 Fine sun sensors

On each fine sun sensor, two perpendicular slits are carved into the compartment. On the bottom of the compartment, a reticle pattern is carved such that an electrical signal is produced when light hits the bottom (Chen and Lerner, 1999). The fine sun sensors output a sun vector and together with the current position of the spacecraft and an ephemeris of the sun the current attitude can be estimated to within 1° with 1 $\sigma$  accuracy when the sun is within a cone of 35°. DISCO-2 is fitted with 6 sun sensors, one on each face of the CubeSat.

#### 4.5.4 Magnetometers

Magnetometers can be modelled as inductors in an inductorcapacitor (LC) circuit. The effective inductance in such a circuit depends on the ambient magnetic field through the coil and thus the magnetometers can estimate the strength of the magnetic field surrounding the CubeSat. Together with the position of the spacecraft and the international geomagnetic reference field, the attitude can be estimated within 1°.

## 4.5.5 Multiplicative extended Kalman filter (MEKF)

The onboard attitude determination is estimated through a MEKF. The filter uses the available measurements from the sensors to update the attitude and state of the spacecraft as well as the associated uncertainties. The filter utilizes knowledge of the control torques to propagate the latest estimate of the attitude. The sensor measurement uncertainties can be chosen freely and the process noise can be modified. The initial process noise is used when propagating the attitude estimate between observations (Lawrence Fallon, 1999).

#### 4.5.6 Linear quadratic regulator (LQR)

The LQR considers the current state, the attitude, and the angular velocity and determines the control voltage on the reaction wheels (Helmy et al., 2023). The LQR input is a user-defined gain, allowing the operators to tailor the behavior of the LQR.

# 5 Harness

Three different wire types are used in the satellite. The first type is a 1 mm diameter wire that requires a bending radius of greater than 3 mm. This wire type is the most commonly used in this satellite. The second wire type is a 3 mm diameter wire with a minimum bending radius of 15 mm. This wire is used for the UHF and S-band antennas. The minimum bending radius is determined based on the dynamical stress impacting the wires due to vibrations during launch. The third wire is an interconnecting PCB to reduce mass and complexity compared to wire-splicing. The PCB connects the IMU, OBC, and one of the PDUs.

Due to a lack of power outputs from the PDUs, some components must share a power output which could be achieved using wire splicing. Generally, splicing wires is not optimal, as it risks developing into a failure point. Therefore, the payload PCB is used to divide the power inputs into multiple outputs for the power-sharing components in the payload as well as the fine sun sensors on top of the satellite. However, one wire is spliced to share power between the two fine sun sensors not placed on top.

# 6 Mission timeline

The DISCO-2 project began on the 18th of November, 2020. In 2021, the project welcomed its first students and prioritized fundraising for a successful startup. The design phase was a major focus in 2022 and 2023. The preliminary design review (PDR) took place in May 2023, followed by the critical design review (CDR) in September of the same year. A similar PDR and CDR with focus on the payload and ground station was conducted in January 2024 (PDR) and September 2024 (CDR). In 2023, the DISCO-2 project also expanded with the addition of new students, enhancing the expertise of each project group. One team is responsible for the assembly and test training scheduled

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	Fundraising																																					
	Student groups on submodules																																					
	Payload design and development																																					
vilestone 2	PDR – Preliminary Design Review																																					
	CDR - Critical Design Review																																					
	Satellite operations training																																					
	Payload and Ground station – PDR																																					
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TABLE 3 Orbital parameters used for DRAMA software to estimate orbital lifetime, collision risk and risk assessment during re-entry.

Semi-major axis	Eccentricity	Inclination	Right Ascension of the Ascending Node (RAAN)	Argument of perigee	Mean anomaly
6881	$1 \cdot 10^{-3}$	97.4°	69.7	0	180

for mid-2024, which will then be used to create the CubeSat later that year. The launch of DISCO-2 is scheduled to happen in June 2025, followed by the launch and early operations phase (LEOP). The timeline of DISCO-2 from 2021 to 2026 is depicted in Figure 7 as a simplified Gantt chart where the detailed steps are summarized.

# 7 Analysis, test, and verification

In Denmark it is mandatory to apply for launch permission from the Danish state under the Danish Space Act to be allowed to launch a satellite. This application must include a risk analysis of the mission, focusing on three key aspects: ensuring the satellite will be deorbited within 25 years after the mission ends, preventing the satellite from breaking up in orbit during this period, and ensuring that all parts of the satellite will burn up in the atmosphere during reentry. Additionally, since DISCO-2 will be launched from the US, it must comply with the Federal Communications Commission (FCC) space safety and orbital debris policy. This policy, in addition to the Danish Space Act requirements, mandates "complete disposal as soon as practicable following the end of the mission, and no later than 5 years after the end of the mission." To document that DISCO-2 meets these requirements, we have designed the test and verification program described below.

# 7.1 DRAMA analysis

We use the Debris Risk Assessment and Mitigation Analysis (DRAMA) tool provided by the European Space Agency (Braun et al., 2020) to analyze the orbital lifetime, collision probability, and re-entry risk for DISCO-2. The satellite mass is set to 6 kg based on



the current mass budget with an added safety margin varying between 5 - 10% for the individual components. The initial analysis of the satellite using the Cross section of Complex Bodies Tool (CROC) assuming a "worst case" situation with a random tumbling satellite with deployed solar panels yields an average cross-section of 0.067 m<sup>2</sup>. Applying the orbital parameters specified in Table 3 with the average cross-section gives an estimated lifetime of  $\approx 2.5$  years as shown in Figure 8.



From the following collision analysis it is estimated that the accumulated risk of collisions with an object larger than 1cm is well below the collision threshold probability of  $1 \cdot 10^{-3}$  stated in the recent "ESA Space Debris Mitigation Requirements" essb-st-u 007, 2023 as illustrated in Figure 9. The reentry analysis from DRAMA with a rudimentary model of the satellite further shows no risk for surviving debris during an uncontrolled re-entry of the satellite, with the main components of the structure and solar panels diminishing becoming uncritical at an estimated altitude and between 10 - 80 km.

## 7.2 Pointing analysis

The narrow-angled camera has a FoV of roughly 3° from the center of the image plane. Simulation tests have been conducted, determining whether or not the pointing accuracy of the ADCS is high enough to maintain target within the camera field of view during an overpass. Two tests over two different points of interest (PoIs) have been simulated. Over the first PoI, the CubeSat flies shaded by Earth and over the second PoI, the CubeSat is in illuminated flight. This allows us to test the ADCS with and without the fine sun sensors. The test results can be seen in Figure 10. The pointing error is below 3° around all body axes with uncertainty less than 1° both in illuminated and shaded flights. As the FoV for the high resolution camera is  $57 \times 57$  km. We estimate that we would be able to capture images of the desired target in more than half the images during a single pass. Furthermore, the tests show that the CubeSat requires a longer time to stabilize when the CubeSat is shaded from the sun versus when it is illuminated and can use fine sun sensors. However, this result should be tested in the commissioning and early operation phase. Lastly, we can expect higher precision with further tuning of the MEKF and LQR.

# 7.3 Functional and environmental test

A team of students will assemble and conduct functional components tests of DISCO-2 during assembly. The test scheme is developed at Space Inventor and will serve as a base to develop a protocol for DISCO-2 when performing functional inspections of components between each step of environmental tests. DISCO-2 will be subjected to component testing throughout the assembly to test both the individual components and redundant communication between subsystems. This will also serve as a way to develop the final test procedure when conducting environmental tests on the assembled satellite. The environmental test will consist of a vibrational analysis of the assembled satellite with random vibration on each axis, with a functionality test between each axis-specific vibration. Further DISCO-2 will be subdue to thermal-vacuum cycling test, to ensure the functionality of the CubeSat in a space like environment prior to launch.

## 8 Operations

Because DISCO-2 is a student satellite we can expect that people with different backgrounds would need to be able to operate the satellite. To accomplish this, students are developing an operations and mission control platform, which will aid the collaboration between operators and make the satellite accessible with a low barrier of entry, not requiring a background in programming to be able to operate it. This platform consists of a web interface, that operators will use to set up flight plans and configure the satellite and image processing. A flight plan is a list of commands set to be executed at a specific time. This allows detailed planning of the satellite operations while not in direct contact with the ground station. Using what is then essentially a block programming language, the control panel makes it possible to set up a flight



around z The axis of lowest moment of inertia is the z-axis

plan for the satellite without having to worry about what in a programming term is called boilerplate code–repeated code required to enable other functionality, i.e., the default commands that should always be run (e.g., for collecting the housekeeping data) and properly managing variables and component configurations. It will also be possible to easily configure the DIPP with existing or new modules, that have been compiled for it. An overview of the data flow and communication processes for operations tasks is shown in Figure 11. Flight plans get compiled to the Proc DSL and during the satellite overpass, they can automatically be sent through the ground station to the satellite where it is stored in the payload scheduler running on the M7 core. With this automation, operators will be able to set up unsupervised transmissions e.g., for communicating at night when everyone is asleep, and it makes supervised control more efficient compared to manual control, which is limited by the operator's reaction speed to parse command output and react correspondingly.



# 9 Conclusion

The 3U+ DISCO-2 satellite presented here is one of the most ambitious student CubeSat in the world to date. The EO student CubeSat mission started in 2020 as a collaboration between the three danish universities Aarhus University, Southern Danish University and IT University of Copenhagen. DISCO-2 is planned to launch in 2025 and will be placed into a near-polar orbit with a nominal mission lifetime of  $\approx 2$  years to cover several seasonal changes. DISCO-2 is optimized to reach the scientific objectives of the mission; to contribute to the study of climate change in the arctic. The two optical cameras and the thermal camera will produce high quality data at every passage of the regions of interest in Greenland in combination with novel in-orbit analysis capabilities applying deep learning to optimise the data quality for downlinked observations. Throughout the whole mission from idea-phase to the current state, student involvement have been highly prioritized to provide a unique opportunity for students to gain insight in various aspects of space science, EO and CubeSats. DISCO-2 serves also as a showcase and prototype mission to gain knowledge on the capabilities of such a small and cost-effective satellite to investigate how data can be used for the scientific research areas in question. The first part of the mission has now successfully passed and the satellite will soon be ready for launch. Upon launch, concentrated efforts on mastering the control and communication with the satellite will be essential to fully realize and utilize the potential of this small satellite.

# Data availability statement

The original contributions presented in the study are included in the article/supplementary material, further inquiries can be directed to the corresponding author.

# Author contributions

AD: Writing-original draft, Writing-review and editing. MFA: Writing-original draft, Writing-review and editing. JP:

editing. Writing-original draft, Writing-review NE: and Writing-original Writing-review editing. draft, and MF: Writing-original draft, Writing-review and editing. CM: Writing-original draft, Writing-review and editing. TR: draft. Writing-review editing. Writing-original and NN: Writing-original draft, Writing-review editing. and CS: Writing-original draft, Writing-review and editing. MJA: Writing-original draft, Writing-review and editing. SiS: Writing-review and editing. TL: Writing-review and editing. JD: Writing-original draft, Writing-review and editing. LP: Writing-original draft, Writing-review and editing. RI: Writing-review and editing. SR: Writing-review and editing. JK: Writing-review and editing. RB: Writing-review and editing. CC: Writing-review and editing. EC: Writing-review and editing. IG-L: Writing-review and editing. RG: Writing-review and editing. BH: Writing-review and editing. JH: Writing-review and editing. JJ: Writing-review and editing. DK: Writing-review and editing. ML: Writing-review and editing. JL: Writing-review and editing. MM: Writing-review and editing. OM: Writing-review and editing. TM: Writing-review and editing. SN: Writing-review and editing. AN: Writing-review and editing. GN: Writing-review and editing. MP: Writing-review and editing. AP: Writing-review and editing. PR: Writing-review and editing. SoS: Writing-review and editing. ÍS: Writing-review and editing. GS: Writing-review and editing. NS: Writing-review and editing. SaS: Writing-review and editing. AT: Writing-review and editing. JT: Writing-review and editing. NV: Writing-review and editing. AV: Writing-review and editing. CK: Writing-original draft, Writing-review and editing.

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# Conflict of interest

The authors declare that the research was conducted in the absence of any commercial or financial relationships that could be construed as a potential conflict of interest.

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