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We have revisited recent results on the ideal ignition of  $H<sup>11</sup>B$  fuel, in the light of the latest available reactivity, an alternative self-consistent calculation of the electron temperature, an increased extent of the suprathermal effects and the impact of plasma density. At high density, we find that the ideal ignition temperature is appreciably relaxed (e.g.,  $T_i \approx 150 \,\text{keV}$  for  $n_i \sim 10^{26} \,\text{cm}^{-3}$  and an optimal <sup>11</sup>B/H concentration  $\varepsilon = 0.15$ ) and burn becomes substantial. We have then investigated central hot-spot ignition in both isobaric and isochoric inertial confinement configurations. Although implosion-driven ignition appears to be unfeasible, the isochoric self-heating conditions foster favourable preliminary conclusions on the utilization of proton fast ignition. In the isochoric case, we find a broad minimum in the ignition energy at  $\rho R \approx 8.5$  g/cm<sup>2</sup> and 220 ≤  $T_i \le 340$  keV  $(80 \leq T_e \leq 95 \text{ keV})$ , for  $\varepsilon = 0.15$ .

#### KEYWORDS

proton-boron fusion, inertial confinement fusion, hot-spot ignition, proton fast ignition, laser boron fusion, aneutronic fusion

# 1 Introduction

The  $^{11}B(p,3\alpha)$  fusion reaction, with a Q-value of 8.6 MeV, is experiencing a renewed interest for energy production purposes, in the light of recent experimental and theoretical findings [[1](#page-4-0)–[11](#page-4-1)]. The reaction is aneutronic and involves only abundant and stable isotopes. Moreover, the α particles in its final state may release all their energy to the fusion plasma. The reaction is also of interest for studies in stellar evolution, where relative abundances of <sup>11</sup>B, Li and Be provide insight into stellar processes [\[12\]](#page-4-2). Proposed approaches for energy production span magnetic [\[13\]](#page-4-3), magneto-inertial [[14,](#page-4-4) [15\]](#page-5-0) and laser-driven [[5](#page-4-5), [16,](#page-5-1) [17\]](#page-5-2) fusion. The exploitation of  $H^{11}B$  fuel, however, remains extremely challenging because of its low reactivity and high radiative losses at temperatures attainable in present-day fusion devices.

The existence of ideal ignition conditions has been demonstrated only lately by Putvinski et al. [[6\]](#page-4-6), who have used a recent fusion cross section dataset [[4](#page-4-7)] for the calculation of the thermal reactivity and added to this latter a contribution coming from kinetic (particularly, suprathermal) effects, calculated self-consistently. Suprathermal effects are due to elastic collisions between the fusion-born α′s and background thermal protons [\[6](#page-4-6), [18\]](#page-5-3), which develop a bolder tail in the proton energy spectrum compared to the Maxwell-Boltzmann distribution [\[19](#page-5-4)–[21](#page-5-5)]. Putvinski et al. [\[6\]](#page-4-6) have found fusion power to overcome bremsstrahlung losses only marginally, for 250 ≤  $T_i$  ≤ 380 keV, in a dilute plasma ( $n_i = n_p + n_B = 10^{14}$  cm<sup>-3</sup>) at the optimal <sup>11</sup>B/H ion concentration  $\varepsilon \equiv n_B/n_p = 0.15$ , with  $T_e$  calculated self-consistently (standard notation is used).

In this Brief Research Report, we first revisit those findings in the light of the latest available reactivity, an alternative self-consistent calculation of  $T_e$  as well as the actual extent of the suprathermal effects. We then show how ideal ignition conditions vary depending on the plasma density regime, the extent of suprathermal effects and the boron-to-hydrogen concentration. We find a relaxed ignition temperature and a significantly larger fusion-to-



<span id="page-1-0"></span>bremsstrahlung power ratio at high density. Consequently, we study ignition in actual isobaric and isochoric hot-spot fuel configurations, and draw preliminary conclusions on fast ignition.

We recall that hot-spot fuel configurations are relevant to laserdriven inertial confinement, which is a promising method to achieve fusion energy [[22](#page-5-6)]. Ignition of DT fuel has recently been achieved at the US National Ignition Facility [[23](#page-5-7)], by exploiting an indirectdrive scheme based on a nearly isobaric fuel configuration [\[24\]](#page-5-8). Fast ignition is a technique alternative to hot-spot ignition and is based on the ignition of precompressed fuel by means of an external trigger. Laser-driven fast ignition was proposed by Tabak et al. [[25](#page-5-9)] 30 years ago and it is today the subject of significant theoretical and experimental investigation [\[26,](#page-5-10) [27](#page-5-11)].

# 2 Ideal ignition

<span id="page-1-3"></span>The ideal ignition conditions of Putvinski et al. [[6\]](#page-4-6) have been recalculated and plotted in [Figure 1A](#page-1-0) (blue curve) in terms of the ratio  $P_f/P_b$ , where

$$
P_f(T_i) = (1+s)P_{th}(T_i) \tag{1}
$$

is the fusion power (per unit volume),  $P_{th}$  is the thermal fusion power, s is a parameter expressing the suprathermal contribution, and  $P_b(T_e)$ is the bremsstrahlung power (see [Appendix A](#page-6-0) for formulas). While  $s = 0.1$  has been found by Putvinski *et al.*, large-angle scattering,

particularly by the effect of the nuclear (strong) interaction, does not appear to have been taken into account in their α-p collision calculations. One of us has shown  $[20]$  $[20]$  that in a H<sup>11</sup>B plasma at high density ( $n_e \sim 10^{26}$  cm<sup>-3</sup>) and electron temperature ( $T_e \ge 50$  keV), suprathermal effects calculated on the basis of the complete elastic α-p cross section can be approximately two times higher than those found upon the assumption of a pure Coulomb scattering. Accordingly, we put forward that s is very likely to reach 0.2. This assumption is in line with the earliest findings of Weaver et al. [[19](#page-5-4)], who calculated a suprathermal increase of the  $H^{11}B$  reaction rate up to 15% at high plasma density and temperature, based on kinetic simulations that included both Coulomb and nuclear large-angle scattering (but that were biased by the poor knowledge of the relevant elastic and fusion cross sections at that time).

From [Figure 1A](#page-1-0), we note that ignition is not possible if the suprathermal contribution is not accounted for  $(s = 0, \text{black curve})$ . We also note that the region where  $P_f/P_b \geq 1$  extends to higher values of  $T_i$  (up to 440 keV) compared to Ref. [\[6](#page-4-6)], which is very likely due to the fact that we have used a more accurate (and appreciably higher) reactivity [[28](#page-5-13)] in the calculation of  $P_{th}$ .

<span id="page-1-2"></span><span id="page-1-1"></span>As for the self-consistent calculation of  $T_e$ , we recall that in the  $T_i - T_e$  plane of an ideal plasma, the self-burn region is bounded by the solutions of the steady-state power balance equations [\(29\)](#page-5-14)

$$
P_f - P_b = 0 \tag{2}
$$

$$
\eta_i P_f - P_{ie} = 0 \tag{3}
$$

where  $\eta_i$  is the fraction of the fusion power transferred to the ions by the  $\alpha$  particles and  $P_{ie}$  is the power transferred from the ions to the electrons (see [Appendix A](#page-6-0)). Eq. [2](#page-1-1) is the balance equation for the entire system and gives the maximum possible  $T_e$ ,  $T_{e, max}$ , at any  $T_i$  higher than the ideal ignition temperature,  $T_{id}$ , while Eq. [3](#page-1-2) holds for the ion fluid only and gives the minimum possible  $T_e$ ,  $T_{e, min}$  [\(Figure 1B](#page-1-0)). All the possible trajectories of the system during burn, which are determined by the time-dependent power flow equations (and their initial conditions) for the ion and electron fluids, lie between the  $T_{e,min}$  and  $T_{e,max}$  curves. We have used Eq. [3](#page-1-2) to obtain the  $T_e(T_i)$ relationship (blue curve in [Figure 1B](#page-1-0)), whereas Putvinski et al. have used the power balance equation for the electron fluid, i.e.,

$$
(1 - \eta_i)P_f + P_{ie} - P_b = 0 \tag{4}
$$

<span id="page-2-1"></span>which yields slightly higher values of  $T_e$ . Those result in a lower  $P_f/P_b$  ratio compared to [Figure 1A](#page-1-0) as  $P_b$  increases with  $T_e$ .

While ignition and self-burn appear less marginal than previously found, low-density plasmas remain of a primary interest for magnetic confinement approaches, which can operate at sub-ignition. More meaningful conclusions can be drawn for ignition-based schemes, at high density. While the explicit square-density dependence of the P-terms cancels out in Eqs [2](#page-1-1), [3](#page-1-2), a residual dependence on density remains in Eq. [3](#page-1-2) through the Coulomb logarithms of  $P_{ie}$  −cp. Eqs [A3,](#page-6-1) [A4.](#page-6-2) Typical values of  $ln \Lambda_{ie}$  at  $T_e = 100 \text{ keV}$  are:  $ln \Lambda_{pe} \approx 24$ ,  $ln \Lambda_{Be} \simeq 22$  for  $n_e = 10^{14}$  cm<sup>-3</sup>, and  $ln \Lambda_{pe} \simeq 8$ ,  $ln \Lambda_{Be} \simeq 6$  for  $n_e = 10^{26}$  cm<sup>-3</sup>. The change of the Coulomb logarithms upon density causes the  $T_{e,min}$  curve in [Figure 1B](#page-1-0) to shift downward while moving from a dilute to a dense plasma. As a consequence,  $T_{id}$  decreases and the self-burn region enlarges. At high density, ignition and burn are quite substantial for  $H^{11}B$  fuel. Note that this effect is amplified by the strong decoupling between  $T_e$  and  $T_i$ , without which  $P_{ie} \approx 0$ .

As  $P_b$  is minimum, throughout the burn region, along  $T_e = T_{e,min}(T_i)$ , this latter condition also yields the maximum  $P_f/P_b$  ratio attainable at a given  $T_i$ . This has been plotted in [Figure 1C](#page-1-0) for  $n_i = 10^{26}$  cm<sup>-3</sup> (corresponding to a mass density  $\rho \approx 250 \text{ g/cm}^3$  for  $\varepsilon = 0.15$ ) and representative values of s. For s = 0.2,  $T_{id}$  lowers to about 150 keV, while  $P_f/P_b$  overcomes 2. Moving to  $n_i = 10^{27} \text{cm}^{-3}$  does not change  $T_{e,min}(T_i)$  substantially and decreases  $T_{id}$  only by a few keV. On another note,  $T_{e,min}$  and  $T_{e, max}$  are mildly sensitive to s, for s  $\ll 1$ . For instance, it is easy to see that  $T_{e, max}$  scales approximately as  $(1 + s)^2$ .

The high  $P_f/P_b$  ratio of [Figure 1C](#page-1-0) encourages the analysis of ignition conditions in hot-spot configurations, where additional loss terms come into play, and shows the potential to withstand fuel depletion and bremsstrahlung emission due to the α-particle ash [\[30,](#page-5-15) [31](#page-5-16)]. It also opens the possibility of working at increased  $^{11}B$ concentration. [Figure 1D](#page-1-0) shows  $P_f/P_b$  curves for the case of  $\varepsilon$  = 0.5. Ideal ignition can still be achieved, however subject to the suprathermal contribution and at the expenses of higher values of  $T_{id}$ . Ignition at equimolarity, i.e.,  $\varepsilon = 1$ , is confirmed to be impossible, at least for  $s\,{\leq}\,0.2.$ 

## 3 Hot-spot ignition

<span id="page-2-0"></span>The power balance condition for a hot spot of radius R and density  $\rho$  at the ignition threshold reads

$$
P_f - P_b - P_h - P_m = 0 \tag{5}
$$

where  $P_h$  is the power density lost through heat conduction and  $P_m$  that lost through mechanical work ( $P_m = 0$  in the isobaric case); see [Appendix A](#page-6-0). All fusion born  $\alpha$ 's are assumed to remain inside the hot spot. Eq. [5](#page-2-0) results in a quadratic equation for  $\rho R$ which coefficients, in general, are functions of  $T_i$  and  $T_e$ , and depend on  $\varepsilon$  and  $s$ . Eq. [5](#page-2-0) is coupled to either of the power balance equations for ions and electrons −analogue to Eqs. [3](#page-1-2), [4](#page-2-1) for the ideal case− through which the variable  $T_e$  can be eliminated. It is convenient to work with the power flow equation for the ion fluid, i.e.,

$$
\eta_i P_f - P_{ie} - P_{m,i} = 0 \tag{6}
$$

<span id="page-2-2"></span>where  $P_{m,i}$  is the component of  $P_m$  exerted by the ions. In the isobaric case, the fact that  $P_{m,i} = 0$  enables the use of the same characteristic curve given by Eq. [3](#page-1-2), in blue in [Figure 1B](#page-1-0). This selfconsistent relationship can be retained also in the isochoric case, inasmuch as  $P_{m,i} \ll P_{ie}$ . The isochoric ignition boundary has been generated upon this assumption [\(Figure 2\)](#page-3-0). As a term of reference, the contour corresponding to  $P_{m,i}/P_{ie} = 0.3$  has also been plotted, which shows that the condition  $P_{mi} \ll P_{ie}$  is reasonably consistent with the large  $\rho R$  values entailed by the isochoric curve, due to the  $(\rho R)^{-1}$  dependency of  $P_{m,i}$ .

By analogy with the DT and DD cases, we expect that 1D simulations of pre-assembled fuel would actually show a lower branch of the ignition curves in the proximity of and after their knee, due to a cooling/re-heating mechanism of the hot spot for initial points located just below the analytic curves [[32](#page-5-17)–[35](#page-5-18)]. In the case of isochoric DT, for instance,  $\rho R$  reduces by a factor of 1.5 when  $T_i$  is twice the minimum of the analytic curve, and the gap increases with  $T_i$  [\[35\]](#page-5-18). The difference is even more dramatic in the case of isobaric DT [\[33](#page-5-19)–[35](#page-5-18)].

Although the confinement parameter is high at the minimum of ignition curves, we have checked that the plasma is still optically thin, i.e., the Planck mean free path,  $l_p$ , is much larger than R. This is due to the high electron temperature. Indeed, one has  $l_P = (\rho \kappa_p)^{-1}$ , where

$$
\kappa_p = 0.43 \left( \langle Z \rangle \langle Z^2 \rangle / \langle A \rangle^2 \right) \rho T_e^{-7/2} \text{ cm}^2 / \text{g} \tag{7}
$$

<span id="page-2-5"></span>is the free-free Planck mean opacity [\[35\]](#page-5-18), with

$$
\langle X \rangle = \frac{n_p X_p + n_B X_B}{n_i} = \frac{X_p + \varepsilon X_B}{1 + \varepsilon} \tag{8}
$$

For instance,  $\rho l_P \approx 9000 \text{ g/cm}^2$  at  $T_i = 200 \text{ keV}$  ( $T_e = 80 \text{ keV}$ ),  $\rho =$  $1000 \text{ g/cm}^3$  and  $\varepsilon = 0.15$ .

The ignition energy,  $E_{ig}$ , has been calculated as the internal (thermal) energy of the hot spot in the ideal gas approximation, i.e.,

$$
E_{ig} = \frac{3}{2} pV = 2\pi ( \Gamma_i T_i + \Gamma_e T_e) \frac{(\rho R)^3}{\rho^2}
$$
 (9)

<span id="page-2-3"></span>where  $V$  is the volume of the plasma sphere and  $p$  is its pressure, as given by

$$
p = p_i + p_e \tag{10}
$$

$$
p_{i(e)} = \Gamma_{i(e)} \rho T_{i(e)} \tag{11}
$$

$$
\Gamma_i = k \big/ \langle A \rangle m_p, \Gamma_e = \langle Z \rangle \Gamma_i \tag{12}
$$

<span id="page-2-4"></span>The quantity  $\rho^2 E_{ia}$  has been plotted vs.  $T_i$  in [Figure 2](#page-3-0) only for the isochoric case, which is expected to enable the higher gain. The curve



<span id="page-3-0"></span>shows a broad minimum for  $220 \le T_i \le 340$  keV, which corresponds to  $\rho R \approx 8.5$  g/cm<sup>2</sup>. One can estimate that at  $\rho$  as high as 4,000 g/cm<sup>3</sup>,  $E_{ia}$  at its minimum is still considerably large (~3 MJ).

# 4 Discussion and conclusion

Despite the fact that self-heating is possible in a pre-formed hot spot, we have verified that implosion-driven formation of the hot spot is hydrodynamically impossible, on the basis of the same argument preventing it in pure D fuel [\[34\]](#page-5-20), i.e., a cooling timescale shorter than the hot-spot confinement time  $t_c \sim R/c_s$ , being  $c_s$  the isothermal sound velocity. Even without this issue, considerations on the required implosion velocity and hydrodynamic instabilities would prevent this scheme from being viable. These circumstances point toward fast ignition as possibly the only scheme to ignite inertially confined  $H^{11}B$ fuel, apart from the trivial, low-gain case of volume ignition. Nevertheless, isochoric self-heating conditions provide a preliminary estimate of fast ignition requirements [[35](#page-5-18)].

Proton fast ignition [[36](#page-5-21)] is particularly suited to  $H^{11}B$  fuel, not only because of its superior ballistic properties in the energy deposition and the potential capability of inducing the hot-ion mode, but also because of the additional heating provided by the in-flight fusion reactions of the proton beam [\[21,](#page-5-5) [37](#page-5-22)]. It has been put forward [\[21\]](#page-5-5) that in a fully degenerate <sup>11</sup>B plasma, under certain conditions, this contribution could become as large as the initial kinetic energy of the proton beam. Such an effect could then appreciably reduce the ignitor energy required in a H<sup>11</sup>B mixture. Taking also into account the reduction of  $E_{ia}$  because of the abovementioned lowering of the higher- $T_i$  branch of the isochoric curve –note that  $E_{ig}$  scales as  $(\rho R)^3$  – ignitor energies of a few hundreds kJ can be expected at densities around  $4000 \text{ g/cm}^3$ . The laser energy required to drive the implosion is estimated at about 1.3 MJ per mg of fuel, by assuming an overall laser-target coupling efficiency of 15% (direct drive), a unit is entrope parameter, and  $\varepsilon = 0.15$ .

The ignitor pulse will have to be delivered to the compressed target within a timescale shorter than  $t_c$ ; e.g.,  $t_c \approx 6$  ps for  $T_i =$ 220 keV ( $T_e = 85 \text{ keV}$ ),  $\rho R = 10 \text{ g/cm}^2$  and  $\rho = 4,000 \text{ g/cm}^3$ . Due to the progressive heating induced, the protons in the bunch will experience rapidly and drastically changing plasma conditions upon their arrival onto the hot spot, and even during their slowing down [[38](#page-5-23)]. Plasma degeneracy will shift from strong at the onset of the ignitor pulse to very weak on its tail. Matching the proton range to the hot spot confinement parameter along the evolving plasma conditions will require a suitably tailored proton spectrum. The determination of such spectrum can only be carried out upon a self-consistent approach to the ignitor-fuel interaction, through accurate simulations. While this task is beyond the scope of the present study, here we estimate, for instance, that 2.5 MeV protons have a 10  $g/cm<sup>2</sup>$  range at the ignition conditions which have been used to calculate  $t_c$  and which correspond to the last stage of the hot-spot heating process (see [Appendix A](#page-6-0) for details on the stopping power). At the early stage of hot-spot heating, for e.g.,  $T_i = T_e = 10 \text{ keV}$ , the 10 g/cm<sup>2</sup> range corresponds to a much higher proton energy, about 200 MeV. In practice, protons with a mean energy of a few tens MeV will most probably be needed. Assuming a 300 kJ ignitor, an overall number of protons of the order of  $10^{16}$ – $10^{17}$  is estimated accordingly. As a term of comparison, a highly directional beam of  $10^{13}$  protons with an approximately Maxwellian spectrum at an effective temperature of 6 MeV has been produced under intense laser irradiation (600 J, 0.5 ps,  $3 \times 10^{20}$  W/cm<sup>2</sup>) of thin foils, through the Target Normal Sheath Acceleration (TNSA) mechanism [\[39\]](#page-5-24). Driving a TNSA-based proton ignitor for H<sup>11</sup>B fuel will therefore require a multiple-beam laser firing scheme and a suitably engineered foil target (e.g., multi-spot designed, heavily H-loaded, convex-shaped for focusing the ignitor beam). The placement of such an extended foil target sufficiently close to the hot spot to limit time-of-flight dispersion of the ignitor power will require cone-guiding through the fuel capsule [\[40](#page-5-25)], with a wide cone aperture. A conically guided capsule will also limit the implosion driver energy while largely preserving the gain [\[40\]](#page-5-25).

With a 300 kJ ignitor and the highest reported laser-to-proton energy conversion efficiency, 15% [\[41](#page-5-26), [42\]](#page-5-27), an overall laser energy of 2 MJ will be needed to drive the ignitor. This energy will have to be delivered to the foil target over a timescale of 1 ps. Suitable laser amplifiers and laser architectures will have to be developed to this extent as well as for the ns-scale implosion of the fuel, where driver energies above 10 MJ are expected. Both Diode-Pumped Solid-State Laser (DPSSL) and excimer laser systems show the potential to be scaled up to the large energy outputs required for compression and fast ignition of  $H^{11}B$  fuel, on both the ns and ps timescales [\[43,](#page-5-28) [44\]](#page-5-29).

We finally recall that within the frame of a very specific fast ignition scheme, based on a laser-driven relativistic shock wave, Eliezer et al. [\[45](#page-5-30)] have found that a laser pulse with intensity of  $1.6 \times 10^{25}$  W/cm<sup>2</sup>, duration of 1 ps and energy of 21 MJ impinging on fuel pre-compressed at  $4,800$  g/cm<sup>3</sup> can generate a side, cylindrical hot spot with a depth of 8.3 g/cm<sup>2</sup>,  $T_i \approx 200 \text{ keV}$ ,  $T_e \approx 50 \text{ keV}$ , where ignition is achieved. Such a laser pulse is judged impracticable in the near term.

On the contrary, our preliminary analysis shows that proton fast ignition of isochoric  $H<sup>11</sup>B$  fuel requires compression and ignitor performances which, though challenging, are in line with near-future laser capabilities. We plan to devote further work to demonstrate burn propagation, better quantify ignition parameters and calculate gain in such scheme, considering actual target configurations.

## Data availability statement

The original contributions presented in the study are included in the article/Supplementary Material, further inquiries can be directed to the corresponding author.

# Author contributions

EG: Formal Analysis, Writing–original draft, Writing–review and editing, Data curation, Investigation, Software, Visualization. FB: Formal Analysis, Writing–original draft, Writing–review and editing, Conceptualization, Methodology, Supervision, Validation.

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# Appendix A: Formalism

#### Power density terms

Explicit expressions for the power density terms in Eqs [1](#page-1-3)–[6](#page-2-2) are given hereafter (electrostatic cgs units are used):

$$
P_{th} = n_p n_B \langle \sigma v \rangle Q \tag{A1}
$$

where  $\langle \sigma v \rangle$  is the Maxwellian reactivity [\[28\]](#page-5-13);

$$
P_b = 8.511 W_0 \hat{T}_e^{-1/2} \left\{ Z_{eff} \left( 1 + 1.78 \hat{T}_e^{-1.34} \right) + 2.12 \hat{T}_e \left( 1 + 1.1 \hat{T}_e - 1.25 \hat{T}_e^{-2} \right) \right\}
$$
(A2)

<span id="page-6-1"></span>where  $\hat{T}_e = kT_e/m_ec^2$  and  $W_0 = e^6n_e^2/m_ec^2\hbar$  $W_0 = e^6n_e^2/m_ec^2\hbar$  $W_0 = e^6n_e^2/m_ec^2\hbar$  [6, [46\]](#page-5-31);

$$
P_{ie} = \frac{3}{2} (n_p \nu_{pe} + n_B \nu_{Be}) k (T_i - T_e)
$$
 (A3)

<span id="page-6-2"></span>where  $\nu_{ie} = \nu_{ie}^* \Delta$ ,

$$
\nu_{ie}^* = \frac{(4\pi)^2 n_e e^4 Z_i^2 \ln \Lambda_{ie}}{3\sqrt{2}\pi^{3/2} m_i m_e} \left(\frac{m_e}{kT_e}\right)^{3/2} \tag{A4}
$$

is the classical heat exchange rate [[29\]](#page-5-14),

$$
\Delta = \frac{\left(1 + 2\widehat{T}_e + 2\widehat{T}_e^2\right)\sqrt{\pi \widehat{T}_e^3/2}}{\int_0^\infty \exp\left\{-\left(\sqrt{1 + x^2} - 1\right)\right/\widehat{T}_e\right\}x^2dx}
$$
(A5)

is its relativistic correction [[6](#page-4-6)], and

$$
ln \Lambda_{ie} = \frac{3}{2Z_i e^3} \sqrt{\frac{4.2 \times 10^5 k}{\pi n_e}} kT_e
$$
 (A6)

according to Spitzer [[47](#page-5-32)];

$$
P_h = -\epsilon \left( Z_i \right) \kappa \nabla T_e S / V \tag{A7}
$$

where  $-\nabla T_e \approx T_e/R$ , S is the area of the hot-spot surface  $(S/V = 3/R)$ ,

$$
\kappa = 20 \left(\frac{2}{\pi}\right)^{3/2} \frac{(kT)^{5/2}k}{m_e^{1/2}e^4 Z_i \ln \Lambda_{ie}} \delta T(Z_i)
$$
 (A8)

is the Spitzer thermal conductivity, with  $Z_i \rightarrow Z_{eff}$  for a multi-Z plasma, and the correction factors  $\epsilon$  and  $\delta T$  have been calculated through interpolation ( $\epsilon = 0.407$ ,  $\delta T = 0.426$ ) [\[47\]](#page-5-32);

$$
P_m = P_{m,i} + P_{m,e} \tag{A9}
$$

$$
P_{m,i(e)} = 3 \frac{p_{i(e)}}{R} u
$$
 (A10)



<span id="page-6-3"></span>where  $u \approx \sqrt{3p/4\rho}$  is the velocity of the material behind a strong shock in isochoric fuel [\[35](#page-5-18)] and the pressures  $p$ ,  $p_i$ ,  $p_e$  are given by Eqs [10](#page-2-3)–[12.](#page-2-4)

#### Fusion energy partition and stopping power

<span id="page-6-4"></span>The fusion power fraction to ions,  $\eta_i$ , has been calculated according to Levush and Cuperman [[48](#page-5-33)] ([Figure A1](#page-6-3)). A stopping power of the form

$$
dE/dt = E/t_E + \gamma / \sqrt{E}
$$
 (A11)

has been used for the fusion-born  $\alpha$  particles, with  $t_E$  and  $\gamma$  given by Corman et al. [\[49\]](#page-5-34). A simplified α spectrum has been utilised, as explained in Ref. [\[20](#page-5-12)].  $\eta_i$  approaches or even overcomes 90% at values of  $T_i$  and  $T_e$  of interest for ignition and burn. The same stopping power model of Eq. [A11](#page-6-4) has been utilised for energy-range calculations for the ignitor protons.

#### <span id="page-6-0"></span>Useful relations

The densities  $n_p$  and  $n_B$  are linked to  $n_e$  and  $\rho$  by the obvious relations

$$
n_p = n_e / (Z_p + \varepsilon Z_B)
$$
 (A12)

$$
n_B = \varepsilon n_e / (Z_p + \varepsilon Z_B) \tag{A13}
$$

$$
n_e = \langle Z \rangle n_i \tag{A14}
$$

$$
\rho = n_i \langle A \rangle m_p \tag{A15}
$$

with  $\langle Z \rangle$  and  $\langle A \rangle$  given by Eq. [8.](#page-2-5)