



Rock Magnetism of Late Cretaceous to Middle Eocene Strata in the Lesser Himalaya, Western Nepal: Inferences Regarding the Paleoenvironment

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Khatri DB, Zhang W, Fang X, Meng Q, Zhang T, Zhang D and Paudayal KN (2021) Rock Magnetism of Late Cretaceous to Middle Eocene Strata in the Lesser Himalaya, Western Nepal: Inferences Regarding the Paleoenvironment. Front. Earth Sci. 9:744063. doi: 10.3389/feart.2021.744063 The growth of the southern piedmont of the Himalayan boundary and its depositional setting has changed since uplift of the Himalaya due to continental Indian-Eurasian collision, which has resulted in variation in magnetic minerals in marine- and terrestrialfacies sediments. In this paper, we utilize rock magnetism data from the late Cretaceous to middle Eocene strata, including the Amile and Bhainskati formations from the Lesser Himalaya (western Nepal), to understand the mechanism controlling magnetic susceptibility (χ). The active tectonics strongly influenced saturation isothermal remanent magnetization (SIRM), HIRM, and hysteresis loops, forming both lowcoercivity minerals in sediments with low χ from the terrestrial facies (zones I, IIIA, and V) and high-coercivity minerals in the sediments with high χ from the marine facies (zones II, IIIB and IV). Thermomagnetic κ -T curves and frequency-dependent χ (χ fd%) values show that sediments with low χ and high χ carry magnetite with coarse non-superparamagnetic (SP) grains and hematite with SP grains, respectively. Comparing the χ data with the lithologic, sedimentary environments, geomorphic features, and sea level data, we propose that low χ values were mainly produced by an increase in terrigenous detrital influx during the regression period of the Tethys Sea, while high χ values formed in marine sediments, which prompted the appearance of ferromagnetic-antiferromagnetic and paramagnetic minerals during the transgression of the Tethys Sea.

Keywords: rock magnetism, Cretaceous to Eocene strata, paleoenvironment, transgression and regression, Lesser Himalaya

INTRODUCTION

The Himalayan orogeny and tectonic stress produced the existing Himalayan arc as a result of the Indian-Eurasian collision since the late Cretaceous time (Beck et al., 1995). This has attracted the attention of various scientists to study its effect on geomorphological features (Molnar and Tapponnier, 1975), provenance (Garzanti, 1999; DeCelles et al., 2004; DeCelles et al., 2014), climate (Ghosh et al., 1995), environment (Bosboom et al., 2014) and ecology (Smith et al., 2015). In Nepal, some paleomagnetic studies have been carried out in the Siwalik (Tokuoka

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et al., 1986; Gautam and Appel, 1994; Gautam et al., 1995; Gautam and Pant, 1996; Gautam and Fujiwara, 2000; Gautam, 2008; Ojha et al., 2009), Lesser, Higher and Tethyan Himalaya (Yoshida and Sakai, 1984; Appel et al., 1991; Pant et al., 1992; Crouzet et al., 2001; Gautam et al., 2011). The Tansen area, which is located at the front of the Lesser Himalaya, incorporates both marine and terrestrial sediments since the continental collision and uplift of the Himalaya. It acts as a remarkable area for reconstructing the tectonic and climatic evolution before and after the collision. Various studies related to detrital geochronology and geochemistry, lithological transitions, flora and faunal fossils have been reported from this region (Sakai, 1983; Sakai, 1984; Kimura et al., 1985; Matsumaru and Sakai, 1989; Sah and Schleich, 1990; DeCelles et al., 2004), showing diverse opinions on the paleoenvironmental condition. Generally, four distinctive views can be perceived on the paleoenvironment changes as 1) the humid climate prevailed during Cretaceous to Eocene (Neupane and Zhao, 2018; Neupane et al., 2021), 2) there was a paleoenvironmental transition from the pteridophytes dominated wet and humid environment during late Cretaceous (Mohabey et al., 1993; Srivastava, 2011; Prasad et al., 2018) to more humid environment in the entire Eocene revealed by evergreen broad-leaved forest forming coal and carbonaceous shale (Mehrotra, 2003; Srivastava, 2011; Shukla et al., 2014; Spicer et al., 2014; Samant et al., 2020), 3) the paleoclimate during Cretaceous to Eocene had been affected by the Tethys Sea and the Himalayan tectonic uplift rather than the global climate change (Bosboom et al., 2011; Licht et al., 2013; Licht et al., 2014), and 4) the Inter-tropical Convergence Zone (ITCZ) had driven the Eocene monsoon climate (Boos and Kuang, 2010; Spicer et al., 2016). The poor preservation of late Cretaceous marine and early Cenozoic terrestrial sequences in South Asia has been the main factor causing these debates. The Palpa section in this area consists of a late Cretaceous-Paleocene (so called Amile Formation) to middle Eocene (called as Bhainskati Formation) sequence of both marine and terrestrial sediments. Their respective ages were constrained by U-Pb detrital zircon ages, Nd isotope analyses and trace element compositions (Robinson et al., 2001; Neupane et al., 2017; Bhandari et al., 2019) and fossil records (Sakai, 1983; Matsumaru and Sakai, 1989). The rock magnetic study has been rarely reported from this area (Gautam, 1989a, b) and is expected to effectively diagnose the paleoenvironmental changes in the Lesser Himalaya, western Nepal.

The uplift of the Himalaya-Tibetan Plateau after the Indian-Eurasian collision has affected the provenance and environmental settings in this region. In addition, the transgression and regression of the Tethys sea within this region also caused the change in depositional settings (either terrestrial or marine setting). The depositional processes have been widely investigated through the rock magnetic properties of sedimentary sequences in the recent years (Verosub and Roberts, 1995; Liu et al., 2012). The magnetic parameters have been also applied for obtaining better discrimination of lithological units such as marine limestone or mudstone and continental clastic rock (Murdock et al., 2013). The Palpa section, on the southern piedmont of the Himalaya records abundant environmental and tectonic information. It provides an excellent opportunity to explore the paleoenvironment through the study of rock magnetic analysis. In this paper, we examine detailed rock magnetic analyses from the Palpa section in the Tansen area to identify the depositional setting that most likely controlled the magnetic susceptibility (χ) variation and make an initial attempt to gain insights into the relationship between the depositional environment and the tectonic uplift during and after the Indian-Eurasian collision.

GEOLOGICAL AND GEOGRAPHICAL SETTING

The Indian-Eurasian collision results in the development of four different tectonostratigraphic units (Figure 1A). The northern Tethys Himalayan Sequence (Neoproterozoic through Cenozoic) is mainly composed of sedimentary rocks with numerous fossiliferous horizons (Stöcklin, 1980; Garzanti and Frette, 1991; Yin and Harrison, 2000; Guillot et al., 2008; Najman et al., 2017). In its south, the high-grade metasedimentary rocks of the Higher or Greater Himalayan Sequence (Neoproterozoic through Cambrian) is thrust southward along the Main Central Thrust (MCT) on the uppermost part of the Lesser Himalayan Sequence (Paleoproterozoic to Paleozoic) (Parrish and Hodges, 1996; DeCelles et al., 2000; Martin, 2017). Furthermore, the weakly metamorphosed Indian continental crust and sedimentary rocks of the Lesser Himalayan Sequence (Valdiya, 1980; Sakai, 1985; Parrish and Hodges, 1996; Upreti, 1999; DeCelles et al., 2000; Martin et al., 2011) are separated by the Main Boundary Thrust (MBT) in the south. The southernmost part consist of marginal deposits eroded from the higher mountain belts, deposited in the foreland basin as the Siwalik (or Churia zone) or Sub-Himalayan Sequence (Paleogene and Neogene), and now lie above the Main Frontal Thrust (MFT) with recent alluvial deposits of the Indo-Gangetic plain (Gansser, 1964; Stöcklin, 1980; DeCelles et al., 1998; White et al., 2001; Najman et al., 2005; van der Beek et al., 2006; Yin, 2006; Ojha et al., 2009). In Nepal, the late Permian to Eocene sediments (called the Gondwana System) typically contains both marine facies, including fossiliferous argillaceous limestones and siltstone/sandstone with molluscs, foraminifers, coral and vertebrate fossils, and terrestrial facies, including ferruginous quartz arenite, conglomeratic quartz arenite, glacial diamictite, interbedded sandstone, and shale. These lithostratigraphic units are sparsely distributed in the Barahachhetra and Katari areas, Tansen, Tulsipur, and the Birendranagar areas extending from east to west, indicating the existence of scattered depositional basins with variable thicknesses. The tectonic stress on the depositional basins (Gondwana System) in the Lesser Himalaya of Nepal has resulted in the formation of klippes (such as the Palpa Klippe).

The study area lies within the region 27°42′0″ N to 27°52′0″ N and 83°27′0″ E to 83°33′0″ E in the Tansen area (Palpa district) in the northern part of the Siwalik zone and to the north of the MBT (**Figure 1B,C,D**). The sampling sites along the so-called Palpa section are located near Charchare village along the Bhaiskatta Khola (the local term for the river is Khola) flowing south of the Tansen city. In the Tansen area, the Kali Gandaki Supergroup (KGS) comprises bioturbated sediments to low-grade



FIGURE 1 Geological framework of Nepal. (A) Simplified geological map of the Nepal Himalaya showing the location of the study area (Palpa section) in the Tansen area, modified after Khatri et al. (2017). (B) Detailed geological map of the western Nepal. The black rectangular box indicates the Palpa section, modified after Sakai (1983), Sakai (1984), Dhital (2015). (C) Google map (http://earth.google.com/) showing the sampling route in the Bhaiskatta Khola from 27°47′40.74″ 83°32′39.38″ to 27°47′45.62″ 83°32′40.94″. (D) Detailed geological map of the study area and the cross-section for the samples profile showing the late Cretaceous to Oligocene-early Miocene strata.



metamorphic rock, argillaceous, carbonate rocks with stromatolites of late Precambrian to early Paleozoic age (Sakai, 1984). The Tansen Group, assigned to the late Carboniferous to early Miocene age by paleontology and stratigraphic position (Sakai, 1983), disconformably overlies the KGS and forms a large synclinorium with a klippe (Figure 1B). Lithostratigraphically, the Tansen Group is divided from bottom to top into the Sisne Formation with mudstone intercalated sandstone, and conglomerate; the Taltung Formation with silty shale, sandstone, and conglomerate; the Amile Formation with quartz sandstone with thick interlayers of black shale and limestone; the Bhainskati Formation with shales and limestone; and the Dumri Formation with sandstone and conglomerate interbedded with shale (Sakai, 1983; Sakai, 1984; Sakai, 1989). The Amile Formation typically consists of terrestrial and marine facies. The lower and upper parts of the Amile Formation are terrestrial facies composed of ferruginous

quartz arenite, conglomeratic quartz arenite, and interbedded sandstone and shale, whereas the middle part is marine facies composed of argillaceous limestone and siltstone/sandstone yielding molluscs, echinoids, corals, and vertebrate fossils. In contrast, the marine Bhainskati Formation is characterized by the predominance of fossiliferous black and green shales yielding molluscs, foraminifers, and vertebrate fossils. It also includes bioturbation with thin beds of limestone and oolitic hematic mottles (Sakai, 1983; Sakai, 1984) with hematitic oxisol layers (DeCelles et al., 1998).

SAMPLING AND LABORATORY PROCEDURES

The study area covers late Mesozoic to early Paleogene rocks lying in the southwestern part of the Lesser Himalaya of Nepal, including the Amile, Bhainskati, and partly Dumri formations (Figure 2). In the study area, 120 samples were collected (Figure 1C) with a portable petrol-powered drill at an interval of 1 m. Later, these core samples were cut into standard-size of 2.2 cm diameter and 2.5 cm in length. The Amile Formation in the study area is approximately 72 m thick and is composed mainly of shale, fine-grained to medium-grained siliceous sandstone, and pebble conglomerate, and occasionally contains iron nodules and asymmetrical ripple marks. The Bhainskati Formation is approximately 41 m thick and composed of light gray, dark gray, green shale, and siltstone with no fossiliferous horizons. The Dumri Formation is composed of gray sandstone, gray-red intraformational conglomerate, and gray siliceous sandstone and is approximately 7 m thick (Figure 2).

All the sampled specimens of the Palpa section were measured by the AGICO MFK1-FA multifunction Kappabridge to determine low- and high-frequency magnetic susceptibility ranging from 976 to 1,561 Hz. The mass-specific susceptibility (χ) was measured at these frequencies and the χ fd% was calculated as $\chi fd\% = (\chi_{976Hz} - \chi_{1561Hz})/\chi_{976Hz} \times 100\%$. The anhysteretic remanent magnetization (ARM) process was carried out in a 100 mT peak alternating field and a superimposed 0.05 mT direct current biasing field with a D-2000 alternating demagnetizer, and later the susceptibility of ARM (χ ARM) was calculated as ARM/0.05 mT. The saturation isothermal remanent magnetization (SIRM) process was performed in an IM-10-30 pulse magnetizer with a 1,000 mT field. In addition, the "hard" isothermal remanent magnetization (HIRM) was calculated using the formula HIRM = (SIRM + IRM-300mT)/2 (Thompson and Oldfield, 1986).

Altogether, 46 samples of shales, siltstones, and sandstones were measured for diffuse reflectance spectroscopy (DRS). As it shows a high sensitivity to the concentration of iron oxides (Scheinost et al., 1998; Torrent and Barrón, 2002). The data obtained from a Purkinje General TU1901 UV-Vis spectrophotometer with a reflectance sphere from 360 to 850 nm at 1 nm intervals were processed by the first derivative to reveal the hematite content (Zan et al., 2018). The determination of redness (percent reflectance in the red color band) offers an additional quantitative view of one aspect of sediment color (Ji et al., 2005). Redness percent was calculated as red area (630-700 nm)/peak area (400-700 nm) × 100. Similarly, eight representative samples in powdered form with a mass of ~0.2 g were chosen for thermomagnetic analysis of magnetic susceptibility, each measured by using a CS-4 apparatus in an argon atmosphere coupled to a AGICO MFK1-FA multifunction Kappabridge, to prevent oxidation during heating from room temperature to ~700°C and later cooled to room temperature. The temperature rate was used as 12°C/min (medium-type), with an approximate duration of 2 h. These representative samples were further measured by a Lakeshore Model 8,600 vibrating sample magnetometer (VSM) to determine the magnetic parameters, such as induced magnetizations [hysteresis parameters-hysteresis loops, saturation magnetization (Ms), saturation remanent magnetization (Mrs), and coercivity (Hc)] and remanent magnetizations in variable magnetizing fields. An alternating gradient force of a magnetometer was used for hysteresis loops

analysis and pulse magnetizer for Isothermal Remanent Magnetizations (IRMs), with a maximum field of 1,000 mT. Magnetic coercivity components acquired from IRM acquisition measurements were achieved using the MAX Unmix application (Maxbauer et al., 2016). These experiments were carried out in the Key Laboratory at the Institute of Tibetan Plateau Research, Chinese Academy of Sciences, Beijing, China.

RESULTS

χ , χ fd%, χ ARM, Saturation Isothermal Remanent Magnetization, High Field Isothermal Remanent Magnetization and Diffuse Reflectance Spectroscopy Results

On the basis of χ , frequency-dependent χ (χ fd%), χ ARM, SIRM, high field isothermal remanent magnetization (HIRM), redness, hematite content, and lithology, the Palpa section can be divided into five-rock magnetic zones (Figure 3). These magnetic zonations cover both terrestrial and marine deposits. The zone I (0-30 m), zone IIIA (47-72 m), and zone V (114-120 m) are defined as terrestrial deposits, and zone II (31-46 m), zone IIIB (73-98 m), and zone IV (99-113 m), as marine deposits (Figures 2, 3). The χ values for zone I, IIIA and V range from -0.10 \times $10^{-8} \text{ m}^3 \text{kg}^{-1}$ to $15.41 \times 10^{-8} \text{ m}^3 \text{kg}^{-1}$, from $-0.12 \times 10^{-8} \text{ m}^3 \text{kg}^{-1}$ to 18.05 \times $10^{-8}\,m^3kg^{-1},$ and from 3.17 \times $10^{-8}\,m^3kg^{-1}$ to 8.53 \times 10^{-8} m³kg⁻¹, respectively and their respective mean are 3.25 \times 10^{-8} m³kg⁻¹, 6.29 \times 10^{-8} m³kg⁻¹, 6.29 \times 10^{-8} m³kg⁻¹ and 5.73 \times 10^{-8} m³kg⁻¹. The sandwich of some nonferromagnetic phases, such as coal-bearing shale and thin carbonaceous layers cause a sudden drop in γ in the siltstone and/or sandstone layers. In contrast, χ fd% is always lower in samples with high χ , SIRM, and HIRM values. The γ ARM values are almost 1 or 2 times higher in the terrestrial facies than in the marine facies. In addition, the χ values for zone II, IIIB and IV range from $20.08 \times 10^{-8} \text{ m}^3 \text{kg}^{-1}$ to $65.81 \times 10^{-8} \text{ m}^3 \text{kg}^{-1}$, from $0.08 \times 10^{-8} \text{ m}^3 \text{kg}^{-1}$ to $23.87 \times 10^{-8} \text{ m}^3 \text{kg}^{-1}$ and from $7.72 \times 10^{-8} \text{ m}^3 \text{kg}^{-1}$ to $77.94 \times 10^{-8} \text{ m}^3 \text{kg}^{-1}$, respectively, and their respective mean are $38.26 \times 10^{-8} \,\mathrm{m^3 kg^{-1}}$, $8.99 \times 10^{-8} \,\mathrm{m^3 kg^{-1}}$ and $26.26 \times$ $10^{-8} \,\mathrm{m^3 kg^{-1}}$. A small χ peak in the zone I occur due to the presence of ferruginous matter in the sandstone.

Noticeable sharp changes in χ , SIRM, and HIRM occur in the fine sediments of the marine facies, and low values usually occur in the coarser sediments of the terrestrial facies. The SIRM and HIRM are consistent with the χ , and the highest peaks are 21.01 × $10^{-8} \text{ m}^3 \text{kg}^{-1}$ and 26.42 × $10^{-8} \text{ m}^3 \text{kg}^{-1}$, respectively, at 41 m. The redness and hematite content show trends similar to those of the rock magnetic parameters (**Figure 3**), which are almost consistent with the magnetization and hysteresis parameters of **Figure 4**. As χ , $\chi fd\%$, χARM , and SIRM were investigated for identifying variable magnetic mineral concentrations, further investigation such as temperature magnetic susceptibility (κ -T), induced magnetizations (hysteresis loops), and IRM curves from rock samples in variable magnetizing fields are also discussed below for detailed magnetic mineralogy of the sediments that enhance paleoenvironmental study.



Their comparison with global sea level (Müller et al., 2008) and correlation for the Palpa section, modified after Sakai (1983). The zones I (0–30 m), IIIA (47–72 m), and V (114–120 m) are defined as terrestrial deposits and zones II (31–46 m), IIIB (73–98 m) and IV (99–113 m) as marine deposits. Note that the orange color represents marine and white represents the terrestrial environment.



FIGURE 4 Lithology and hysteresis parameters (Ms, Mrs, Hc and Mrs/Ms) variations of the late Cretaceous to middle Eocene strata in the Palpa section. The samples within the marine regime show greater peaks of hysteresis parameters as compared to the terrestrial regime. The change in sea level condition may have affected the development of these peaks.

High-Temperature Magnetic Susceptibility (κ-T Curves)

Figure 5 displays each of the κ -T curves with a noticeable decrease near 580°C, which suggests the ample occurrence of magnetite in the sediments. A subsequent drop in heating curves (Figures 5A–C) until around 700°C in siliceous and silty sandstone samples, indicates magnetite as the dominant magnetic carrier. But, the heating curve of the sample at 2 m (Figure 5A), possibly implies a maghemite-like phase (drop at about 350°C). There is evidence of the neoformation of magnetite in most samples (Figures 5B–F,H), either by reduction due

to burning of organic matter and/or neoformation of minor ferromagnetic phase from iron-bearing silicates or possibly Fe-bearing paramagnetic minerals (most likely iron sulfides such as pyrite and greigite). In these samples, the heating curve firstly decreases steadily to 300°C and then increases between 350°C and 520°C. Meanwhile, the heating curve progressively decreases nearly to zero at 690°C in the siltstone and shale (**Figures 5D-H**), inferring some hematite as a magnetic carrier (Shouyun et al., 2002; Kruiver et al., 2003; Ao et al., 2010). Thus, both magnetite and hematite act as a major magnetic carrier in these samples.



FIGURE 5 | Representative temperature-dependent susceptibility curves from selected samples at various horizons of the Palpa section. The samples with a low χ value are shown in (A–D) and a high χ value in (E–H). Note: Heating and cooling cycles are indicated by red and blue lines, respectively.



and a high χ value in (E–H). Note: Black and red lines indicate the curves before and after paramagnetism correction.

Hysteresis Loops

The hysteresis loops show that the intrinsic coercivities for all representative samples range from 5.51 to 193.76 mT. The very low Hc ranges from 5.5 to 10.6 mT, revealing ferromagnetic minerals as the major contributor. The hysteresis loops are either sigmoid-shaped slightly open to thin straight lines (closed type) at low fields (**Figures 6A–C**) and sometimes noisy (**Figure 6D**) but one with negative χ (at 65 m),

suggesting the predominance of low-coercivity ferromagnetic minerals (e.g., magnetite) masked by para- and/or diamagnetic contributions. However, the wasp-waisted (Figures 6E–G) and flat and wide shapes (Figure 6H) indicate high coercivity, indicating a hard magnetic component (i.e., hematite). Generally, the wasp-waisted shapes are observed (Figures 6E–G) in samples with less than 10% of hematite or more than 50% hematite (Da Silva et al., 2012). It is thus worth



pointing out that these samples are typical for mixtures of lowand high-coercivity components and/or magnetic grain sizes (Roberts et al., 1995; Tauxe et al., 1996).

Isothermal Remanent Magnetization

Typical IRM acquisition curves (**Figures 7A–C**) show a sharp rise below 200 mT and display weak magnetic intensities with very noisy intensities, representing the low-coercivity ferromagnetic minerals or soft magnetic minerals (e.g., magnetite and/or maghemite) as the major magnetic carrier. In contrast, the higher intensities of magnetization are represented by the gradual growth of IRM curves beyond 200 mT and the fact that the curves do not saturate completely in a 1,000 mT field, signifying the occurrence of some hard magnetic minerals of high coercivity, most likely hematite, contributing to the remanence (**Figures 7A,D-H**). A slow increase in IRM can be observed after the applied magnetic field

reaches 300 mT, indicating less hard magnetic minerals with a relatively high coercive force exist, which is consistent with the thermomagnetic curves of magnetic susceptibility (**Figure 5**).

Similarly, the methods for demagnetization curves or decomposing IRM acquisition provide insights into the low- and high-coercivity mineral assemblages (Heslop et al., 2002; Egli, 2004; Maxbauer et al., 2016). After decomposing the IRM acquisition data, the representative samples (Figure 7B) reveal three- and twocomponent model fits. The three distinct components were obtained from the IRM component analyses (Figures 7A-C,E,F) (Maxbauer et al., 2016). The parameter that shows the mean coercivity of an individual grain population from an assemblages of grains of a single magnetic grain is defined by the parameter B_h. Dispersion parameter (DP) is another parameter used as a measure of the variability of the physical and chemical processes that affect the grain microcoercivity. In order to calculate the relative contribution of each component to the total measured magnetization for each model component, the integrated area under individual component distributions are determined by the magnetic mineral contribution factor called as observed (OC) and extrapolated (EC) contribution (Egli, 2004; Maxbauer et al., 2016). The major contributor to the curves (Figures 7A-C) (component 1) have a mean coercivity of an individual grain population (B_h) of 1.31, 1.26, and 1.49, respectively, and a dispersion parameter (DP) of 0.30, 0.69 and 0.49, respectively, indicating distinctive pedogenic or detrital soft components close to magnetite-like or phase. However, component 1 varies from 1.53 to 1.88, and the DP has values of 0.33 and 0.42 (Figures 7B,E,F), typically indicating partially oxidized pedogenic magnetite (or hematite). Intermediate component 2 (Figures 7A-C) is characterized by a B_h of 0.5, 0.74, and 0.91 and a DP of 0.34, 0.58, and 0.74, respectively, comparable to the partially oxidized pedogenic or detrital soft component, but B_h varied from 1.17 to 1.35 and DP varies from 0.32 to 0.52 in Figures 7B,E,F, which are typical of detrital soft components. The final component 3 (Figures 7A-C) features a B_h of 2.24, 2.32, and 2.15 and a DP of 0.78, 0.53, and 0.47, respectively, indicating detrital soft components, but (Figures 7B,E,F) has a Bh of 0.62 and 2.23 and a DP of 0.21 and 0.34, indicating partially oxidized pedogenic magnetite (or hematite). The decomposition of the IRM acquisition data reveals two primary magnetic components: component 1, characterized by a Bh of 2.06, 2.91, and 3.29 and a DP of 0.66, 0.4, and 0.81, respectively, and component 2, characterized by a B_h of 1.37, 1.43, and 1.71 and a DP 0.52, 0.44 and 0.27, respectively (Figures 7B,D,G,H). These components reveal finegrained hematite (component 1) and partially oxidized pedogenic magnetite (component 2) with either some detrital magnetite or a rare biogenic soft component (Egli, 2004; Lascu and Plank, 2013; Maxbauer et al., 2016).

DISCUSSION

Magnetic Concentration-dependent Parameters From Different Magnetic Minerals in Relation to Tectonics

The magnetic concentration-dependent parameters such as χ , χ ARM, SIRM, and HIRM are shown in **Figure 3**. In **Figure 3**, the

low values of magnetic parameters in the terrestrial facies suggest low contents for the total magnetic mineral assemblage and antiferromagnetic components. These cases could be caused by the dissolution of fine-grained magnetite and the dissolution of antiferromagnetic minerals through microbial activities. The other possible driving factor could be the irregular deposition of sediments with low heavy mineral concentrations (Snowball and Thompson, 1990). In Figure 4C, the maximum Hc was observed in between 31-46 m and 99-113 m (marine facies). In addition, 2-29 m and 49-73 m (terrestrial facies) and also in 74-97 m (marine facies) display relatively lower Ms values, Hc, and Mrs/Ms ratios (Figures 4A,B,D). When very strong linear relationships exist between Mrs and χ , it suggests that the source of the magnetic minerals remains unaffected (Caitcheon, 1993; Eriksson and Sandgren, 1999; Meena et al., 2011). However, in **Figures 8A,B**, Mrs and χ show a very poor linear relation (R^2 = 0.0006) for 31–46 m and $R^2 = 0.1321$ for 73–113 m, inferring more possible sources for these existing magnetic inclusions (Figure 3) in the marine realm in the study area. In addition, a positive linear relationship ($R^2 = 0.8251$ and 0.883) exists between Ms and Mrs (Figures 8C,D) that denotes the magnetic concentration is controlled by magnetic mineralogy.

Figures 3, 4 show two distinct breaks in the curves respective to the significant change in the depositional setting (Figure 2), showing higher concentrations of primary magnetic grains with higher coercivity mineral assemblages, especially antiferromagnetic minerals or harder minerals, such as hematite. The presence of paramagnetic minerals (e.g., Fesilicates and Fe-oxyhydroxides) or weak magnetic oxides in the study area show low χ values (Figure 3A), probably associated with the allochthonous type and migrated through different fluvial channels or in situ environments. A gradual increase in χ fd% (Figure 3B) occurs in the siltstone, siliceous sandstone, and conglomerate beds of terrestrial facies, while the shale and siltstone beds bear less-weathered mineral grains resulting in higher χ and lower χ fd% spikes in the marine facies. In contrast, the marine facies exhibit higher SIRM, HIRM, and χ values but lower χ ARM values than the terrestrial facies (Figure 3A,C-E). The greater values of these magnetic concentration-dependent parameters (Figure 3A,D,E) are probably induced by the drying phase of soil wetting/drying cycles through magnetite and Fe³⁺ oxide dissolution and/or drying by Fe³⁺ oxide during neoformed ultrafine magnetites (Fischer, 1988; Maher, 1998). Meanwhile, the yARM shows higher values (Figure 3C) formed by weak early diagenesis. Generally, the weak early diagenesis is affected by the local climatic environment during intense moist and sporadically reduced conditions, exhibiting wet and/or wet and warm conditions (Maher, 1998; Maher and Thompson, 1999; Maher et al., 2003). During this time, minor fluctuations in sea level have occurred (Müller et al., 2008), which in turn enhanced precipitation-induced erosional rate and deposits influx of magnetite- and hematite-bearing channel siltstone and shale (Figures 3G,H) into the existing basin or the continental shelf. Importantly, there was a constant collision-induced uplift of the Himalayan-Tibetan Plateau in this region. This leads to change in the sedimentary depositional environment (terrestrial and/or



FIGURE 8 | (A-B) Bivariate plot of χ versus Mrs, a very low linear relation among both the concentration dependent. **(C-D)** Linear relation between Ms and Mrs display the magnetic mineral concentration for marine environment (31–46 m and 73–113 m).

marine environments) with soft and hard magnetic minerals (Figures 8A,B) and provenance, which altered the channel inflows, continental weathering, erosion rate, and production of organic material within the depositional basin and modified the magnetic mineralogy to higher coercivity mineral assemblages (Figures 8C,D and Figure 9). The active tectonics highly influenced the concentrations of ferromagnetic minerals, such as magnetite, and hard minerals (e.g., hematite), which govern the magnetic signal within deposits forming in depositional environments that can be affected by changes in the regional climate (Figure 9).

Prevailing Tectonic Control on the Depositional Basin

The dynamic geomorphological features formed as a result of collision and uplift of the Himalaya, inducing cycles of transgression and regression, precipitation, and weathering causing rapid denudation. During the course of northward advancement of the Indian Plate, the basin plains might have had low areas with relatively constant deposition, in which changes in subsidence caused shifts between terrestrial and marine environments, adjusting the basin framework (e.g., the basins of Barahakshetra in the east, Tansen in the center, Tulsipur in the west and Birendranagar in the far west of Nepal). A simpler model can be postulated for the paleodepositional environment assessment of the late Cretaceous to middle Eocene strata (Figure 9). Zone I starts with a basal fluvial detritus in the form of siliceous sandstone followed by a pebbly conglomerate (Figures 2, 3), through local high-energy ephemeral channel systems derived from the adjacent highlands (not highlyelevated mountains), as a result of erosional processes, in response to active tectonics. These sediments are enriched in Fe-silicates and Fe-oxyhydroxides and some ferruginous nodules, probably showed earlier diagenetic processes under anoxic conditions (reducing conditions) developed by bacterial activities in the organic materials, and the high sulfate content possibly under brackish conditions, and occasionally under nonmarine conditions (Figure 2 and Figure 9A) (Harder, 1989; Yoshida et al., 1998; Shouyun et al., 2002). In addition, the multiple thin coal layers in zones I, II, and IIIA, with some ash and sulfur, sandwiched between greenish-gray sandstone and siltstone in zone I (Figure 2), originated from a peat swamp in a thick forest moor in a humid and marshy environment. These deposits probably formed within a reducing environment during the still-stand phase of marine transgression (Potter et al., 1980; Riegel, 1991; Sarkar and Prasad, 2000; Singh et al., 2010). Besides, the zones I, IIIA, and V consist of sandstones, conglomerates, and interbedded shales and sandstone which exhibit comparatively lower rock magnetic parameters values (χ , SIRM, and HIRM



values) (Figure 3), than zones II and IV and characterizes soft magnetic mineral assemblages. The repeated interbedding of carbonaceous sandstone and shale in zones I and IIIA (Figures 2, 3) by prograding low-energy fluvial systems can be linked with marginal marine influences, such as delta front environments (Figure 9B) (Mazumder et al., 2017). Meanwhile, the asymmetrical ripple marks in light gray siliceous sandstone (e.g., zone IIIA) also indicate deposition under fluvial conditions (Debenay and Guillou, 2002). These fluvial cycle changes after the transgression caused thick graygreen sandstone and occasional gray-red sandstone deposits in zone IIIA (Figures 2, 3).

Besides, the red-yellow weathering in the carbonaceous siltstone (e.g., zone II) and the gray-green silty shale followed by dark carbonaceous shale (e.g., from zone IIIB to IV) (**Figure 2**), formed due to higher organic production and relative oxygen-deficient bottom water caused by dysoxic or anoxic conditions, probably in a warm climate (**Figure 9A**) (Wignall, 1991; Lash and Blood, 2014). They show relatively high values of χ , SIRM, and HIRM, though not χ ARM, and relatively low values of χ fd% (**Figure 3**). These high values are due to higher coercivity mineral assemblages such as the predominance of hematite over magnetite. The formation of a black shale to the upper part of green-red shale (e.g., zone IV) (**Figure 2**), indicates erosion from surrounding highlands and a eustatic sea level rise leading to marine invasion (**Figure 3H**).

This condition usually occurs in calm water with low flow velocities associated with a weakened flooding phase (Allen, 1980; Sahni et al., 2004; Sisodia and Singh, 2000), and signifies the gradual regression of the Tethys Sea (Valdiya, 1980; Wells and Gingerich, 1987; Mathur, 1990; Srivastava and Kumar, 1996) (in Nepal, we termed it the Bhainskati Sea) with a hiatus in the deposition. This type of sedimentation in a euxinic environment might have been developed in a shallow sea with some tidal influence, such as in a protected lagoon with mud zones and tidal flats (Figure 9B) (Wallace-Dudley and Leckie, 1993; Willis et al., 1999; Bhatia and Bhargava, 2006; Singh et al., 2010; Krim et al., 2017). The fine sediments (e.g., zone II) are thought to have been derived from the Indian cratonic succession and the remobilization of previously deposited arenaceous sequences (Garzanti, 1999), originally by a northward-flowing drainage system in the late Cretaceous time (Gansser, 1964; Sakai, 1983; DeCelles et al., 1998; DeCelles et al., 2004; DeCelles et al., 2014). However, the sediments of zone IIIB and IV include mixed recycled Indian craton concealed basement rocks (probably from the northeastern volcanic rocks of India) and the Himalayan units (probably the Tethyan sedimentary rock sequences) (DeCelles et al., 2004; Gehrels et al., 2011). The iron ore deposits in high-grade metasedimentary and lowgrade volcano-sedimentary rocks in the Singhbhum and Bundelkhand Cratons of greenstone belt successions (from the Indian subcontinent) (Saha, 1994; Mukhopadhyay et al., 2008;

Chattopadhyay et al., 2015; Mukhopadhyay, 2020) might have been transported to the Amile and Bhainskati formations. Also, the metasediments from the Bastar Craton (Dora et al., 2020) and other metamorphosed rocks from the Aravalli Craton (Crawford, 1970) are thought to be present occasionally in the late Cretaceous to middle Eocene strata.

It is interesting to point out that the relative rise in sea level drowned the southern highlands, and a shallow marine environment developed briefly in the late Cretaceous strata and ceased with a progressive shift to continental sedimentation. Some larger foraminifera, oysters, sharks, pelecypods (bivalves), gastropods, echinoids, screleactinean corals, and vertebrate fossils are reported from the argillaceous limestone of the Amile Formation (middle part) (e.g., zone II) (Sakai, 1983), acted as barriers (Figure 9B) (Wright and Burchette, 1996). Furthermore, the larger foraminifers (Sakai, 1983; Matsumaru and Sakai, 1989), bivalves, gastropods, Teleostei, Asteracantus sp., Chelonia and Trionichidae carapaces (Sakai, 1983), and pristichampsinae (Sah and Schleich, 1990) in the Bhainskati Formation (e.g., zone IV) have been interpreted as indications of an aquatic environment, allowing productive shallow sea level conditions (Reineck and Singh, 1980). As the reflooding caused the deposition of basal intraformational conglomerate (mudpebble conglomerate) beds (e.g., zone v) (Dumri Formation) (Figure 2) eroded from the northern uplifted mountain ranges during the late Oligocene to early Miocene. It represents the final regression uplift caused by the tectonic stress in the Himalayan belt. The cessation of deposition could correspond to the timing of the India-Asia collision and the disappearance of the Bhainskati seaway in the Nepal region as a result of Himalayan uplift. Furthermore, the lack of mud cracks in the study area indicates that the basin was not completely in the shallowing stage and that exposure of the sediments did not occur under totally dry conditions. We also argue that the local geomorphology played a significant role in the development of magnetic mineralogy in this section and that the nature of sediment influx (of either autochthonous or allochthonous origin) was influenced by local drainage patterns and seawater conditions in the depositional basin, with enhancement by active tectonics.

CONCLUSION

The results suggest the complex mechanisms operating during the collision between the Indian and Eurasian plates and the early stages of uplift of the Himalayan orogen. Firstly, the major change inferred from abrupt breaks in the χ , SIRM and HIRM, and lithological differences, which are concluded to have developed due to the modification of the depositional environment during the deposition of the late Cretaceous to middle Eocene strata.

Secondly, the active tectonics influenced SIRM, HIRM, and hysteresis loops with the concentrations of ferromagnetic minerals, such as magnetite, and hard minerals (e.g., hematite) formed in the terrestrial (zones I, IIIA, and V) and marine (zones II, IIIB and IV) facies respectively. We suggest that the forementioned rock magnetic changes from the sediments during the late Cretaceous to middle Eocene had been affected by the alternating warm-dry and warm-humid regional climate associated with the regression and transgression of the Tethys Sea as a result of the Indian-Eurasian collision.

DATA AVAILABILITY STATEMENT

The datasets presented in this study can be found in online repositories. The names of the repository/repositories and accession number(s) can be found below: http://figshare.com/s/ 6de59e47d56dd81aca78.

AUTHOR CONTRIBUTIONS

DK: Conceptualization, Methodology, Formal analysis, Writing–Review and Editing, Investigation. WZ: Conceptualization, Formal analysis, Investigation, Resources, Supervision. XF: Investigation, Resources, Supervision. QM: Investigation. TZ: Investigation. DZ: Investigation. KP: Investigation.

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