frontiers in Astronomy and Space Sciences

CORRECTION
published: 25 October 2019
doi: 10.3389/fspas.2019.00067

Corrigendum: On the Nature of the Core of α Centauri A: The Impact of the Metallicity Mixture

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Edited by:
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Reviewed by:
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Specialty section:
This article was submitted to Stellar and Solar Physics, a section of the journal Frontiers in Astronomy and Space Sciences

Received: 02 July 2019
Accepted: 04 October 2019
Published: 25 October 2019

Citation:
Nsamba B, Campante TL, Monteiro MJPFG, Cunha MS and Sousa SG (2019) Corrigendum: On the Nature of the Core of α Centauri A: The Impact of the Metallicity Mixture. Front. Astron. Space Sci. 6:67. doi: 10.3389/fspas.2019.00067

Keywords: HD 128620, asteroseismology, stellar modeling, fundamental parameters, convection, radiation

A Corrigendum on

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by Nsamba, B., Campante, T. L., Monteiro, M. J. P. F. G., Cunha, M. S., and Sousa, S. G. (2019). Front. Astron. Space Sci. 6:25. doi: 10.3389/fspas.2019.00025

In the original article, there was a mistake in **Figure 3** as published.

This figure aimed at exploring trends regarding the contribution from the dominant model properties (i.e., mass, initial metal mass fraction, and overshoot parameter) toward the convective core sizes of the best-fit models of α Centauri A. However, the data plotted in **Figure 3** in the original article represents the models with convective cores that satisfy a probability threshold required to fit the sets of observables (i.e., seismic observables, classical observables, and an interferometric radius), instead of showing the best-fit models with convective cores that satisfy the sets of observables within 1σ. The corrected **Figure 3** appears below.

The authors apologize for this error and state that this does not change the scientific conclusions of the article in any way. The original article has been updated.

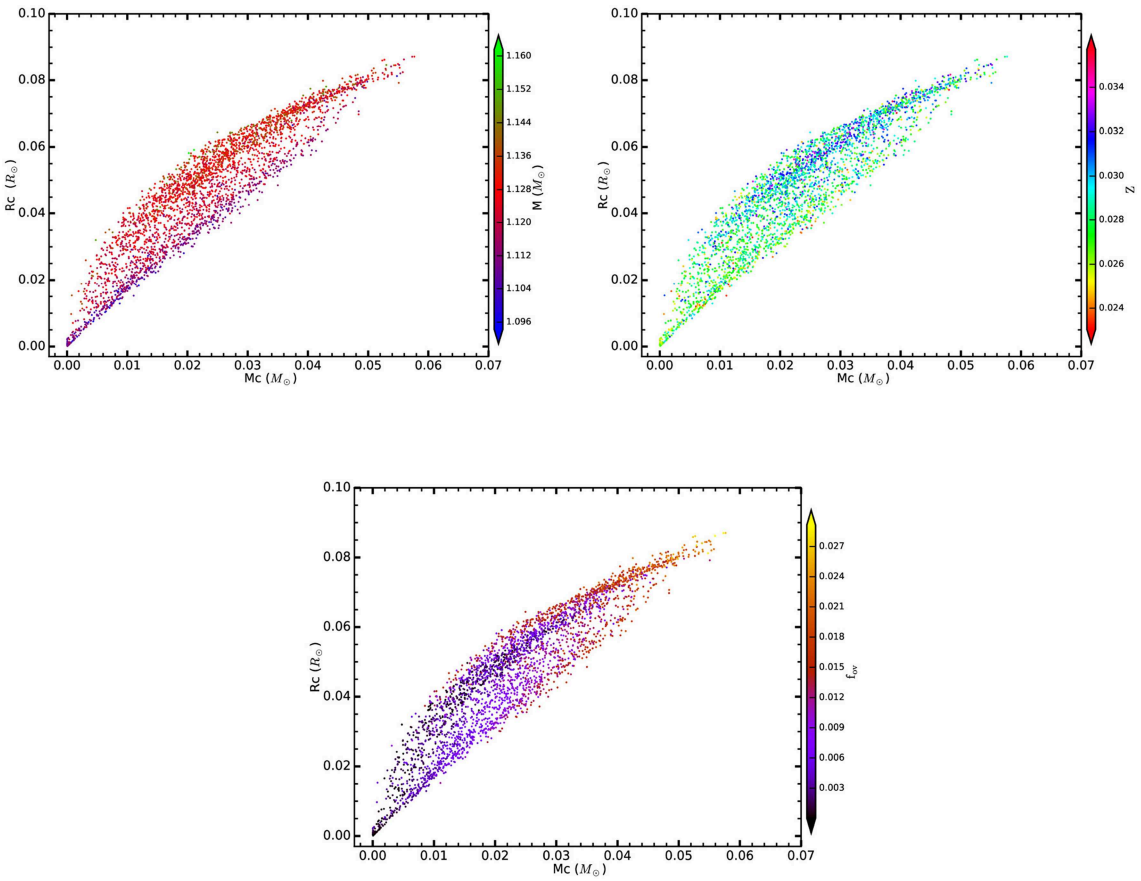

FIGURE 3 | Scatter plots showing core radius (R_c) vs core mass (M_c) for best-fit models with convective cores obtained using grid B (Run I). Top **left** is color-coded according to the model mass (M), top **right** is color-coded according to the initial metal mass fraction (Z), and bottom panel is color-coded according to the overshoot parameter (f_{ov}).