



Effect of Magnesium Incorporation on Solution-Processed Kesterite Solar Cells

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The introduction of the alkaline-earth element Magnesium (Mg) into $\text{Cu}_2\text{ZnSn}(\text{S},\text{Se})_4$ (CTZSSe) is explored in view of potential photovoltaic applications. $\text{Cu}_2\text{Zn}_{1-x}\text{Mg}_x\text{Sn}(\text{S},\text{Se})_4$ absorber layers with variable Mg content $x = 0 \dots 1$ are deposited using the solution approach with dimethyl sulfoxide solvent followed by annealing in selenium atmosphere. For heavy Mg alloying with $x = 0.55 \dots 1$ the phase separation into Cu_2SnSe_3 , MgSe_2 , MgSe and SnSe_2 occurs in agreement with literature predictions. A lower Mg content of $x = 0.04$ results in the kesterite phase as confirmed by XRD and Raman spectroscopy. A photoluminescence maximum is red-shifted by 0.02 eV as compared to the band-gap and a carrier concentration N_{CV} of $1 \times 10^{16} \text{ cm}^{-3}$ is measured for a Mg-containing kesterite solar cell device. Raman spectroscopy indicates that structural defects can be reduced in Mg-containing absorbers as compared to the Mg-free reference samples, however the best device efficiency of 7.2% for a Mg-containing cell measured in this study is lower than those frequently reported for the conventional Na doping.

Keywords: thin film solar cells, kesterite, Mg, solution processing, structural defects

INTRODUCTION

Kesterite-type material $\text{Cu}_2\text{ZnSn}(\text{S},\text{Se})_4$ (CZTSSe) has been recognized as a promising candidate for low-cost thin-film solar cells due to its large absorption coefficient, tunable band-gap E_g between 1.0 and 1.5 eV adjusted via S/Se-ratio, low toxicity and earth-abundant nature. This technology has achieved a 12.6% maximum performance (Wang C. et al., 2014), still far away from that of 22.6% for $\text{Cu}(\text{In},\text{Ga})\text{Se}_2$ solar cells (Jackson et al., 2016). The main performance limitation of kesterite-based solar cells is the open circuit voltage deficit ($E_g/q - V_{\text{oc}}$). One of the reasons is the non-optimal quality of kesterite absorber and the presence of secondary phases (Siebentritt and Schorr, 2012). Another reason can be the unfavorable alignment of the conduction band minimum (CBM) at the CZTSSe/CdS interface (Platzer-Björkman et al., 2015). Gokmen et al. (2013) pointed out that the $[\text{Cu}_{\text{Zn}} + \text{Zn}_{\text{Cu}}]$ defect cluster could be the origin of electrostatic potential fluctuation in the CZTSSe absorber layer. That fluctuation is probably a significant factor that decreases the photovoltaic (PV) device performance. It was suggested that the substitution of Cu or Zn by other elements as Magnesium (Mg) could suppress the antisite defects Cu_{Zn} and/or Zn_{Cu} formation that limit kesterite solar cells efficiency (Zhong et al., 2016).

Several studies about the Mg incorporation into kesterite have recently been reported, but the observed effects of Mg are contradictory. CIGSe bulk material doped with Mg was

deposited by liquid phase sintering method measuring a decreased hole concentration with the increase in Mg content (10 at %, $\text{Mg}/(\text{In}+\text{Mg}) = 0.1$), which was attributed to the Mg_{Cu} donor defect (Monsefi and Kuo, 2014). At the same time, an increased hole concentration was observed for 5 at % Mg explained by the substitution of In with Mg ion to form Mg_{In} acceptor defect. $\text{Cu}_2\text{MgSnS}_4$ thin films grown by ultrasonic co-spray pyrolysis showed p-type conductivity and band-gap energy of 1.76 eV (Guo et al., 2016). In contrary, n-type conductivity was estimated for $(\text{Cu}_{2-x}\text{Mg}_x)\text{ZnSnSe}_4$ bulk materials with $x = 0.1-0.4$, which was attributed to the formation of the donor-type Mg_{Cu} antisite defects (Kuo and Wubet, 2014). The formation of stable $\text{Cu}_2\text{MgSn}(\text{S},\text{Se})_4$ was calculated based on density functional theory (Zhong et al., 2016), whereas a complete phase separation was predicted by Wang W. et al. (2014).

The purpose of this work is to study the effect of Mg addition in various concentrations to CZTSSe solar cell absorbers. Mg is incorporated to the absorber thin films by adding a magnesium salt to the precursor solution. Two experimental series are presented. The first experiment involves the replacement of Zn for Mg concentrations in order to evaluate possible alloying effects on the absorber band-gap for $\text{Cu}_2\text{Zn}_{1-x}\text{Mg}_x\text{Sn}(\text{S},\text{Se})_4$ thin films, however a complete phase separation is observed for $x = 0.55$ and 1. For a lower Mg content of $x = 0.04$ the kesterite phase is obtained, and respective absorbers and solar cells are studied to reveal any structural and electronic effects of Mg-containing sample as compared to a nominally Mg-free one.

MATERIALS AND METHODS

Absorber Preparation

The $\text{Cu}_2\text{Zn}_{1-x}\text{Mg}_x\text{Sn}(\text{S},\text{Se})_4$ thin films absorbers were deposited from the precursor solution with dimethyl sulfoxide (DMSO) as the solvent onto Mo/SiO_x/soda-lime glass (SLG) substrates with a subsequent selenization using the methodology described in the previous work (Haass et al., 2015). 1 mm-thick SLG was cleaned in three different supersonic baths at a temperature of 80°C. The first bath consisted of salt-free water with soap (Borer Deconex), the second bath consisted of a weak acetic solution (~5%) and the final bath contained only de-ionized water (18 MΩ·cm). The residual water was blown off with nitrogen and the substrates were dried in vacuum prior to the subsequent layer deposition. The SiO_x barrier layer was deposited at 200°C substrate temperature on SLG substrate by sputtering of a Si target in an Ar/O₂ atmosphere. The layer thickness of SiO_x is approximately 200-300 nm. The SiO_x barrier layer was used to reduce the alkali elements out-diffusion from the SLG substrate although the presence of Na could not be eliminated since it can also be transported via the gas phase during the annealing step (Abzieher et al., 2016). The CZTSSe precursor solution contained thiourea (99%+, Sigma-Aldrich), $\text{SnCl}_2 \cdot 2\text{H}_2\text{O}$ (98%, Sigma-Aldrich), ZnCl_2 (99.99%, Alfa Aesar) and CuCl_2 (98%+, Alfa Aesar) dissolved in DMSO (99.9%, Alfa Aesar). Mg was introduced by adding $\text{Mg}(\text{CH}_3\text{COO})_2 \cdot 2\text{H}_2\text{O}$ to the precursor solution. **Table 1** shows

the nominal Mg composition and metal ratios in the precursor solutions.

The solution was spin-coated onto a Mo/SiO_x-coated and dried on a hotplate at 320°C in air. The spin-coating and drying steps were repeated 12 times in order to obtain the desired precursor film thickness of approximately 1.5–2 μm (Haass et al., 2015). Samples were annealed in a rapid thermal annealing (RTP) reactor (AS-ONE 150 from Annealsys) inside a silica-coated graphite box with selenium pellets (0.8 g). A three-stage temperature gradient was employed for annealing samples with holding temperatures at 300, 500, and 550°C for 30, 45, and 5 min, respectively. After selenization the absorbers were immersed into a 10 wt% KCN solution for 30 s in order to remove any copper-rich phases and clean the surface from contaminations and oxides (Haass et al., 2015).

Device Preparation

A 50–70 nm thick CdS buffer layer was deposited by chemical bath deposition, and 70 nm/250 nm i-ZnO/Al:ZnO bilayer was sputtered. A Ni/Al top grid and an antireflection coating of MgF_2 were deposited by e-beam evaporation. Individual solar cells were mechanically scribed to an area of $0.30 \pm 0.02 \text{ cm}^2$ (Haass et al., 2015).

Materials Characterization

Inductively coupled plasma mass spectrometry (ICP-MS) measurements were carried out to determine the Mg content in the selenized absorber layers, which is listed in **Table 1**. The ICP-MS measurements were done on the full cells (without antireflection coating) for samples A and B with prior etching of the window layer by 5% acetic acid for 60 s and rinsing in distilled H₂O. For the sample C with much lower Mg concentration we used the ICP-MS results of the absorber layer on SLG (no Mo) without any etching. The Mg content measured by ICP-MS for sample C is 0.04, which is lower than the nominal content of 0.17, indicating that Mg is lost during the absorber fabrication, and the reason needs to be investigated. Secondary Ion Mass Spectrometry (SIMS) measurements were recorded on a TOF-SIMS system from ION-TOF using O²⁺ primary ions with 2 keV of ion energy, a current of 400 nA and a raster size of $400 \times 400 \mu\text{m}^2$. An area of $100 \times 100 \mu\text{m}^2$ in the case of depth profiles was analyzed using Bi⁺ ions with 25 keV of ion energy. Energy dispersive X-ray (EDX) analysis and X-ray fluorescence (XRF) were used to quantify the composition of matrix elements. The XRF measurements were done on selenized absorber layers without any etching. Scanning Electron Microscopy (SEM) and EDX measurements were done on a Hitachi S-4800 electron microscope. X-ray diffraction (XRD) patterns were recorded in 2θ scan mode using a Bruker D8 diffractometer with $\text{CuK}\alpha$ radiation ($\lambda = 1.5418 \text{ \AA}$, beam voltage: 40 kV, beam current: 40 mA, calibrated using Si100 and Si111 single crystals), a step size of 0.04° and a scan rate of 0.5 s/step (Haass et al., 2015). Grazing incidence XRD (GIXRD) measurements were collected with a PaNAlytical X'Pert Pro MPD diffractometer using $\text{Cu K}\alpha$ radiation ($\lambda = 1.5418 \text{ \AA}$, beam voltage: 40 kV, beam current: 40 mA), a step size of

TABLE 1 | Mg content and nominal metal ratios in the precursor solution (sol) and selenized absorbers (abs) for all the samples.

Sample	Nominal Mg content x in $\text{Cu}_2\text{Zn}_{1-x}\text{Mg}_x\text{Sn}(\text{S},\text{Se})_4$ in the precursor solution	Nominal metal ratios in precursor solution (sol) and metal ratios measured by ICPMS and/or XRF* in selenized absorbers (abs)			Actual Mg content measured by ICPMS in selenized absorbers
		Cu/Zn	Cu/Sn	Zn/Sn	
A	$x = 1.00$	N.A.	1.08 (sol)	0 (sol)	$x = 1.00$
		N.A.	1.19 (abs)	0 (abs)	
B	$x = 0.50$	2.46 (sol)	1.08 (sol)	0.44 (sol)	$x = 0.55$
		2.72 (abs)	0.99 (abs)	0.36 (abs)	
C	$x = 0.17$	1.28 (sol)	1.34 (sol)	1.05 (sol)	$x = 0.04$
		1.45 (abs)	1.74 (abs)	1.20 (abs)	
		1.41*(abs)	1.82*(abs)	1.29*(abs)	
D	$x = 0.00$	1.27 (sol)	1.40 (sol)	1.05 (sol)	$x = 0.00$
		1.40*(abs)	1.74*(abs)	1.24*(abs)	

*Measurements carried out by XRF.

0.02°, a scan rate of 2 s/step and a multilayer mirror (Caballero et al., 2015). Detector scans with incident angles of 1°, 3°, and 5° were carried out. Raman scattering measurements were performed in back scattering configuration using a highly sensitive Raman apparatus developed at IREC consisting in a Horiba Jobin Yvon iRH320 spectrometer coupled with a low noise CCD detector cooled at -70°C . In this system, excitation and light collection were made through a macro optic system with a laser spot diameter of about 70 μm . Back-scattering measurements were performed under 633 and 785 nm excitation wavelengths by focusing laser spot directly onto the layer surface which allows the assessment of the absorber layer without any contribution from the upper layers. In the case of back surface measurements, laser spot was directly focused on the back surface after mechanical lift-off process. Excitation power was kept below 26 W/cm^2 in order to avoid presence of thermal effects in spectra. The first-order Raman spectrum of monocrystalline silicon (Si) was measured as a reference before and after each Raman spectrum acquisition, and spectra were corrected by imposing Si first order at 520 cm^{-1} (Oliva et al., 2017).

Device Characterization

The J - V characterization of solar cells was performed under standard test conditions (100 mW/cm^2 , 25°C, AM1.5G) using a solar simulator calibrated with a certified Si diode. External Quantum Efficiency (EQE) spectra were recorded using a chopped white light source (900 W halogen lamp) with a LOT MSH-300 monochromator, which was calibrated with Si and Ge photodiodes. The illuminated area on the sample was 0.1 cm^2 including grid lines. Photoluminescence (PL) spectra were measured on a FT300 fluorescence lifetime spectrometer from PicoQuant with a 639 nm pulsed diode laser as excitation source (pulse width 90 ps, repetition rate 10 MHz) and a thermoelectric cooled Hamamatsu NIR-PMT module H10330A-45 (rise time 0.9 ns, transit time spread 0.4 ns). Capacitance-Voltage (C-V) room temperature measurements were carried out

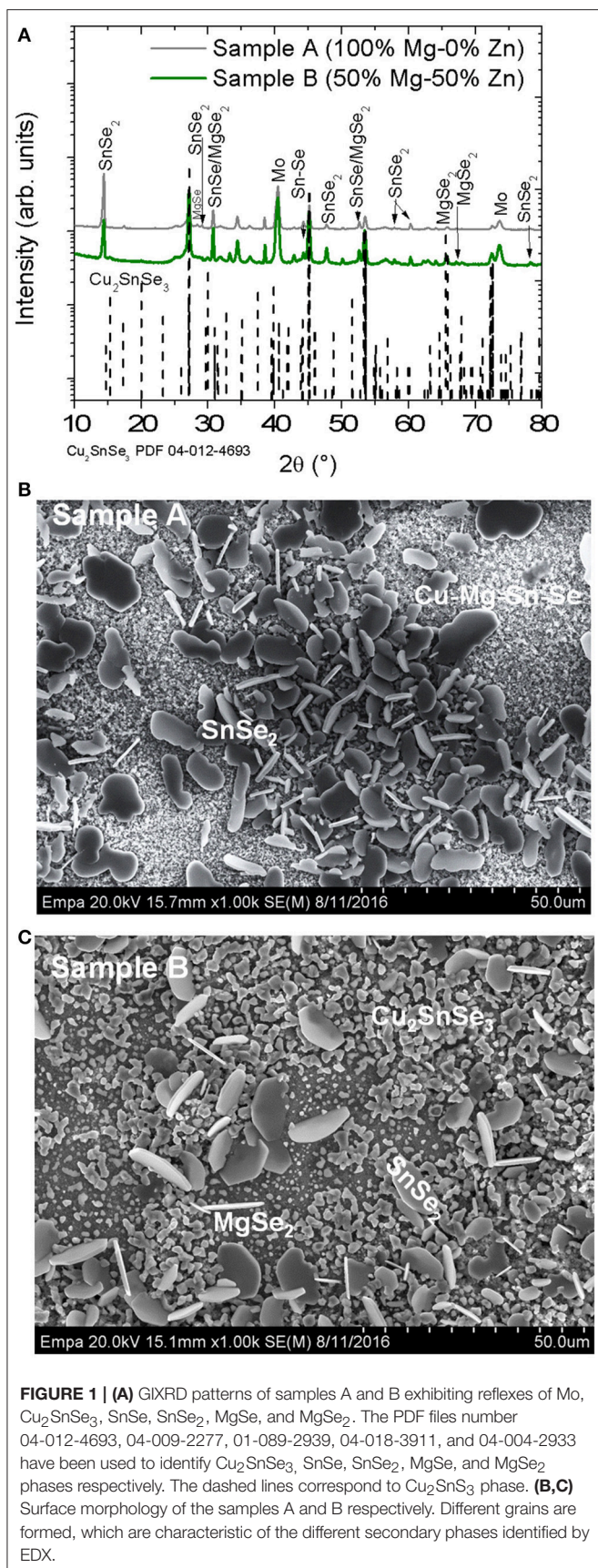
with a LCR-meter from Agilent (E4990A) with an AC-voltage of 30 mV (Haass et al., 2015). J - V -T was carried out in the temperature range from 138 to 298 K. For temperature dependent current-voltage measurements the solar cell was placed on temperature controlled Cu stage inside an evacuated cryostat cooled with liquid nitrogen and illuminated by a 900 W halogen lamp. The sample temperature was measured by thermocouples and regulated by a PID controller. The intensity of the incident light was varied by 2 orders of magnitude from approximately 1 - 140 mW/cm^2 using neutral density filters.

RESULTS AND DISCUSSION

Alloying with High Concentration of Mg: $\text{Cu}_2\text{Zn}_{1-x}\text{Mg}_x\text{Sn}(\text{S},\text{Se})_4$

The first experiment series involved heavy alloying of Zn with Mg in $\text{Cu}_2\text{Zn}_{1-x}\text{Mg}_x\text{Sn}(\text{S},\text{Se})_4$ (CZMTSSe) thin films. **Figure 1A** shows XRD spectra of samples A ($x = 1$) and B ($x = 0.5$). One can observe a complete phase separation, identifying the reflexes corresponding to MgSe_2 , MgSe , SnSe_2 , and Cu_2SnSe_3 in both samples. The presence of CZTSSe and ZnSe phase is difficult to exclude due to the closeness of their diffraction peaks with those of Cu_2SnSe_3 . Already during the deposition of the $\text{Cu}_2\text{Zn}_{1-x}\text{Mg}_x\text{SnS}_4$ (CZMTS) precursor layer, it was observed that the solution was inhomogeneous with some precipitates.

Figures 1B,C illustrate SEM surface morphology of samples A and B. Different types of grains are observed. For sample A ($x = 1$), SnSe_2 is observed on top of smaller grains with the presence of Cu (15.1 at %)-Mg (16.9 at %)-Sn (10.1%)-Se (44.7 at %)-S (13.2 at %) (see **Figure 1B**). For sample B with $x = 0.5$, larger grains corresponding to SnSe_2 and the smallest grains consistent with MgSe_2 and Cu_2SnSe_3 are detected as shown in **Figure 1C**.



The observed phase separation is in agreement with the predictions of Wang et al. who describes the instability of the $\text{Cu}_2\text{MgSnS}_4$ by the following reaction:

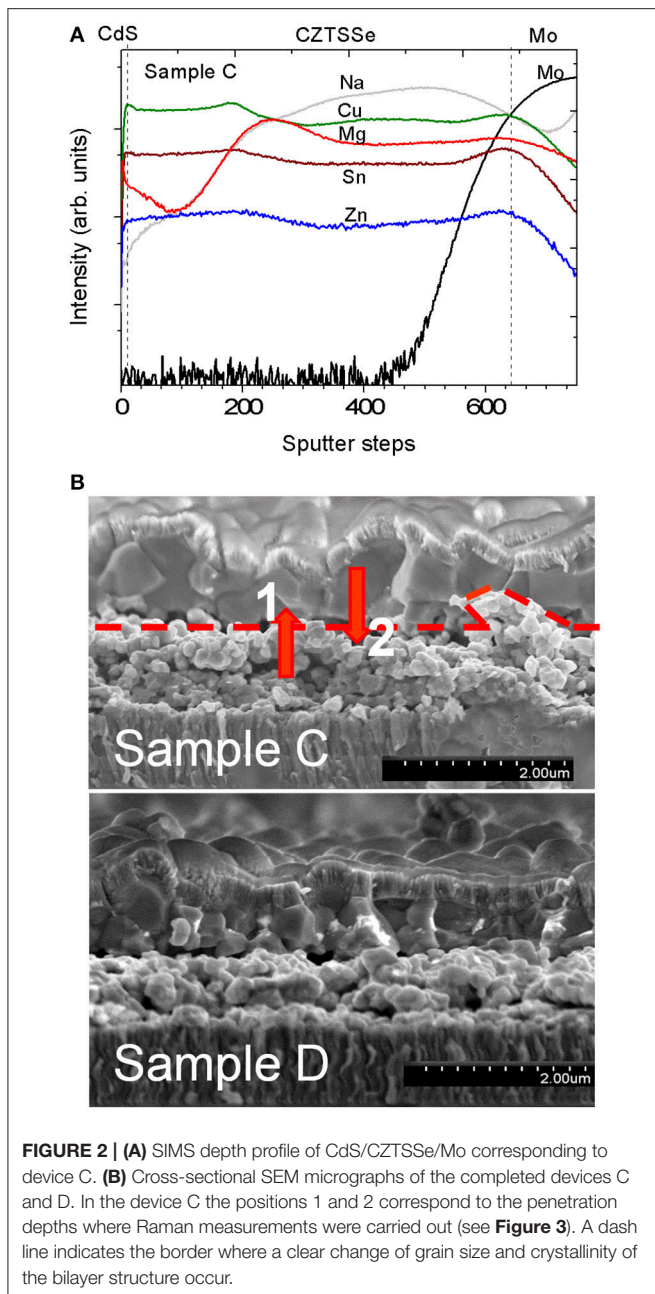


The calculated energy change of this reaction is exothermic ($\Delta E = -0.01 \text{ eV} < 0$), meaning that the phase separation of $\text{Cu}_2\text{MgSnS}_4$ proceeds spontaneously, in accordance with the disappearance of the stable region in the chemical potential space. The instability for this compound is directly related to the fact that for group IIA element, $\text{Cu}_2\text{MgSnS}_4$ is more stable in the ionic rocksalt structure with S than in the more covalent tetrahedral environment as in kesterite. When an $\text{I}_2\text{-II-IV-VI}_4$ compound is unstable, the corresponding elements, in this case Mg, can only be incorporated into a stable $\text{I}_2\text{-II-IV-VI}_4$ with low concentration (Wang W. et al., 2014). No working solar cells could be obtained using samples A and B as a consequence of the phase separation.

Effect of Low Mg Concentration on Kesterite Solar Cells

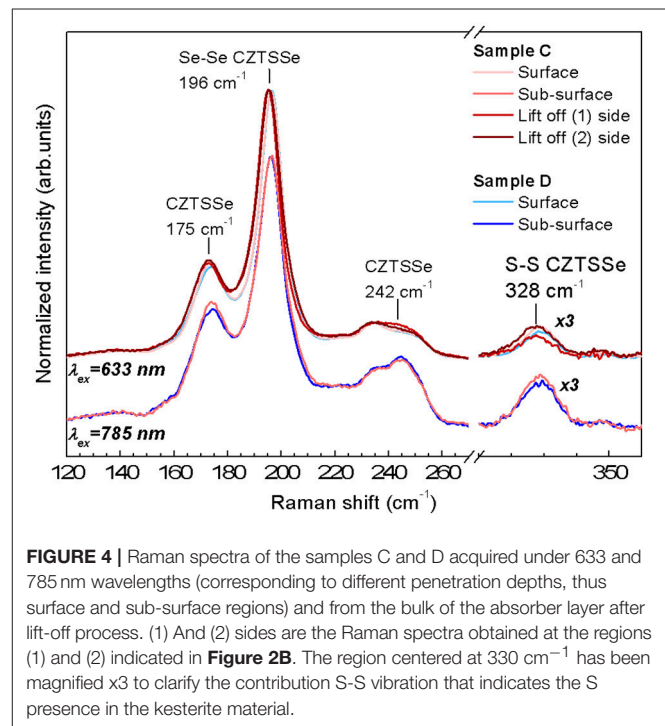
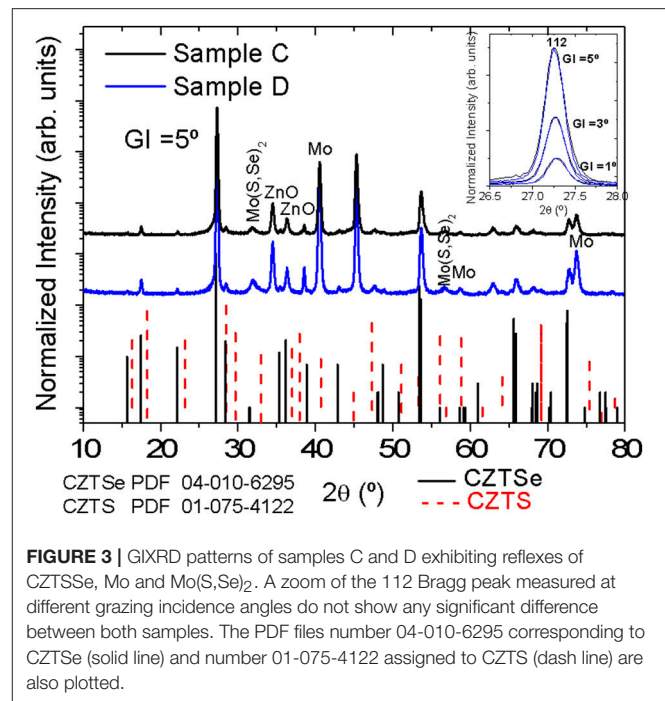
The effect of the addition of low Mg concentration was studied by comparing sample C with a Mg content of $x = 0.04$ (measured by ICP-MS in Table 1) to the nominally Mg-free, sample D. The distribution of Mg inside sample C measured by SIMS is not homogeneous (Figure 2A), with a higher Mg signal in the lower part of the absorber. This can be correlated to the morphology of samples C and D visible in the cross-section. Both absorbers exhibit a bi-layer structure with large-grain material on top and smaller grains close to the back contact as reported for the solution-processed kesterite layers (Haass et al., 2015). We assume that a higher Mg content in the lower part of sample C is due to a larger number of grain boundaries which can accommodate more Mg in comparison to large grains of the top crust. It is worth mentioning that a significant Na signal can be detected by SIMS in both samples C and D despite the fact that no Na was intentionally added to the precursor solution and a SiO_x barrier was employed. ICP-MS measurements were carried out to quantify the Na concentration incorporated into the absorber layer. A Mg-free reference sample had a sodium concentration of only 70 ppm, while samples without SiO_x barrier layer were reported with around 2,000 ppm (Sutter-Fella, 2014). This fact shows the effectiveness of the alkali barrier layer. This confirms that Na can also be transported via the gas phase during selenization (Abzieher et al., 2016).

Figure 3 displays GIXRD diffractograms of solar cells corresponding to samples C and D. Reflexes at 15.7° , 17.5° and 22.2° confirm unambiguously the CZTSSe phase. Moreover, reflexes at 31.8° and 56.6° appear, indicating the presence of the $\text{Mo}(\text{S,Se})_2$ layer formed during the selenization. No other secondary phases can be detected, although the presence of $\text{Zn}(\text{S,Se})$ and $\text{Cu}_2\text{Sn}(\text{S,Se})_3$ phases cannot be ruled out since their diffraction reflexes coincide with those of CZTSSe. Reflexes of ZnO are also visible due to the window layer of $i\text{-ZnO/AZO}$ in complete devices. A zoom of the 112 Bragg reflex of the kesterite phase measured at different grazing incidence angles does not



show significant differences between both samples at different depths.

Raman spectra measurements of the completed devices corresponding to samples C and D are shown in **Figure 4** for two wavelengths of 633 and 785 nm. At these excitation wavelengths the contribution from the ZnO and CdS layers do not interfere with the absorber thin film signal allowing the direct characterization of the surface (penetration of <70 nm for the 633 nm) and subsurface regions (<150 nm for the 785 nm) (Dimitrievska et al., 2016b; Oliva et al., 2016). In order to access CZTSSe bulk, additional Raman measurements under 633 nm excitation were performed at the back of the Mg-containing CZTSSe after its mechanical delamination (lift-off process),

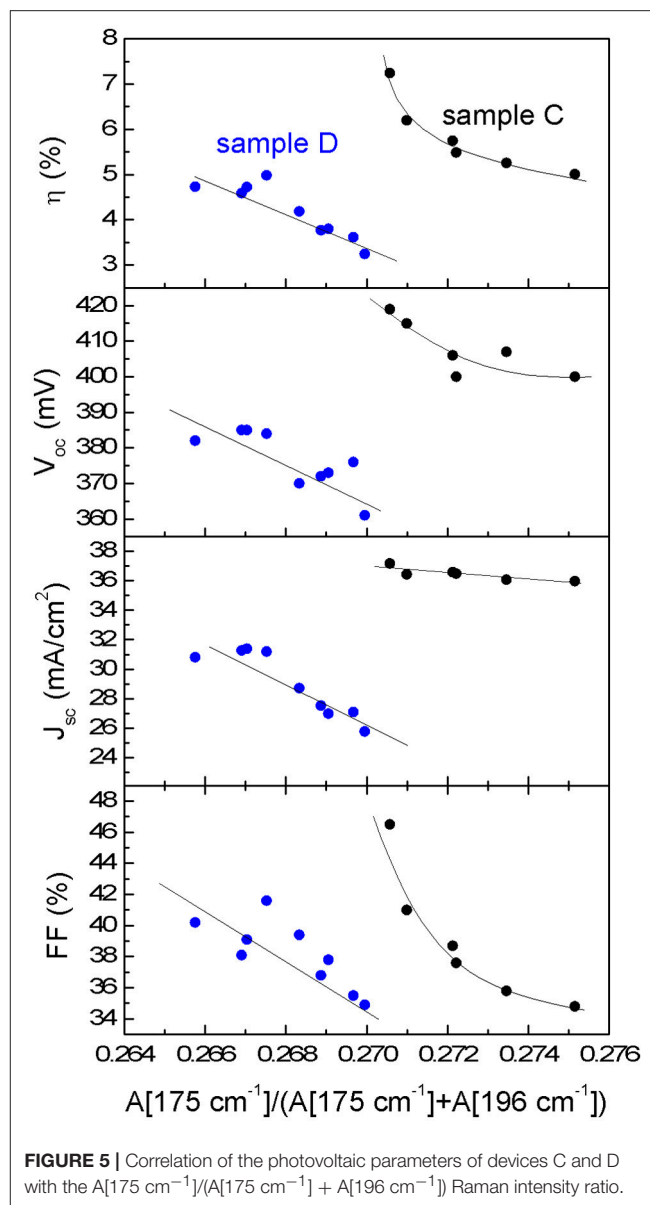


corresponding to interface regions 1 and 2 (shown as red dotted line in **Figure 2B**). All Raman spectra exhibit characteristic features of highly Se-rich CZTSSe solid solution ($[S]/([S]+[Se]) \approx 3\%$ evaluated by Raman spectroscopy) yet. A detailed analysis of the Raman spectra of samples C and D can bring more insights about the impact of low Mg concentration on the CZTSSe absorber. Raman related parameters are summarized in

Tables S1–S3 (See Supporting Information). The analysis of the Raman shift (RS) and the full-width-at-half-maximum (FWHM) of the Se–Se A' mode of kesterite phase (peak located at 196 cm^{-1}) from Mg-containing/Mg-free samples at surface/sub-surface regions show that the differences are within the experimental error. The apparent FWHM increases between the surface and the sub-surface measurements (around 2 cm^{-1}) (see Tables S1, S2 in Supporting Information) is attributed to the activation by Raman pre-resonant process of weak contributions overlapped with the 196 cm^{-1} peak. The absence of significant variations of the characteristic properties of the Se–Se peak suggests that both samples (Mg-containing/Mg-free) present a similar crystalline quality at the surface and sub-surface regions.

For the Mg-containing sample C an extended analysis has been performed to compare the surface/bulk and bulk regions. The results show a red-shift of the main peak as well as an increase of the FWHM. These changes are an indication of crystal quality degradation, in agreement with the smaller grains at the back of the absorber layer observed by cross-sectional SEM images (see **Figure 2B**). Additionally, this red-shift of the main peak could be compatible with the replacement of Cu or Zn by a lighter atom as Mg. As shown in **Figure 2A**, a minimum Cu SIMS-signal coincides with a maximum Mg SIMS-signal and is located near the border between the big and small grains shown in **Figure 2B**. These results seem to suggest that a replacement of Cu by Mg could take place in that region, which could reduce the Cu_{Zn} antisite defects formation.

The analysis of the changes at 175 and 242 cm^{-1} regions is interesting in order to evaluate the concentration variation of the defect clusters $[\text{Zn}_{\text{Cu}}+\text{V}_{\text{Cu}}]$ and $[2\text{Zn}_{\text{Cu}}+\text{Zn}_{\text{Sn}}]$ (Dimitrievska et al., 2015). A significant variation of the 175 cm^{-1} peak intensity between the samples C and D at the surface and the sub-surface is observed. The 175 and 196 cm^{-1} peak intensity ratio has been correlated with the concentration of the $[\text{Zn}_{\text{Cu}}+\text{V}_{\text{Cu}}]$ (A-type) defect cluster and with the V_{oc} of the device (Dimitrievska et al., 2015). For the Mg-containing samples an increase of this ratio is observed. This suggests that with the inclusion of Mg, a reduction of V_{Cu} and/or Zn_{Cu} antisite formation is promoted. It is in agreement with the hypothesis that for low Mg concentration in CZTSSe samples, Mg replaces Cu and/or the Mg allows the inhibition of Zn_{Cu} antisites formation during CZTSSe synthesis (Zhong et al., 2016). Additionally the analyses of the bulk region show similar values of the 175 cm^{-1} peak intensity than those observed for the surface indicating a similar defect property at the surface and bulk of the absorber layer. Furthermore, the variation of the PV parameters of both samples with the $A[175\text{ cm}^{-1}]/([A175\text{ cm}^{-1}]+A[196\text{ cm}^{-1}])$ ratio (see **Figure 5**) shows a clear correlation between the improvement of V_{oc} , short circuit current density J_{sc} , fill factor FF and efficiency η with the increase of the $[\text{Zn}_{\text{Cu}}+\text{V}_{\text{Cu}}]$ defect cluster concentration (reduction of the $A[175\text{ cm}^{-1}]/([A175\text{ cm}^{-1}]+A[196\text{ cm}^{-1}])$ ratio), similarly as what has already been reported in previous works for pure CZTSe (Dimitrievska et al., 2016a,b). Therefore, optoelectronic parameters increase with the increase of the defect cluster concentrations as it is shown in **Figure 5** for both samples, Mg-containing and Mg-free. This fact suggests that Mg improves material properties allowing better device



performances, but reduces the $[\text{Zn}_{\text{Cu}}+\text{V}_{\text{Cu}}]$ defect clusters density.

The evolution of the peak intensity at 242 cm^{-1} shows that for the Mg-containing and Mg-free samples, the surface and sub-surface regions are similar while a clear increase is observed for the spectra acquired from the points (1) and (2) of the lift-off for sample C, regions with smaller grain size. This suggests that for this region, the $[2\text{Zn}_{\text{Cu}}+\text{Zn}_{\text{Sn}}]$ defect cluster is promoted inducing the degradation of the crystal quality observed by changes of the FWHM and Raman shift (Dimitrievska et al., 2015).

An average efficiency of 5.8%, $V_{\text{oc}} = 405\text{ mV}$, $J_{\text{sc}} = 36.3\text{ mA cm}^{-2}$ and $\text{FF} = 39.2\%$ was measured for nine Mg-containing solar cells on sample C, whereas the nominally Mg-free sample D yielded an average efficiency of 4.2%, $V_{\text{oc}} = 376\text{ mV}$, $J_{\text{sc}} = 29.0\text{ mA cm}^{-2}$ and $\text{FF} = 38.2\%$ for nine cells. J-V

curves and EQE spectra for best cells from sample C and D are shown in **Figure 6**. It appears that the performance of Mg-containing cells is higher than that for nominally Mg-free devices, in particular by improving the charge collection in the long-wavelength part of the spectrum (**Figure 6B**). However, no reduction of the V_{oc} deficit is measured after adding Mg, being of 0.59 V for both best cells. The band gap energy E_g determined from the inflection point of the EQE spectrum is 1.01 eV and 0.97 eV for samples C and D, respectively. The

comparison of reverse-biased EQE measurement with the EQE without bias is also plotted in **Figure 6B** for the device C. Although the Mg-containing device exhibits an enhanced carrier collection, it can be further improved under negative bias (-1 V). This bias dependence is signature of a poor carrier collection toward the back of the absorber layer (Scheer and Schock, 2011), which can be related to the smaller grain size next to the back contact.

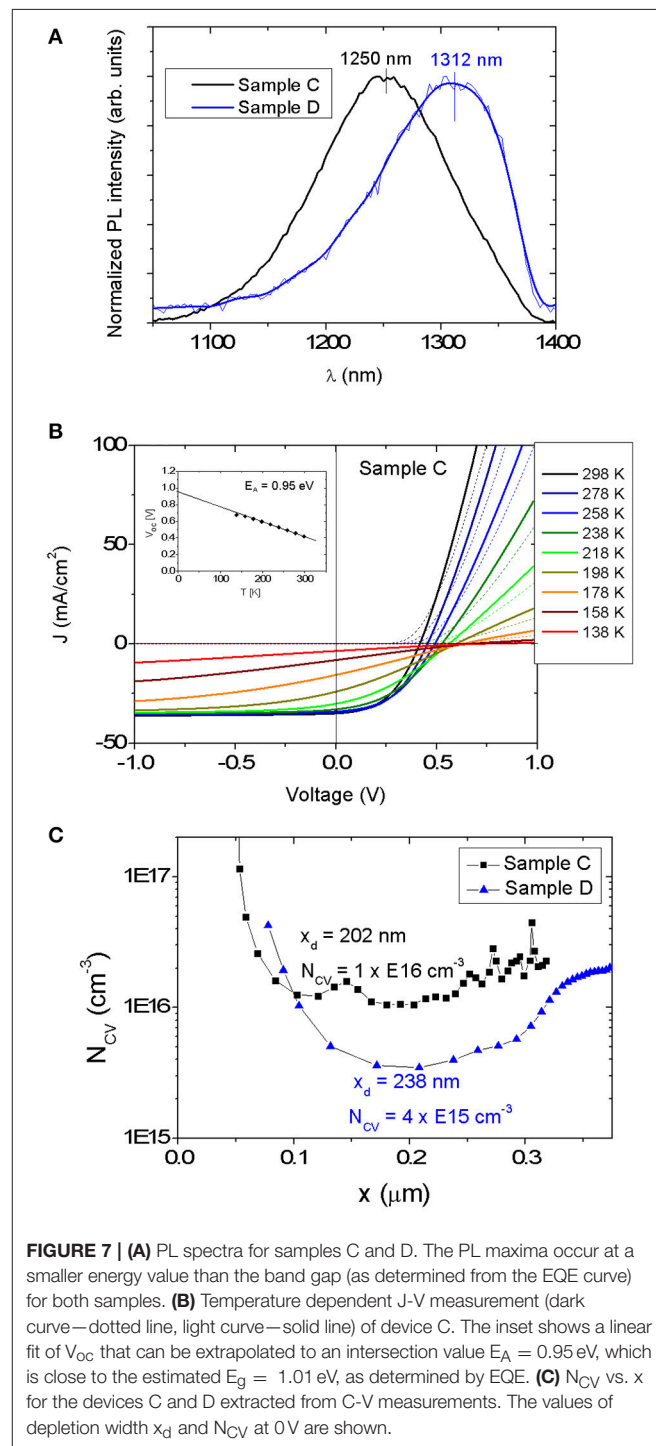
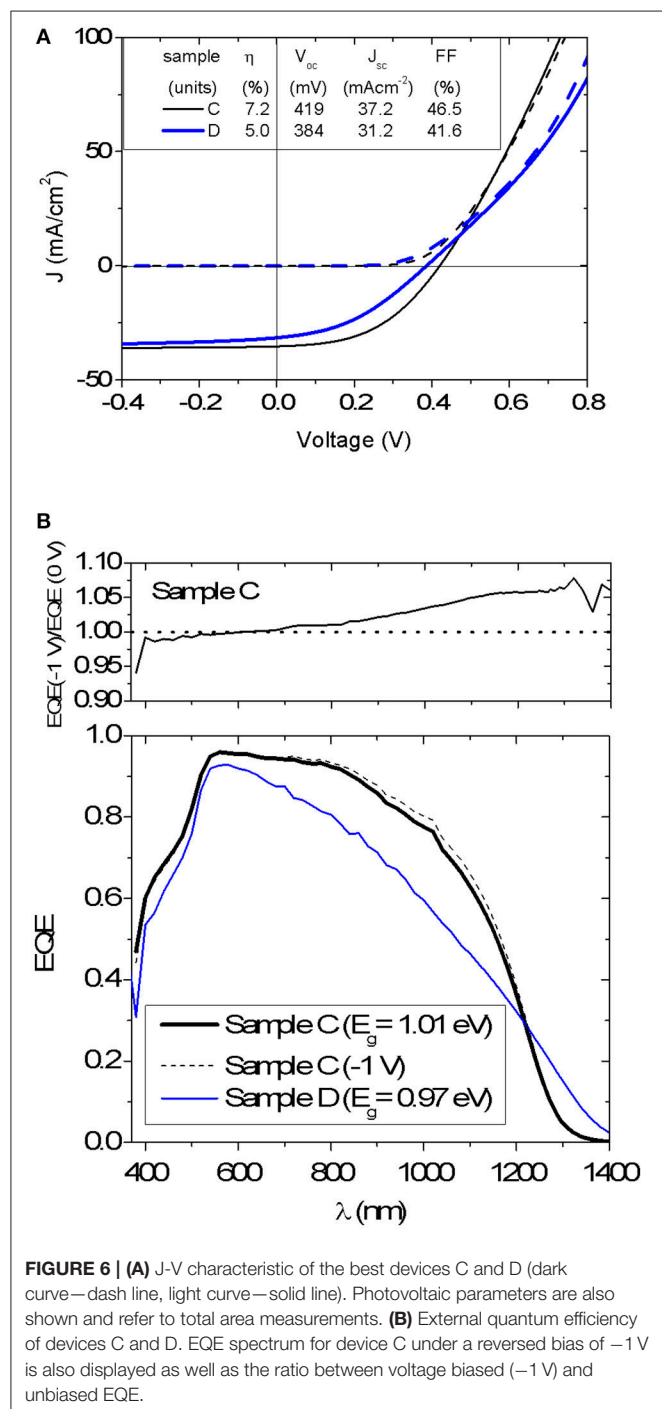


Figure 7A shows PL spectra acquired at room temperature of devices C and D. The PL spectra exhibit a broad peak for both samples. The broad shape indicates the presence of tail states and/or potential fluctuations (Gokmen et al., 2013). Red-shifted PL peak maxima compared to E_g , ΔE_{Eg-PL} , by 20 meV and 25 meV are determined for solar cells C and D, respectively. These values are in the range obtained for CZTSe grown by vacuum-based methods (20 meV) (Lee et al., 2015).

Figure 7B displays the JV-T measurement of the device C. No roll-over effect is observed even at the lowest temperature of 138 K. The temperature dependence of the V_{oc} extrapolated to $T = 0$ K provides an activation energy $E_A = 0.95$ eV, which can correspond to the main recombination channel. Since this E_A value is very near the E_g of 1.01 eV determined by EQE, it appears that the dominant recombination path is located within the bulk of the kesterite absorber layer rather than at the CdS/CZTSSe interface.

Furthermore, capacitance-voltage (C-V) measurement was carried out for both solar cells (see **Figure 7C**). The depletion width (x_d) of the Mg-containing CZTSSe device is approximately of 0.2 μm at a bias of 0 V. A carrier concentration N_{CV} of $1 \times 10^{16} \text{ cm}^{-3}$ can be extracted for the Mg-containing absorber, whereas a lower carrier concentration N_{CV} of $4 \times 10^{15} \text{ cm}^{-3}$ was measured for the Mg-free sample, indicating that the presence of Mg increases acceptor density. The doping level is in the same range as for Na-doped CZTSe solar cells fabricated using a NaF precursor layer and by diffusion from SLG substrates ($x_d = 0.23 \mu\text{m}$ and $N_{CV} = 1 \times 10^{16} \text{ cm}^{-3}$) (Lee et al., 2015).

CONCLUSIONS

The incorporation of Mg into $\text{Cu}_2\text{Zn}_{1-x}\text{Mg}_x\text{Sn}(\text{S},\text{Se})_4$ layers leads to the complete phase separation for $x = 0.5 \dots 1$, whereas a lower content of $x = 0.04$ can preserve the kesterite phase. As compared to the nominally Mg-free sample, low quantities

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of Mg can improve the grain growth, reduce the number of structural defects and increase the acceptor density in the solar cell absorber. In this respect Mg appears to behave similar to the conventional alkali dopants such as Na or K. The first Mg-containing kesterite cell has been fabricated with the highest conversion efficiency of 7.2%. This value is lower than the efficiency of 11–12% obtained for Li, K or Na-doped devices (Haass et al., 2017), indicating that the addition of Mg is less effective.

AUTHOR CONTRIBUTIONS

RC, SH, and YR designed the research and experiments. RC and SH fabricated solar cells, characterized layers and solar cells. CA assisted with the analysis. LA and FO carried out Raman spectroscopy measurements. VI-R assisted with Raman spectra analysis. RC, SH, FO, and YR wrote the paper. All authors contributed with discussions.

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SUPPLEMENTARY MATERIAL

The Supplementary Material for this article can be found online at: <https://www.frontiersin.org/articles/10.3389/fchem.2018.00005/full#supplementary-material>

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Conflict of Interest Statement: The authors declare that the research was conducted in the absence of any commercial or financial relationships that could be construed as a potential conflict of interest.

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