



A Fresh Look at Waves in Ion-Electron Plasmas

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Exploiting the general dispersion relation describing all waves in an ideal ion-electron fluid, we revisit established treatments on wave families in a cold ion-electron plasma. These contain the magnetohydrodynamic Alfvén and fast waves at low frequencies, long wavelengths, but are enriched by short wavelength resonance behaviors, electrostatic and electromagnetic mode types, and cut-off frequencies distinguishing propagating from evanescent waves. Our theoretical treatment exploits purely polynomial expressions, which for the cold ion-electron case only depend on 2 parameters: the ratio of masses over charges μ and the ratio E of the electron gyro frequency to the combined ion-electron plasma frequency. We provide a complete description of all waves, which stresses the intricate variation of all five branches of eigenfrequencies $\omega(k, \vartheta)$ depending on wavenumber k and angle ϑ between wavevector and magnetic field B . Corresponding 5-mode phase and group diagrams provide insight on wave transformations and energy transport. Special cases, like the high frequency modes in magneto-ionic theory following from Appleton-Hartree dispersion relations, are naturally recovered and critically discussed. Faraday rotation for electromagnetic waves is extended to all propagation angles ϑ . The discussion covers all cold ion-electron plasma waves, up into the relativistic regime.

Keywords: waves, cold plasmas, 2-fluid theory, magnetohydrodynamic, electromagnetic wave theory

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1. INTRODUCTION

The theory of wave propagation in ion-electron plasmas is covered in many textbooks (Stix, 1992; Boyd and Sanderson, 2003; Bittencourt, 2004; Chen, 2016; Thorne and Blandford, 2017), and can be considered established. The starting point for many treatments is based on a 2-fluid approach, where one solves for plane wave solutions $\exp[i(\mathbf{k} \cdot \mathbf{x} - \omega t)]$ in an otherwise homogeneous medium, usually magnetized with uniform magnetic field \mathbf{B} . In the rest frame of a homogeneous ion-electron mixture, assumed to be charge-neutral such that number densities obey $n_e = Zn_i$ when ions have charge number Z , the equilibrium electric field and current vanish, while each species has its own pressure p_e, p_i . The dispersion relation between wave frequency ω and wavenumber $k = |\mathbf{k}|$ is then usually obtained from linearizing the Maxwell equations, traditionally introducing dielectric, susceptibility and conductivity tensors, to quantify displacement vector, polarization vector and current vector relations to the electric field, respectively. In general, this leaves a large variety of wave modes that are particularly aware of the orientation angle ϑ between wavevector \mathbf{k} and the magnetic field \mathbf{B} , while the wave properties can differ greatly according to frequency and wavelength. Indeed, when we have electron and ion masses given by m_e and m_i , the plasma is characterized by its ratio of masses over charges $\mu = Zm_e/m_i$, and this background

ion-electron mixture introduces its own set of typical frequencies and wavelengths. The former include the electron and ion plasma frequencies $\omega_{pe} = \sqrt{e^2 n_e / \epsilon_0 m_e}$ and $\omega_{pi} = \sqrt{\mu} \omega_{pe}$ and the electron and ion cyclotron frequencies $\Omega_e = eB/m_e$ (where electron charge is $-e$) and $\Omega_i = \mu \Omega_e$. Lengthscales relate to typical speeds in the system, such as the light speed $c = 1/\sqrt{\mu_0 \epsilon_0}$ and the electron and ion sound speeds. One such lengthscale is the skin depth $\delta = c/\omega_p$, where it turns out convenient to combine the plasma frequencies into $\omega_p^2 = \omega_{pe}^2 + \omega_{pi}^2$. The governing dispersion relation equally follows (e.g., Goedbloed and Poedts, 2004) from a standard linearization of the governing equations of motion for each species, combined with Maxwell equations. We will in what follows rewrite and analyse this general form of the dispersion relation, but specify deliberately to the cold ion-electron case, where the species pressures (and sound speeds) vanish $p_e = 0 = p_i$. We can normalize all frequencies to the plasma frequency $\bar{\omega} = \omega/\omega_p$, and for such cold ion-electron plasma, only two dimensionless parameters remain in the description, namely

$$E = \Omega_e/\omega_p, \quad I = \Omega_i/\omega_p. \quad (1)$$

That only two parameters cover the full complexity of wave modes in a cold ion-electron plasma is well-known, and is at the basis of the classical Clemmow-Mullaly-Allis plots that feature in most textbooks (Stix, 1992; Bittencourt, 2004; Thorne and Blandford, 2017), which classify wave modes and associated wave normal surfaces (plots of ω/k vs. ϑ , which are figures of revolution about \mathbf{B}). In what follows, we will discuss all wave modes and suggest a new classification scheme based on the polynomial form of the governing dispersion relation. That the ideal two-fluid description leads to a 12th order polynomial in the wave frequency ω is well-known (Denisse and Delcroix, 1961), although one frequently exploits lower order polynomials to cover e.g., only the high frequency electromagnetic waves (following e.g., the Appleton-Hartree description), or conversely focuses on the three low frequency branches that relate to the magnetohydrodynamic (MHD) slow, Alfvén and fast mode pairs (Stringer, 1963; Ishida et al., 2005; Damiano et al., 2009; Bellan, 2012; Zhao, 2015). An eight order polynomial approximation, valid above the lower hybrid frequency, was exploited in Zhao (2017) to cover whistler waves and the three high frequency electromagnetic mode pairs. We will instead start from the full 12th order equation, only making the cold plasma assumption, such that the slow MHD modes become marginal solutions $\omega^2 = 0$. We make contact with the mentioned textbook treatments, showing how the usual dielectric tensor treatments are indeed mathematically fully equivalent. The advantage of the polynomial method over the dielectric tensor method is the direct relationship of the solutions of the dispersion equation to the primitive two-fluid variables, which would also be exploited in corresponding numerical time stepping codes. This advantage was also pointed out by Bellan (2012), who gave a similar analysis for the low frequency domain. The present paper exploits the polynomial method to full effect to provide the crucial phase and groups diagrams for all values of the parameters, together with their animations.

2. DISPERSION RELATION FOR COLD ION-ELECTRON PLASMAS

Following Goedbloed and Poedts (2004), the general dispersion relation for an ideal ion-electron fluid can be written as a twelfth-order polynomial in $\bar{\omega}$, where one distinguishes six pairs of forward and backward propagating modes, since the expression is actually sixth order in $\bar{\omega}^2$. This already eliminated a pair $\bar{\omega}^2 = 0$ of marginal entropy-like modes, and the special case of a cold ion-electron fluid can factor out another $\bar{\omega}^2 = 0$ solution, corresponding to the slow magnetohydrodynamic (backward and forward) waves. It is a matter of algebra to show that the remaining 10th order polynomial can be rewritten to

$$\bar{k}^4 \bar{A}(\bar{\omega}^2, \lambda^2) - \bar{k}^2 \bar{\omega}^2 \bar{B}(\bar{\omega}^2, \lambda^2) + \bar{\omega}^4 \bar{C}(\bar{\omega}^2) = 0, \quad (2)$$

where $\bar{k} = \delta k$, $\lambda^2 = \cos^2 \vartheta$, containing three 3rd order polynomials in $\bar{\omega}^2$ given by

$$\begin{aligned} \bar{A} &= \bar{\omega}^6 - (1 + E^2 + I^2) \bar{\omega}^4 \\ &+ [(1 + EI)EI + \lambda^2(E^2 + I^2 - EI)] \bar{\omega}^2 - \lambda^2(EI)^2, \end{aligned} \quad (3)$$

$$\begin{aligned} \bar{B} &= 2\bar{\omega}^6 - (4 + 2E^2 + 2I^2) \bar{\omega}^4 + [2(1 + EI)^2 \\ &+ (1 + \lambda^2)(E^2 + I^2 - EI)] \bar{\omega}^2 - EI(1 + EI)(1 + \lambda^2), \end{aligned} \quad (4)$$

$$\bar{C} = (\bar{\omega}^2 - 1) [\bar{\omega}^4 - (2 + E^2 + I^2) \bar{\omega}^2 + (1 + EI)^2]. \quad (5)$$

Unlike most textbook treatments, which take Equation (2) at fixed frequency and solve for both roots in \bar{k}^2 , we will consider a given wavenumber \bar{k} , and use the dispersion relation to quantify all five roots in $\bar{\omega}^2$. In what follows, we will also systematically drop the overbars on \bar{k} and $\bar{\omega}$, as we will always work with dimensionless frequencies and wavenumbers.

2.1. Cut-Off Frequencies

Cut-off frequencies relate to large wavelength (small wavenumber k) limiting behavior, and Equation (2) reveals instantly that special frequencies then follow from the zeros of the \bar{C} polynomial. These are computed simply as

$$\omega_c^2 = 1, \quad (6)$$

$$\omega_{u,l}^2 = 1 + \frac{E^2 + I^2}{2} \pm |E - I| \sqrt{\frac{(E + I)^2}{4} + 1}. \quad (7)$$

Hence, the $\omega_c^2 = 1$ cutoff happens exactly at the combined plasma frequency (this reads as $\omega^2 = \omega_p^2$ when restoring dimensions). The latter pair $\omega_{u,l}^2$ distinguishes the upper (+) from the lower (−) cut-off frequency. For $E = 0$, the three cut-offs coincide at the plasma frequency (noting that $I = \mu E$).

In general, the three cut-off frequencies are independent of the angle ϑ , but their relative ordering is influenced by the values of both parameters E, I or equivalently, E and μ . Since μ is normally considered fixed in a specific plasma (e.g., it assumes the value $\mu \approx 1/1,836$ in a fully ionized hydrogen plasma), the magnetic field strength in essence determines the other parameter E , going from unmagnetized $E = 0$ cases, to strongly magnetized regimes

$E \rightarrow \infty$. A special case is obtained for an electron-positron or pair plasma where $\mu = 1$, when $\omega_{u,l}^2$ coincide to $\omega_I^2 = \omega_u^2 = 1 + E^2$ (Keppens and Goedbloed, in press). For an arbitrary μ , it is clear that $\omega_u^2 \geq 1$ (and the equality holds for the unmagnetized case $E = 0$), but the lower cut-off may be below, equal to, or above unity. The equality of ω_c^2 with ω_I^2 corresponds to the specific field strength where $E = 1/\mu - 1$. A plot of the three cut-off frequencies for a $\mu = 1/1,836$ hydrogen plasma vs. E is shown in **Figure 1**. The vertical dashed lines indicate special values for E , e.g., where $E = 1/\mu - 1$, where the ordering of the three cut-offs changes. The right panel in **Figure 1** shows a zoom on the behavior near this value. The above exact expressions for the cut-off frequencies, valid for all values of E , can also be used to get approximate expressions in limits of small or large E regimes. Obviously, the small E limit makes all cut-offs approach their unmagnetized regime: all 3 become the plasma frequency. The opposite, large E limit is easily evaluated to give as upper and lower cut-off (both way above the third cut-off at plasma frequency):

$$\lim_{E \rightarrow \infty} \omega_u^2 = E^2, \quad (8)$$

$$\lim_{E \rightarrow \infty} \omega_l^2 = E^2 \mu^2. \quad (9)$$

2.2. Resonances

Resonances occur at large wavenumbers (short wavelengths), and Equation (2) shows that then the zeros of \bar{A} come into play. The three resonances obtained from $\bar{A} = 0$ can be computed using standard root-finding for polynomial expressions (or, could be explicitly obtained through Cardano's formulae for the roots of a third order polynomial). For a given set of parameters (E, I) or (E, μ), the 3 resonances still depend on the orientation angle ϑ . The exact variation of the three resonances at any (ϑ, E, μ) requires to compute the three roots from

$$\omega^6 - [1 + E^2(1 + \mu^2)]\omega^4 + E^2[\mu + \mu^2 E^2 + \lambda^2(1 - \mu + \mu^2)]\omega^2 - \lambda^2 E^4 \mu^2 = 0. \quad (10)$$

For each of the three resonances, we typically find monotonic behavior $\omega(\vartheta)$ between the extremal angles of parallel to perpendicular orientation. **Figure 1** also shows the range in values obtained for the three resonances, for the same fixed value of $\mu = 1/1,836$, as function of the remaining E parameter. Three different colors indicate the three resonance ranges. The top (purple) one is seen to stay above unity, and stays between the upper cut-off and E^2 for large E . This purple resonance range always shows increasing frequencies when going from parallel (solid) to perpendicular (dashed) behavior. The middle (blue) range is bounded from above by E^2 , and stays below unity until we reach the $E = \sqrt{\mu^2 - \mu + 1}/\mu$ value, which lies in between the dotted lines shown at $E = 1/\mu - 1$ and $E = 1/\mu$ (right panel). Ultimately, its range becomes bounded by the $\mu^2 E^2$ curve. This blue resonance will relate to the fast magnetosonic wave branch. The lowest resonance range (red) always stays below the $\mu^2 E^2$ line and below unity. In fact, this resonance range extends all the way to zero at exactly perpendicular orientation, and will be shown to relate to the Alfvén branch.

We can get approximate expressions for the three resonances by setting $\mu = 0$ (which is unphysical, but is at the basis of the often used Appleton-Hartree dispersion relation, see section 5.3). We then obtain one solution at zero frequency (which will relate to the red or Alfvén branch), and the purple and blue resonances become

$$\omega^2 \approx \frac{1}{2} [1 + E^2 \pm \sqrt{(1 + E^2)^2 - 4E^2 \lambda^2}]. \quad (11)$$

Note that setting $\mu = 0$ eliminates the (normalized) ion cyclotron frequency I from the description.

The plots for the general $\mu \neq 0$ case in **Figure 1** reveal how the red resonance range actually always decreases in frequency, with angle going from parallel to perpendicular, while the purple one is always increasing in frequency, going from parallel to perpendicular orientations. The blue resonance range on the other hand first decreases in frequency, up to the value $E = \sqrt{\mu^2 - \mu + 1}/\mu$. From then on, the blue range increases in frequency from parallel to perpendicular. The dashed (perpendicular) and solid (parallel) limits are easily obtained analytically, since

$$\bar{A}(\lambda = 0) = \omega^2 [\omega^4 - (1 + E^2 + I^2)\omega^2 + (1 + EI)EI], \quad (12)$$

$$\bar{A}(\lambda = 1) = (\omega^2 - 1) [\omega^4 - (E^2 + I^2)\omega^2 + (EI)^2]. \quad (13)$$

The solutions to the perpendicular ($\lambda = 0$) case are thus threefold: one at zero frequency, related to non-propagation of Alfvén waves perpendicular to the magnetic field, the other two solutions from $\omega^4 - (1 + E^2 + I^2)\omega^2 + (1 + EI)EI = 0$ are known as the upper and lower hybrid resonance, given by our purple and blue dashed lines in **Figure 1**, respectively. These upper and lower hybrid resonances are thus generally given by

$$\omega_{\pm}^2 = \frac{1}{2} \left[1 + E^2 + I^2 \pm \sqrt{(1 + E^2 + I^2)^2 - 4(1 + EI)EI} \right]. \quad (14)$$

Their limits for low magnetization ($E \ll 1$) become 1 and 0, as the electrostatic mode remains and the fast mode becomes marginal (see further). In the limit of high magnetization ($E \gg 1$), we find that $\omega_{+}^2 \approx E^2$ while $\omega_{-}^2 \approx \mu^2 E^2 = I^2$.

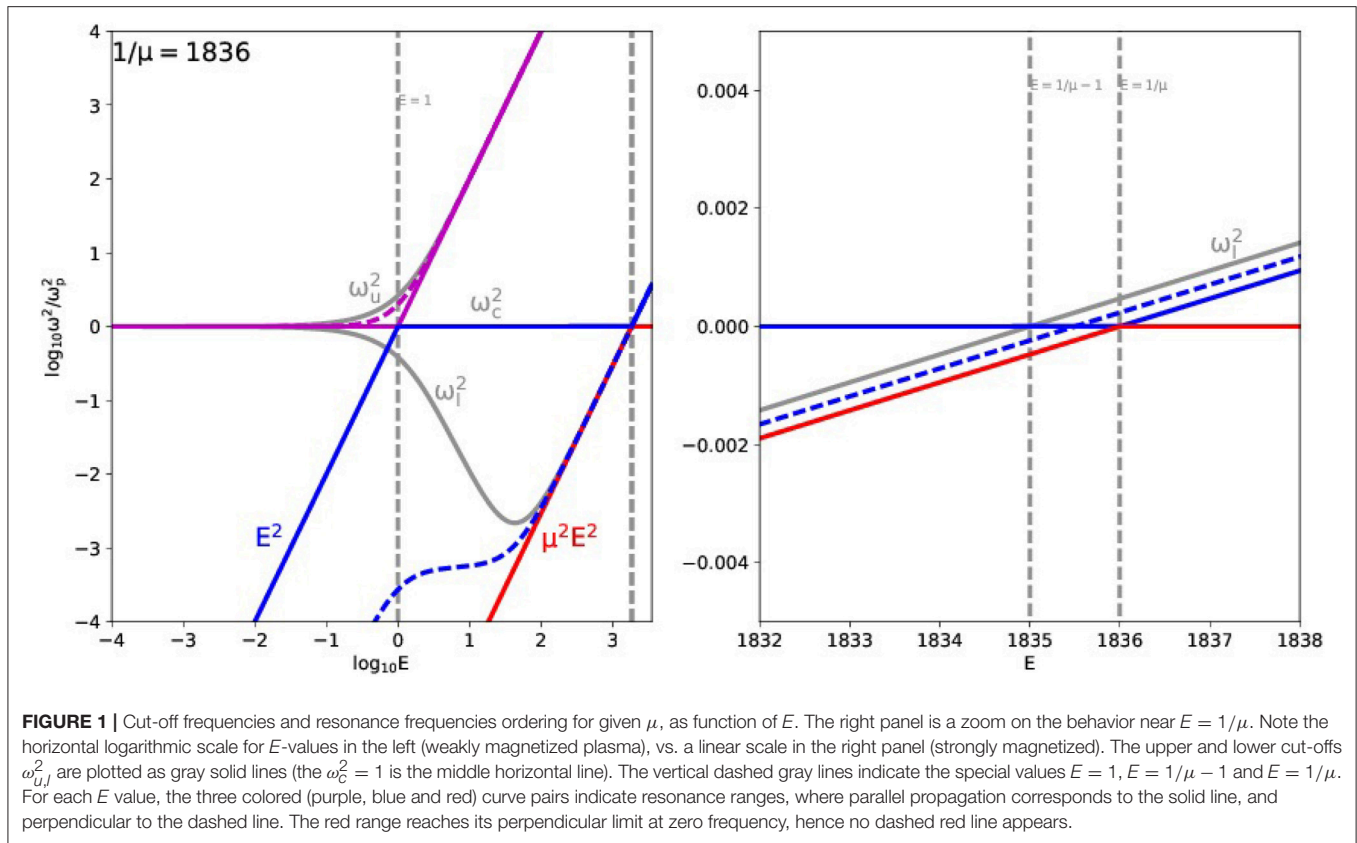
It can be noted that for the pair plasma case where $\mu = 1$, the above discussion of resonances and cut-offs simplifies significantly: e.g., the blue resonance range collapses on the curve E^2 . Also, for a pair plasma, the purple resonance range extends (at exactly perpendicular propagation) to the upper cut-off value ω_u^2 (Keppens and Goedbloed, in press).

2.3. Low and High Frequency Limits

The dispersion relation Equation (2) also shows clearly the limits at both high and low frequencies, where the plane wave phase speed $v_{ph} = \omega/k$ attains a finite value. At high frequencies, we find that we obtain a double-valued solution at light speed behavior, as we can instantly write this limit as

$$(1 - v_{ph}^2)^2 = 0. \quad (15)$$

Hence, we expect two solutions that will behave as electromagnetic waves at high frequencies. These high frequency solutions, together with a finite phase speed, require large



wavenumbers. Hence, together with the three resonances that occur at large wavenumbers, these two will serve to organize the 5 solution branches.

At low frequencies, we similarly retrieve the following solutions

$$v_{ph,F}^2 = \frac{EI}{1 + EI}, \tag{16}$$

$$v_{ph,A}^2 = \lambda^2 \frac{EI}{1 + EI}. \tag{17}$$

These are actually fast (F) and Alfvén (A) waves, at least at large wavelengths (required to make the phase speed finite at low frequencies). Hence, they will augment the 3 cut-offs to organize the 5 solution branches. The combination $EI/(1 + EI)$ can be recognized as the relativistically correct expression for the Alfvén speed v_A . This speed is more commonly written in terms of the magnetization parameter $\sigma = B^2/(\mu_0 c^2 \rho h)$ where $\rho = n_e m_e + n_i m_i$ and the specific enthalpy $h = 1$ for a cold plasma, and we get (temporarily restoring dimensions):

$$\frac{v_A^2}{c^2} = \frac{EI}{1 + EI} = \frac{\sigma}{1 + \sigma} = \frac{B^2}{\rho h \mu_0 c^2 + B^2}. \tag{18}$$

3. DISPERSION DIAGRAMS

We will now analyse dispersion diagrams, which are obtained as solutions $\omega^2(k, \vartheta)$ to the general dispersion relation. We first fix a

parameter set (E, μ) . Then, at fixed wavenumber and orientation, Equation (2) is a polynomial of degree 5 in ω^2 , and due to the symmetry of the underlying determinant (Goedbloed and Poedts, 2004), there will always be five real solutions. We will first address the special orientations of parallel and perpendicular propagation. We plot in **Figure 2** two dispersion diagrams, one for parallel (left) and one for perpendicular (right) orientation, for the choice where $\mu = 1/1,836$ and $E = 1.5$.

3.1. Parallel Propagation

A first observation we can directly make from Equation (2) is that parallel propagation ($\lambda = 1$) can always factor out the solution $\omega^2 = 1$, since $\bar{A}(\omega^2 = 1, \lambda = 1) = 0 = \bar{B}(\omega^2 = 1, \lambda = 1)$. This solution is also retained at all angles when $E = 0$, where it actually represents the electrostatic mode, which is a non-propagating plasma oscillation. We will find that $E \neq 0$ cases will turn this mode into a propagating solution, for all angles away from parallel.

The other 4 pairs in Equation (2) for $\lambda = 1$ are intricately mixed, but it is possible to write the remaining 8th order expression as

$$\begin{aligned} & [\omega^4 + \omega^3|E - I| - \omega^2(1 + EI + k^2) - \omega|E - I|k^2 + k^2EI] \\ & \times [\omega^4 - \omega^3|E - I| - \omega^2(1 + EI + k^2) + \omega|E - I|k^2 + k^2EI] = 0. \end{aligned} \tag{19}$$

Note that $|E - I| = (1 - \mu)E$ as $\mu \leq 1$, and in the case of a pair plasma ($\mu = 1$), this leads to a full factorization

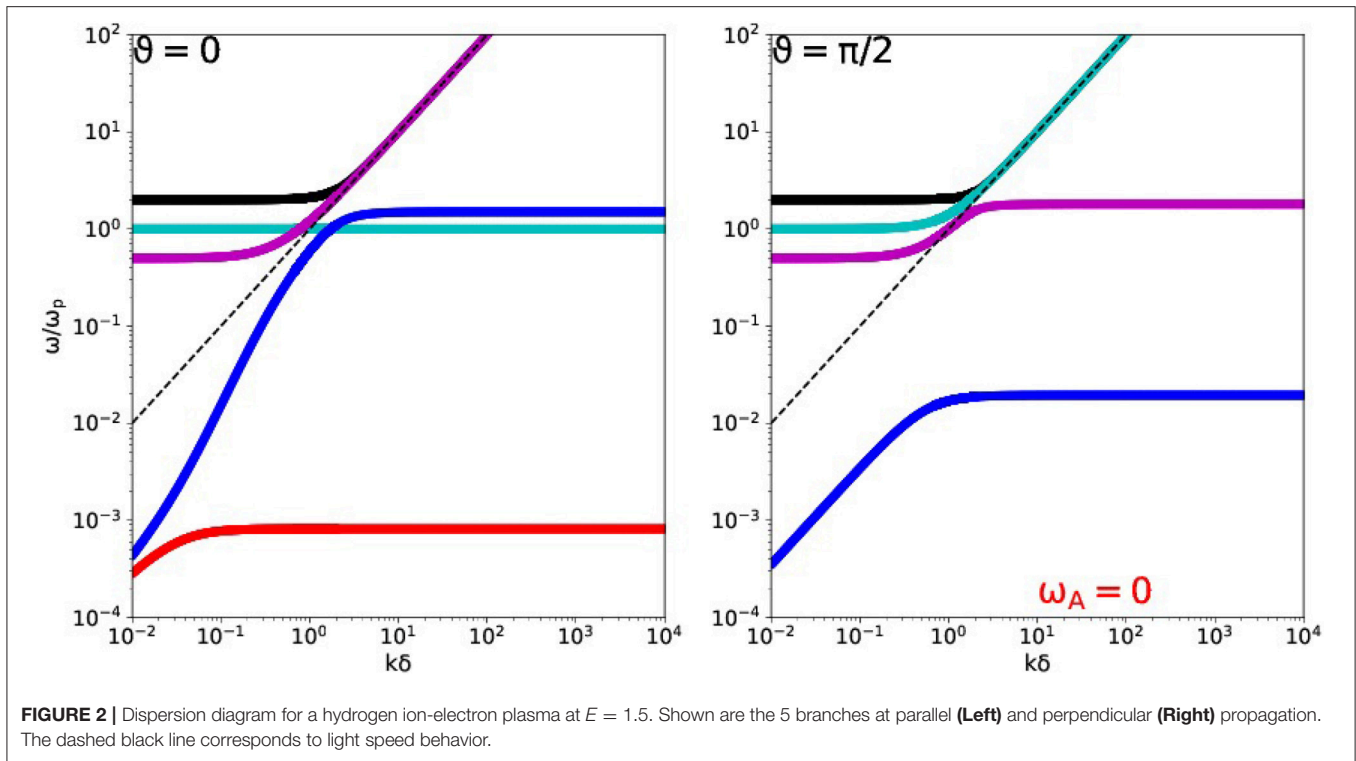


FIGURE 2 | Dispersion diagram for a hydrogen ion-electron plasma at $E = 1.5$. Shown are the 5 branches at parallel (Left) and perpendicular (Right) propagation. The dashed black line corresponds to light speed behavior.

of all solution branches (Keppens and Goedbloed, in press). However, for $\mu \neq 1$, Equation (19) does not seem to offer real advantage, since the original formulation was already a 4th order polynomial in ω^2 , and its four solution pairs ω_1^2 , ω_2^2 , ω_3^2 and ω_4^2 retain their meaning as forward and backward propagating mode pairs. Instead, both factors in Equation (19) mix forward and backward pairs: if ω_i is solution to its first factor, $-\omega_i$ will be a solution to its second factor. Still, many textbook treatments in fact rely on the factorization in Equation (19). This, as we will discuss further on, implies a description in left and right circularly polarized waves, and indeed left and right circular polarizations switch roles when going forward or backward in time, or under mirror symmetry (see also Keppens and Demaerel, 2016). It is important to note that the forward-backward means of categorizing the solutions is fully consistent with combined PT symmetry, i.e., flipping the time direction and mirroring space.

We rather use the original Equation (2) to compute at which k -values any of the other 4 branches intersects with the $\omega^2 = 1$ solution. This will prove useful when we vary the orientation angle ϑ away from parallel, as we will show further on: these k -values are special as they correspond to those locations where we will first witness avoided crossings of branches. This turns out to be at values

$$k^2 = \frac{EI \pm |E - I|}{EI - 1 \pm |E - I|}. \tag{20}$$

Of course, only positive values need to be considered for k^2 . A careful analysis of Equation (20) shows that only one sign combination is positive as long as $E < 1$, and hence in such cases

only one other branch will intersect the $\omega^2 = 1$ range (e.g., see **Figure 2** for $E = 0.5$, left panel: in this case, the cyan and purple branch cross). However, both sign combinations are positive for $1 < E < 1/\mu - 1$, implying that two branches intersect the $\omega^2 = 1$ branch (e.g., see **Figure 2** for $E = 1.5$, left panel: here we find a crossing between the cyan and both purple and blue branches). Further one positive value (i.e., one intersection) follows in the narrow range $1/\mu - 1 < E < 1/\mu$, and both sign combinations are positive when $1/\mu < E$, leading again to two crossings of the $\omega^2 = 1$ branch. When $E > 1/\mu$, the branches crossing $\omega^2 = 1$ are actually the fast (blue) and Alfvén (red) branch, instead of the purple and blue branch found in **Figure 2** (for $E = 1.5$, left panel). In the limit $E \rightarrow \infty$, both sign combinations in Equation (20) lead to $k^2 = 1$, since then the fast and Alfvén branch essentially coincide, as seen from Equations (16–17). Note that the pair plasma case is again naturally contained as a special case where $E = I$. Knowing explicitly the special k -values where branches cross at parallel orientation is also especially useful when wanting to quantify the full phase and group speed variations for all branches, as done in section 4: the phase and group diagrams display intricate wave exchange occurring at these wavenumbers.

3.2. Perpendicular Propagation

At perpendicular propagation ($\lambda = 0$), one marginal frequency pair $\omega^2 = 0$ is contained in Equation (2), and this mode relates to the long wavelength Alfvén waves, which do not propagate perpendicular to the magnetic field. Another solution at perpendicular propagation is found to be $\omega^2 - 1 - k^2 = 0$.

Indeed, for $\lambda = 0$, we can factor Equation (2) into

$$\begin{aligned} & \omega^2(\omega^2 - 1 - k^2) \\ & \{\omega^6 - \omega^4(2 + E^2 + I^2 + k^2) + \omega^2[(1 + EI)^2 + k^2(1 + E^2 + I^2)] \\ & - k^2EI(1 + EI)\} = 0. \end{aligned} \quad (21)$$

In the special case of a pair plasma ($E = I$), a complete factorization can be obtained, as the final term then becomes $(\omega^2 - 1 - E^2)(\omega^4 - \omega^2(1 + E^2 + k^2) + k^2E^2)$. The latter term also factorizes in the unphysical limit where $\mu = 0$, making $I = 0$ (see further on in section 5.3), which makes the 6th order term write as $\omega^2(\omega^4 - \omega^2(2 + E^2 + k^2) + 1 + k^2(1 + E^2))$. The branch $\omega^2 - 1 - k^2 = 0$ is in plasma physics referred to as the ‘ordinary’ or O mode. The three solutions mixed up in the 6th order term are collectively named ‘extraordinary’ or E (or X , depending on the textbook at hand). In line with this terminology, and acknowledging the fact that one of these three solutions will be a high frequency electromagnetic wave, we will use the label ω_X (and the black color) for its highest frequency solution, and call ω_O (in cyan color) the mode that relates to the $\omega^2 = 1 + k^2$ branch.

Figure 2, right panel, shows the 5 solution branches for perpendicular propagation, for a hydrogen plasma case where $E = 1.5$. The marginal (red) Alfvén-related branch is mentioned in the figure for completeness. The ordinary mode $\omega^2 - 1 - k^2 = 0$ branch (cyan) starts horizontally at unit normalized frequency and goes up like $\propto k^2$. Direct comparison with its parallel counterpart (right panel) shows how the (purple and cyan) branches have changed connectivity between their long wavelength vs. short wavelength behavior. This is due to avoided crossings, that show up in oblique orientations, which we discuss next.

3.3. Oblique Orientations

At any orientation different from parallel or perpendicular, we must resort to numerical evaluation of the 5 roots in Equation (2), which is rather straightforward. **Figure 3** shows this for two representative angles for the case with $(E, \mu) = (1.5, 1/1.836)$. Its left panel took a very small angle $\vartheta = 0.001$, while the right panel has $\vartheta = \pi/3$. Comparison with **Figure 2** shows that oblique propagation demonstrates avoided crossings of branches, occurring near the special k values computed from Equation (20). The inset in the left panel shows this avoided crossing quite clearly. Note how the blue, purple and cyan branch are all affected. Indeed, when we animate the dispersion curves for this case $E = 1.5$ with angle, only the parallel case is special in its connectivity showing true branch crossings, and all other angles have the 5 branches neatly ordered in frequency: the red lies below the blue, found in turn below the purple, which is below the cyan, and the black branch always lies at the top.

This behavior is generic, and the avoided crossings occur according to the predictions on the basis of Equation (20), as is illustrated in **Figure 4** for a case where $E = 0.5$ and only one branch intersects the $\omega^2 = 1$ branch at parallel orientation. The left panel of **Figure 4** shows the avoided crossing between the cyan and purple branch, again at $\vartheta = 0.001$ (note the inset).

Since for E -values beyond $E = 1/\mu - 1$, the cut-off ordering changes as discussed before with $\omega_i^2 > \omega_c^2$, avoided crossings affect the blue branch only up to $E = 1/\mu$. Beyond $E > 1/\mu$, both the blue and red branch will demonstrate avoided crossings.

The color scheme of the branches in **Figures 2–4** is chosen in accord with the red, blue and purple resonance ranges we discussed in section 2.2 and quantified in **Figure 1**. We introduce a convenient labeling ($A/F/M/O/X$) for the 5 wave modes through ω_A (red), ω_F (blue), ω_M (purple), ω_O (cyan) and ω_X (black). Indeed, for all but parallel orientations, the red (Alfvén at small wavenumber), blue (Fast at small wavenumber) and purple branch connect to the corresponding resonance value at that angle. The parallel case is different, since then avoided crossings become true crossings, and only then the cyan branch becomes the non-propagating solution $\omega^2 = 1$, its purple branch crosses it to go to electromagnetic behavior, while the blue and red branch connect to the other two resonance frequencies. This is of particular interest, since textbook classifications of wave modes in a cold plasma rely on specific wave properties at parallel or perpendicular orientations. Our analysis suggests that it is more natural to organize the waves according to the 5 branches retained in Equation (2), and handle the special cases when branches can cross separately. In the special case of a pair plasma, a further avoided crossing with the then special (non-propagating) branch $\omega^2 = 1 + E^2$ at perpendicular orientations can occur, and needs to be accounted for (Keppens and Goedbloed, in press).

4. 5-MODE PHASE AND GROUP DIAGRAMS

4.1. Phase Diagrams

We can show the full complexity of the 5 wave mode pairs in a cold plasma in an alternative way. Up till now, we stressed the $i = 1, 2, \dots, 5$ solutions $\omega_i^2(k, \vartheta)$ for fixed ϑ , and plotted them in dispersion diagram views vs. k . We can also stress the full and intricate variation with angle ϑ , by fixing a wavenumber k . It is then insightful to plot the dimensionless phasespeed ω_i/kc for all 5 modes, and all angles. This is best done in polar plots where we show $(\omega_i/kc)\hat{\mathbf{n}}$, where $\hat{\mathbf{n}} = \mathbf{k}/k$. These are surfaces of revolution about the magnetic field direction, so that we can plot their 2D cross-section and take the magnetic field to be horizontal (i.e., $\mathbf{B} = B\hat{\mathbf{e}}_x$), making $\hat{\mathbf{n}} = (\cos \vartheta, \sin \vartheta)$. These phase diagrams then vary with wavenumber k , but the ordering of the branches discussed previously means that at all but parallel orientations, they remain always nested surfaces. The special k values where branches cross at parallel orientations, represent mode transformations occurring when the surfaces locally touch one another.

We plot some representative 5-mode phase diagrams in **Figure 5** for the hydrogen case at $E = 1.5$, shown previously in dispersion diagrams in **Figures 2, 3**. At large wavelengths (small k), we can recognize that the red and blue branch correspond to the long wavelength, low frequency Alfvén and fast magnetohydrodynamic (MHD) modes, respectively. At small wavelengths (large k), the cyan and black branches

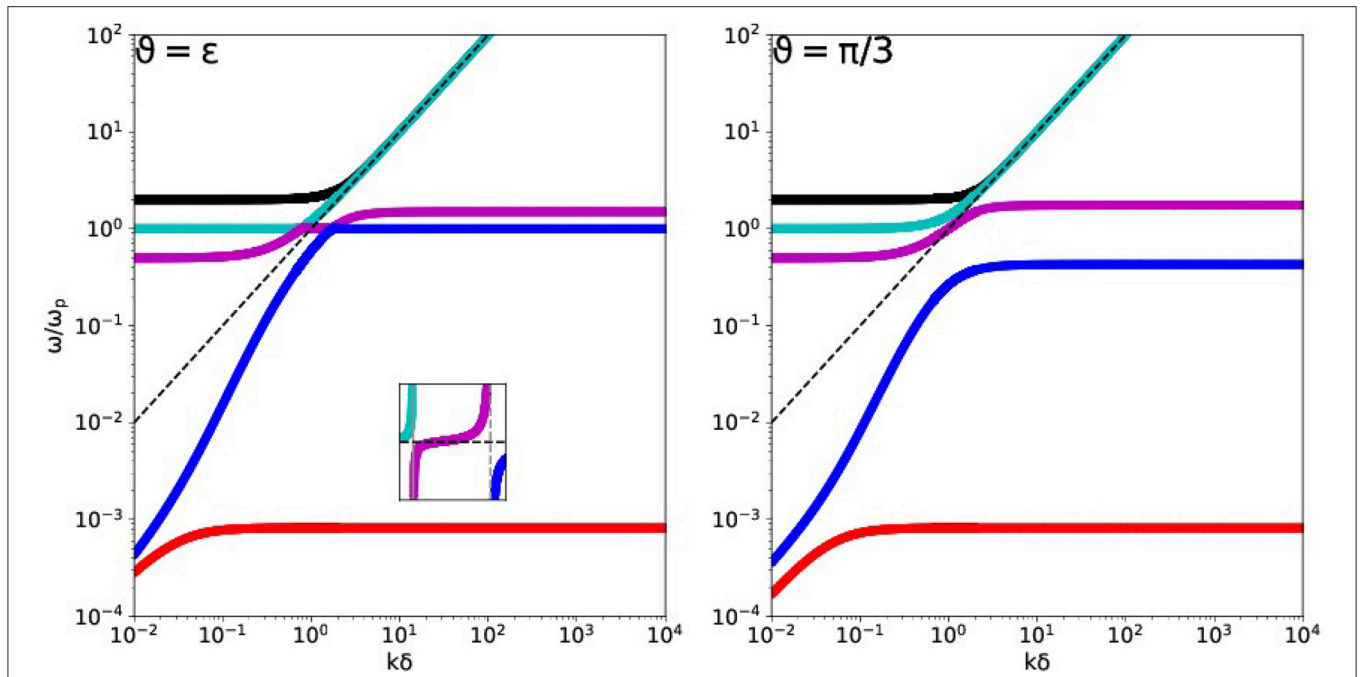


FIGURE 3 | Dispersion diagram for a hydrogen ion-electron plasma at $E = 1.5$. Shown are the 5 branches at near-parallel $\vartheta = 0.001$ (left) and oblique $\vartheta = \pi/3$ propagation. The former shows avoided crossings (see inset). The dashed black line corresponds to light speed behavior.

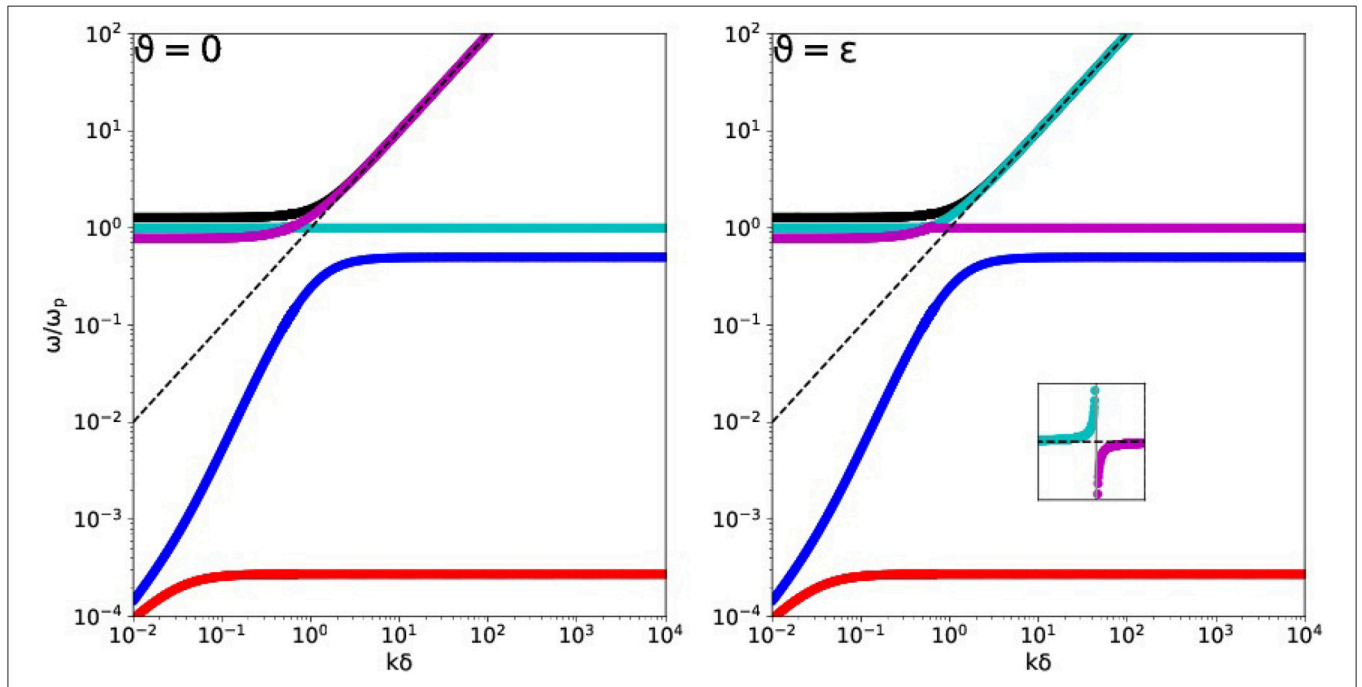


FIGURE 4 | Dispersion diagram for a hydrogen ion-electron plasma at $E = 0.5$. Shown are the 5 branches at parallel (Left) and near-parallel $\vartheta = 0.001$ (Right) propagation. The latter shows avoided crossings (see inset). The dashed black line corresponds to light speed behavior.

correspond to phase speeds nearing the light speed, and they are the familiar electromagnetic waves. Note that all branches that (partly) lie above the dashed vertical line in **Figures 2–4**

have phase speeds above the light speed c , which does not pose any physical problem as we are quantifying phase speeds here. It is even possible to identify when branches go from

superluminal to subluminal phase speeds, happening at specific $k - \vartheta$ combinations for particularly the purple branch. Animated views (available from the first author's homepage¹) of these 5-mode phase diagrams reveal their variations with wavenumber most clearly, as well as the mode transformations happening at parallel orientation, when these nested surfaces of revolution locally touch at specific k values (from Equation 20).

4.2. Group Diagrams

While the 5-mode phase diagrams provide insight in the wave transformations, the way energy is transported is quantified from the group speed. Here, the real power of the polynomial representation comes in, since starting from the dispersion relation (2) written as

$$\omega^{10} + c_4\omega^8 + c_3\omega^6 + c_2\omega^4 + c_1\omega^2 + c_0 = 0, \quad (22)$$

we can easily compute the corresponding group speed $\partial\omega/\partial\mathbf{k}$ by implicit derivation. The dependency of the coefficients $c_i(E, I; k^2, \lambda^2)$ implies that we only need to use $\partial k^2/\partial\mathbf{k} = 2\mathbf{k}$ and $\partial\lambda^2/\partial\mathbf{k} = 2\lambda/k[\hat{\mathbf{b}} - \lambda\hat{\mathbf{n}}]$ where $\hat{\mathbf{b}} = \mathbf{B}/B$ and $\hat{\mathbf{n}} = \mathbf{k}/k$. The resulting expression can be manipulated into

$$\frac{\partial\omega}{\partial\mathbf{k}} = -\lambda k \frac{P_4}{P_9} \hat{\mathbf{b}} + \left(\frac{\lambda^2 k P_4 - k P_8}{P_9} \right) \hat{\mathbf{n}}. \quad (23)$$

In this expression, we find the following polynomials

$$P_4 = \omega^4(EI - E^2 - I^2) + \omega^2[(1 + EI)EI + k^2(E^2 + I^2 - EI) - k^2E^2I^2], \quad (24)$$

$$P_8 = -2\omega^8 + \omega^6\alpha_6 - \omega^4\alpha_4 + \omega^2\alpha_2 - 2\lambda^2E^2I^2k^2, \quad (25)$$

$$P_9 = 5\omega^9 + 4c_4\omega^7 + 3c_3\omega^5 + 2c_2\omega^3 + c_1\omega. \quad (26)$$

In the P_8 polynomial, the $\alpha_{6,4,2}$ coefficients are

$$\alpha_6 = 4 + 2E^2 + 2I^2 + 2k^2, \quad (27)$$

$$\alpha_4 = 2(1 + EI)^2 + (1 + \lambda^2)(E^2 + I^2 - EI) + 2k^2(1 + E^2 + I^2), \quad (28)$$

$$\alpha_2 = (1 + EI)(1 + \lambda^2)EI + 2k^2[(1 + EI)EI + \lambda^2(E^2 + I^2 - EI)]. \quad (29)$$

These (admittedly lengthy) expressions can, however, easily be evaluated for each of the 5 branches, and this for all directions and wavenumbers. Indeed, the only thing we need to do is to compute the 5 solutions as zeros of the polynomial in Equation (22) (which is how all previous results shown were obtained), and then evaluate the RHS of expression (23) accordingly. Zeros of the P_9 polynomial may need special treatment, but those actually correspond to the double roots of the original polynomial. Since we found that the branches almost never intersect (except at the special crossings discussed previously), this situation hardly occurs. At perpendicular orientations, the Alfvén related branch is marginal $\omega_A = 0$, but this case can also be handled separately (it has zero group speed).

¹http://perswww.kuleuven.be/Rony_Keppens

The resulting 5-mode group speed diagrams are intricate, and some are shown in **Figure 6** for the same hydrogen plasma with $E = 1.5$. The variation with wavenumber provides fascinating views on the anisotropy inherent in all 5 wave modes. At the special wavenumbers where modes transform, also these diagrams show drastic deformations, which are only appreciated in animated views (available from the first author's homepage¹). In **Figure 6**, the top left panel corresponds to long wavelength behavior, where the typical Friedrichs diagram containing Alfvén (red) and fast (blue) waves is recovered, while the other three waves have all smaller groupspeeds. This is opposite to the behavior at short wavelengths, where the two electromagnetic modes (purple ω_O or O and black ω_X or X) have group speeds that approach the light circle. As group speeds, all speeds obtained with formula (23) lie within the light circle, but their relative ordering and morphology is rather complicated. This is shown at selected wavenumbers in the four panels of **Figure 6**.

5. RELATION TO TEXTBOOK TREATMENTS

5.1. Refractive Index Views

Textbook treatments (e.g., Stix, 1992; Bittencourt, 2004; Thorne and Blandford, 2017) rather emphasize that the dispersion relation Equation (2) gives direct information on the refractive index $n^2 = k^2/\omega^2$. When we consider a given frequency ω , next to the choice of (E, μ) and an orientation ϑ , the solutions to Equation (2) follow directly from

$$n^2 = \frac{\bar{B} \pm (\bar{B}^2 - 4\bar{A}\bar{C})^{\frac{1}{2}}}{2\bar{A}}. \quad (30)$$

At most two real solutions can exist at given frequency, and this is then used to classify the wave modes in various types. Frequency ranges where only one real solution is found, or no real solution at all, must be accounted for. Of course, two complex solutions can always be found, but then the waveform $\exp[i(\mathbf{k} \cdot \mathbf{x} - \omega t)]$ adopted with real ω and complex k has an evanescent, instead of propagating behavior. Using the solution (30) in Equation (2), after adding a factor $\bar{A}n^2$ on each side of the equality sign, allows one to write the solutions also as (Bittencourt, 2004)

$$n^2 = 1 - \frac{2(\bar{A} - \bar{B} + \bar{C})}{2\bar{A} - \bar{B} \pm (\bar{B}^2 - 4\bar{A}\bar{C})^{\frac{1}{2}}}. \quad (31)$$

The discriminant appearing in Equations (30-31) can be reworked to

$$\begin{aligned} \bar{B}^2 - 4\bar{A}\bar{C} &= [\omega^2(E^2 + I^2 - EI) - EI(1 + EI)]^2 \sin^4 \vartheta \\ &+ 4(\omega^2 - 1)^2(E - I)^2\omega^2 \cos^2 \vartheta. \end{aligned} \quad (32)$$

This expression is general, and can be used to rewrite Equation (31) for the special cases of parallel or perpendicular orientations to the formulae in Equations (19-21).

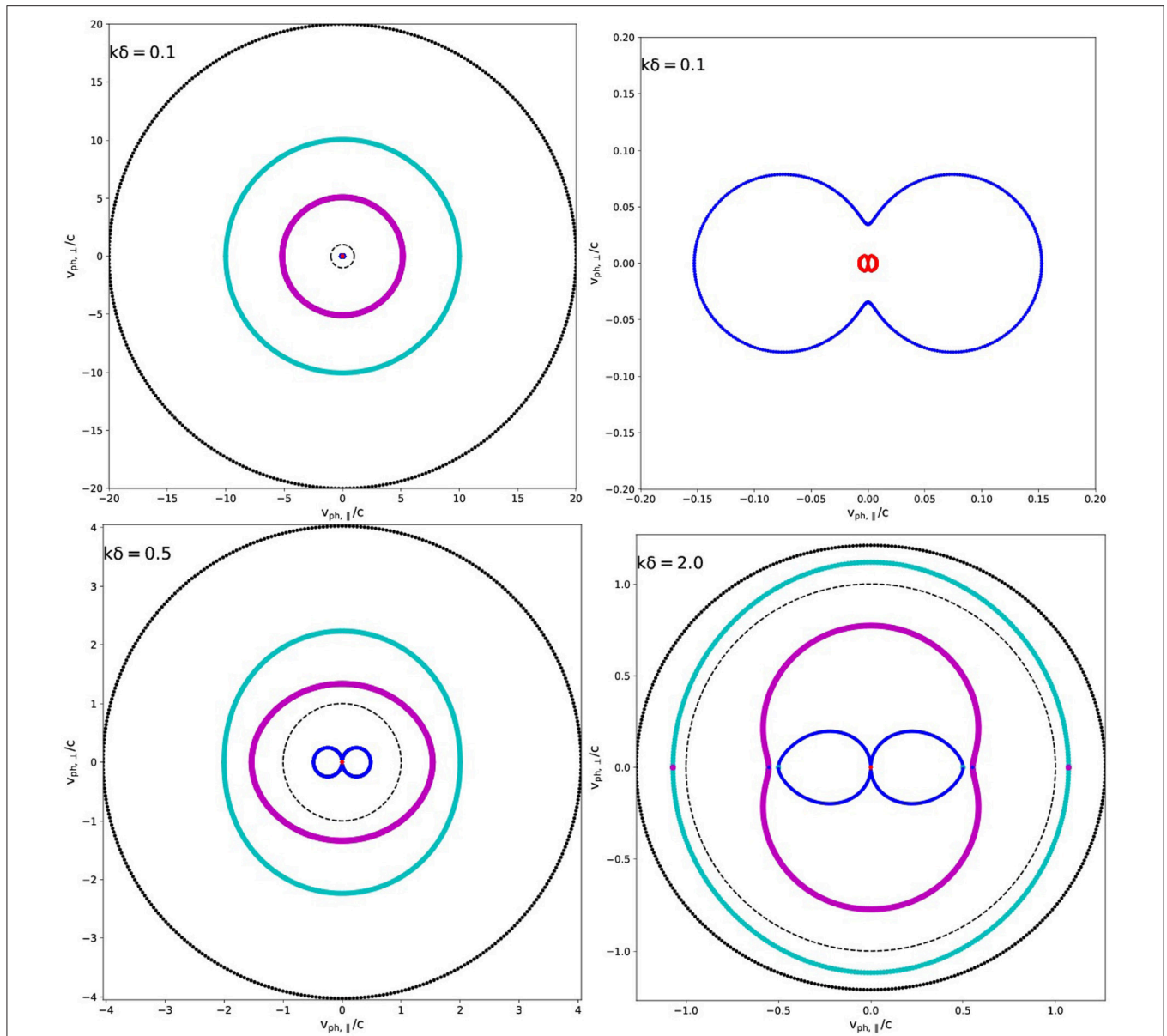


FIGURE 5 | Selected 5-mode phase diagrams at three different values of $k\delta = 0.1$ (top row), 0.5 (bottom left), 2.0 (bottom right). The dashed black circle always indicates the light circle, hence notice the varying scale. This is for a hydrogen plasma with $E = 1.5$, as in **Figures 2, 3**, and we exploit the same color scheme: electromagnetic modes are black (ω_X) and cyan (ω_C), followed by purple (ω_M or Middle), blue (ω_F or Fast) and red (ω_A or Alfvén) wave. Note that in the $k = 2$ (bottom right) plot, at parallel orientation (horizontal central line), mode exchanges occurred between the cyan, purple and blue branches. We always show all 5 branches, such that the red Alfvén branch requires further zooming in to see its details: this is done in the top right panel for the $k = 0.1$ case.

For completeness, we note that textbook treatments typically exploit the following quantities

$$R = \frac{\omega^2 - \omega(E - I) - (1 + EI)}{(\omega - E)(\omega + I)}, \tag{33}$$

$$L = \frac{\omega^2 + \omega(E - I) - (1 + EI)}{(\omega + E)(\omega - I)}, \tag{34}$$

$$P = \frac{\omega^2 - 1}{\omega^2}. \tag{35}$$

These appear in the combinations $S = (R + L)/2$ and $D = (R - L)/2$, which obey the equality $S^2 - D^2 = RL$, and together with P , they build up the cold plasma dielectric tensor, which relates the displacement vector \mathbf{D} (giving $\nabla \times \mathbf{B} = \mu_0 \partial \mathbf{D} / \partial t$, including displacement currents) to the electric field \mathbf{E} through $\mathbf{D} = \epsilon_0 \bar{\bar{\epsilon}} \cdot \mathbf{E}$, with

$$\bar{\bar{\epsilon}} = \begin{pmatrix} S & -iD & 0 \\ iD & S & 0 \\ 0 & 0 & P \end{pmatrix}. \tag{36}$$

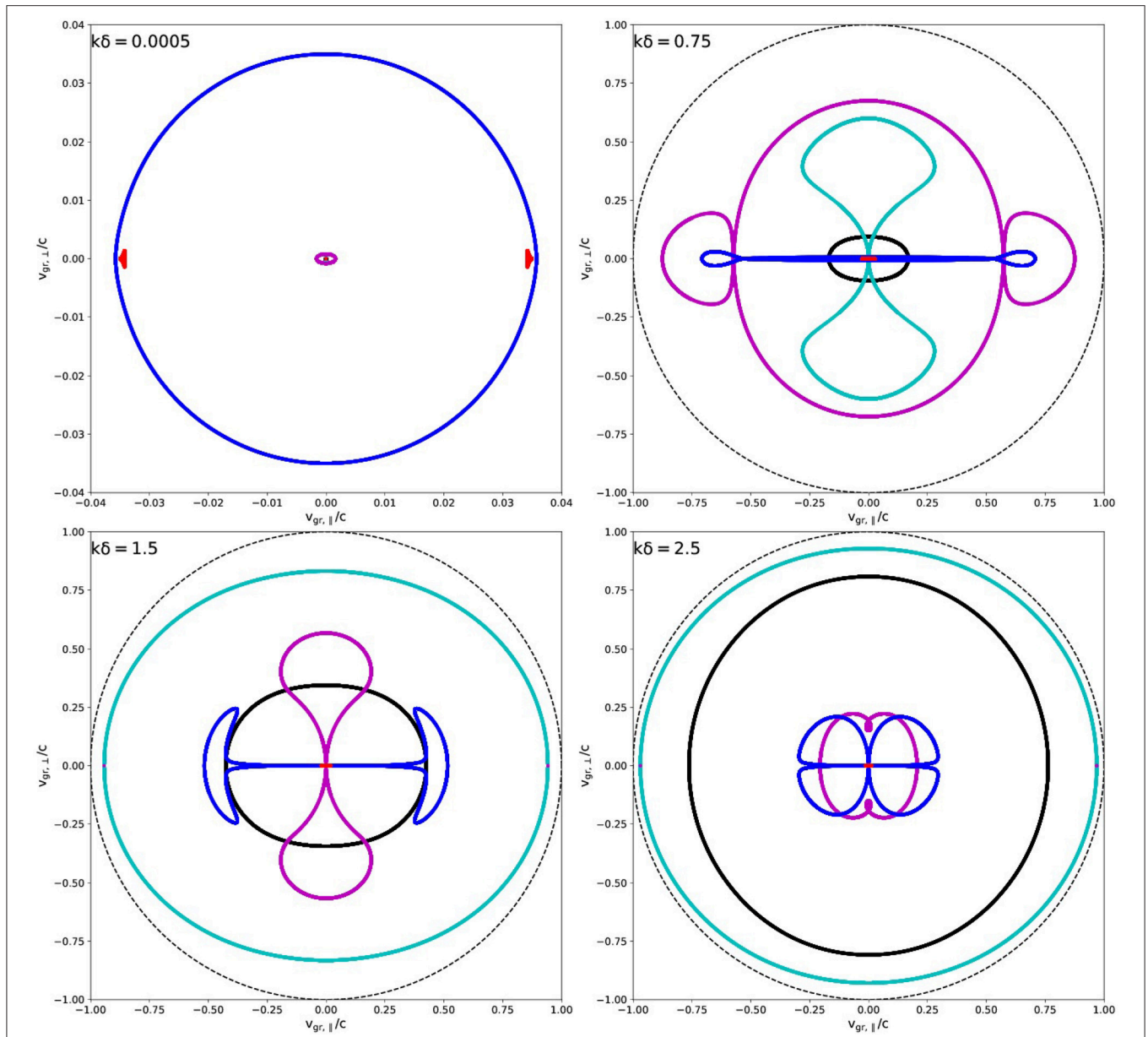


FIGURE 6 | Selected 5-mode group diagrams at four different values of $k = 0.0005, 0.75, 1.5, 2.5$. The dashed black circle (with unit radius) always indicates the light circle, notice the varying scale used in the top left plot. This is for a hydrogen plasma with $E = 1.5$, as in **Figures 2, 3, 5** phase, using the same color scheme: electromagnetic modes are black (ω_X) and cyan (ω_D), followed by purple (ω_M or Middle), blue (ω_F or Fast) and red (ω_A or Alfvén) wave.

We wrote the tensor components here with the 3rd dimension parallel to \mathbf{B} , and the \mathbf{k} - \mathbf{B} plane as the 1-3 plane. The dispersion relation, and all discussions of polarization in terms of the wave electric field orientation w.r.t. \mathbf{B} and \mathbf{k} , then follow from

$$\mathbf{k} \times (\mathbf{k} \times \mathbf{E}) + (\omega^2/c^2)\bar{\bar{\epsilon}} \cdot \mathbf{E} = \mathbf{0}. \quad (37)$$

The governing dispersion relation is still given by Equation (30), where the only difference appears as follows: $\bar{A} \rightarrow A$, where $\bar{A} = \omega^2(\omega^2 - E^2)(\omega^2 - I^2)A$, and similarly for \bar{B} and \bar{C} . In the A, B and C formulations, one can write $C = PRL, A = S \sin^2 \vartheta + P \cos^2 \vartheta$,

and $B = RL \sin^2 \vartheta + PS(1 + \cos^2 \vartheta)$. It is then customary to rewrite the governing dispersion relation as

$$\tan^2 \vartheta = \frac{-P(n^2 - R)(n^2 - L)}{(Sn^2 - RL)(n^2 - P)}. \quad (38)$$

From this latter expression, one can see that $n^2 \rightarrow \infty$ implies $\tan^2 \vartheta = -P/S$, and this latter formula gives $\bar{A} = 0$, which we used to discuss cold plasma resonances. Cut-offs on the other hand are found from $P = 0, R = 0$ or $L = 0$, and these make

$\bar{C} = \omega^2(\omega^2 - E^2)(\omega^2 - I^2)PRL$ vanish. Since further

$$D = \frac{(I - E)\omega}{(\omega^2 - E^2)(\omega^2 - I^2)}, \quad (39)$$

the pair plasma case leaves $D = 0$ (rendering the dielectric tensor diagonal), while $S = L = R$.

One may recognize the dispersion relation for parallel propagation given by Equation (19) as $n^2 = L$ and $n^2 = R$, with $P = 0$ giving the $\omega^2 = 1$ factor. The $n^2 = R$ and $n^2 = L$ correspond to right-hand and left-hand circularly polarized waves, respectively. We note again that this mixes forward and backward mode pairs from the original formulation. Similarly, the “ordinary” $\omega^2 - 1 - k^2 = 0$ branch from perpendicular propagation in Equation (21) can be written as $n^2 = P$. The remaining “extraordinary” term, can then be recognized from Equation (38) as $n^2 = RL/S$.

We can plot all dispersion diagrams in terms of the refractive index, and this is done in **Figure 7**, for exactly the same parameters as taken in **Figure 3**. Note again the avoided crossings at near-parallel orientations (left panel). The horizontal dashed lines are indicating fixed frequencies where $\log_{10} \omega = -4, -0.1, 0.1$, respectively. These return in **Figure 8** below.

5.2. CMA Related Wave Normal Surfaces

The standard way to discuss cold ion-electron plasma waves, uses the refractive index viewpoint and the solutions obtained as in Equation (30). When both solutions are real, these 2 solutions at fixed frequency lead to a first classification as being “fast” or “slow,” depending on their corresponding phase speed. Two more labels relate to the limiting behaviors at parallel and perpendicular orientations. At parallel orientations, we obtained Equation (19) which (artificially) separates 4 of the 5 mode pairs into $n^2 = R$ or right-hand circularly polarized waves, vs. $n^2 = L$ for left-hand polarized waves. A corresponding R or L label is then used to classify the mode. At perpendicular orientation, we noted that Equation (21) contains the “ordinary” (O) mode $\omega^2 - 1 - k^2 = 0$, next to the 3 “extraordinary” (E) ones mixed up in $n^2 = RL/S$. Collectively, the labels fast/slow, L/R , and O/E then serve to identify specific wave mode behavior. We note however that both classifications miss out one of the 5 branches, since the $\omega^2 = 1$ solution is left out from the L/R scheme, and the marginal $\omega_A^2 = 0$ pair is left out from the O/E scheme. To make matters worse, the 5 branches show avoided crossings as soon as one deviates from exactly parallel orientations (and in the case of a pair plasma, this is also true at exactly perpendicular orientations, see Keppens and Goedbloed, in press), and this seems unaccounted for in textbook treatments. That the branches that are left out from the L/R vs. O/E labeling do not correspond at all (one being cyan or ω_O , the other being red or ω_A for the $E = 1.5$ case used in our figures), is yet another aspect to be considered.

A way to categorize the diversity of wave modes in a cold plasma is using the classical Clemmow-Mullaly-Allis or CMA diagram (Stix, 1992). When plotting $1/\omega^2$ vs. EI/ω^2 (or variations thereof, like E^2/ω^2 or E/ω), the lines corresponding to $P = 0$, $R = 0$ or $R = \infty$ and $L = 0$ or $L = \infty$, $S = 0$ and

$RL = PS$ divide this phase-space into sixteen distinct regions. These regions correspond to differences in wave propagation characteristics, best visualized through the wave normal surfaces, which plot phase velocity $(1/n)[\vartheta]$ for all angles ϑ . The 16 regions correspond to topologically distinct morphologies of the wave normal surface plots, and they can contain zero, one or two solutions depending on the chosen frequency. In **Figure 8**, we show this CMA-related view of the wave normal surfaces for the $E = 1.5$ hydrogen plasma considered earlier, and this at three fixed frequencies such that $\log_{10}(\omega) = -4, -0.1, 0.1$, from left to right. These frequencies are also indicated in the refractive index plots shown in **Figure 7**, and we adopted the coloring scheme we introduced for labeling the 5 wave modes ω_A (red), ω_F (blue), ω_M (purple), ω_O (cyan) and ω_X (black). Note e.g., how the blue ω_F branch is in the textbook way a “fast” mode for the left panel, while it becomes a “slow” mode in the middle panel, although in all cases it corresponds to the blue branch that behaves as fast MHD waves at long wavelengths. We did not add the traditional L/R or O/E labels to the surfaces, as they confuse wave modes due to avoided crossings. Indeed, at $\vartheta = 0$ **Figure 2** (left panel) shows that at $\log_{10}(\omega) = 0.1$ one intersects the purple and blue branch which are L and R , respectively, while at any finite angle (see **Figure 2**, right panel or **Figure 3**) the cyan ω_O and purple ω_M branch matter. The $\vartheta = \pi/2$ panel from **Figure 2** has ω_O as cyan being “ordinary” or O , while ω_F , ω_X and ω_M are all “extraordinary” or E , but has no label for the Alfvén mode, while it appears as the “slow” mode in the leftmost panel of **Figure 8**. We rather opt to use the 5-mode identification which appears naturally, and handle the complications of avoided crossings separately.

5.3. Magneto-Ionic Theory and the Appleton-Hartree Relation

A well-known special case of the general cold ion-electron plasma dispersion relation is the Appleton-Hartree equation. It ignores ion motion, and assumes that one is interested in high frequency waves only. In particular, it is valid for frequencies above the electron plasma frequency ω_{pe} (the dimensional quantity where $\omega_{pe}^2 = e^2 n_e / \epsilon_0 m_e$), which should be above the ion plasma and ion gyrofrequency. The Appleton-Hartree relation describes high frequency electromagnetic waves that travel at arbitrary angle ϑ , and is used in magneto-ionic theory (Bittencourt, 2004; Thorne and Blandford, 2017).

The Appleton-Hartree relation turns out to be a rather curious limit of the general dispersion relation, where one sets the charge to mass ratio to the unphysical value $\mu = 0$ (or equivalently, where one sets $I = 0$). This limit of Equation (2) allows one to factor out a $\omega^2 = 0$ pair, and one is left with the 8th order polynomial (4th order in ω^2) given by

$$\begin{aligned} & \omega^8 - \omega^6(3 + E^2 + 2k^2) + \omega^4 [3 + E^2 + 2k^2(2 + E^2) + k^4] \\ & - \omega^2 [1 + k^2(2 + (1 + \lambda^2)E^2) + k^4(1 + E^2)] + k^4 \lambda^2 E^2 = 0. \end{aligned} \quad (40)$$

Textbook discussions (Bittencourt, 2004; Chen, 2016; Thorne and Blandford, 2017) rather write this relation in the form given

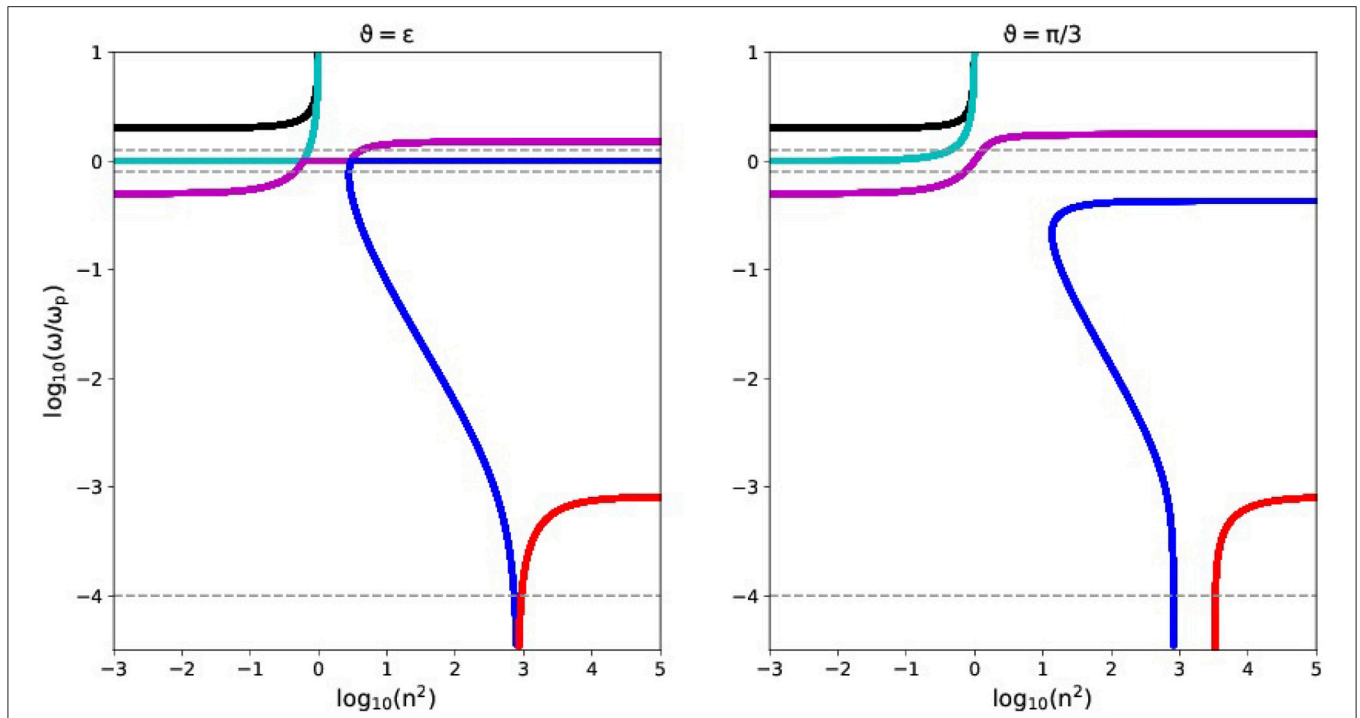


FIGURE 7 | Dispersion diagram for a hydrogen ion-electron plasma at $E = 1.5$, this time displayed in refractive index view. The parameters and panels correspond to the $E = 1.5$ hydrogen plasma at near-parallel $\vartheta = 0.001$ (left) and oblique $\vartheta = \pi/3$ propagation from **Figure 3**. The horizontal dashed lines correspond to the fixed frequencies for which **Figure 8** shows wave normal surfaces.

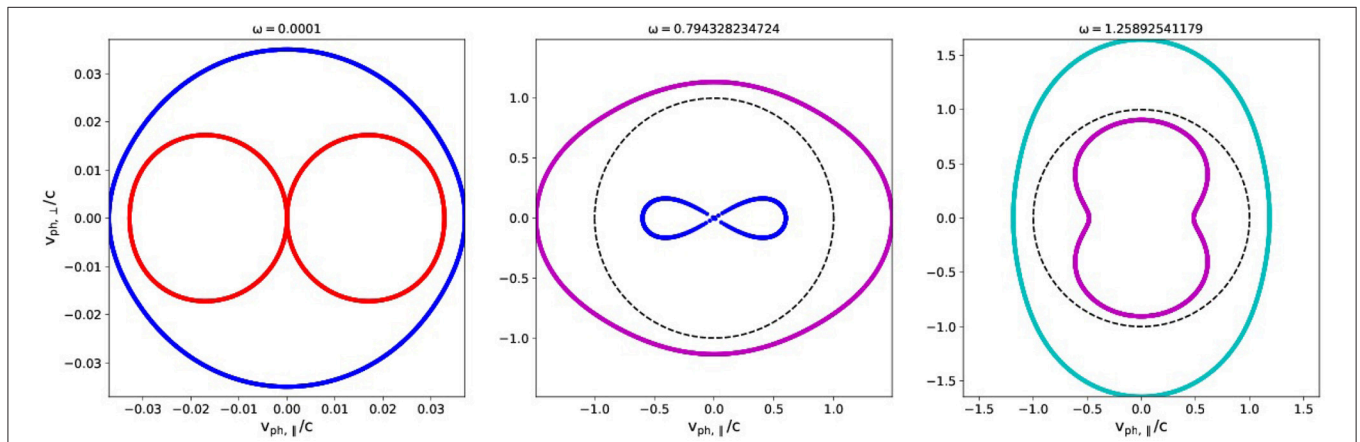


FIGURE 8 | Wave normal surfaces at fixed frequencies for the hydrogen ion-electron plasma at $E = 1.5$.

by Equation (31), where if we exploit expression (32), we can write Equation (40) also as

$$n^2 = 1 - \frac{1/\omega^2}{1 - \frac{(E^2/\omega^2) \sin^2 \vartheta}{2(1-1/\omega^2)} \pm \left(\frac{(E^2/\omega^2) \sin^4 \vartheta}{4(1-1/\omega^2)^2} + (E^2/\omega^2) \cos^2 \vartheta \right)^{1/2}} \quad (41)$$

It is clear that the polynomial form (40) is preferable, to discuss how its four solution pairs relate to the original 5 from the full dispersion relation (2). Note that the assumption of $\mu = 0$

means that the dimensionless frequency used in these expression is actually $\tilde{\omega} = \omega/\omega_{pe}$ (where previously it meant ω/ω_p).

Figure 9 shows the dispersion relation comparison between the 5 modes of the true ion-electron dispersion relation at $E = 1.5$, and the 4-mode approximation made by Appleton-Hartree. The bottom panel actually quantifies the differences in frequencies, which are hardly distinguishable in the top panel. The approximate dispersion relation (40) also has the $\omega^2 = 1$ solution for exactly parallel orientation, and we find indeed at this near-parallel angle a similar avoided crossing behavior for the 4

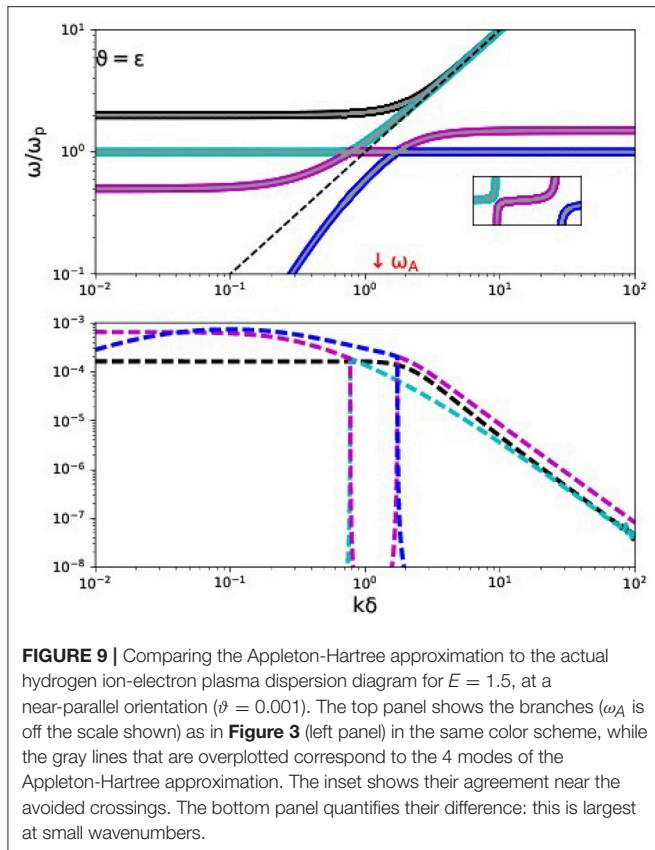


FIGURE 9 | Comparing the Appleton-Hartree approximation to the actual hydrogen ion-electron plasma dispersion diagram for $E = 1.5$, at a near-parallel orientation ($\vartheta = 0.001$). The top panel shows the branches (ω_A is off the scale shown) as in **Figure 3** (left panel) in the same color scheme, while the gray lines that are overplotted correspond to the 4 modes of the Appleton-Hartree approximation. The inset shows their agreement near the avoided crossings. The bottom panel quantifies their difference: this is largest at small wavenumbers.

retained modes. The Appleton-Hartree approximation contains the $I = 0$ limit of the crossings from Equation (20), so it has $k^2 = E/(E \pm 1)$ as special wavenumbers. Note that the Appleton-Hartree relation always discards the Alfvén related branch ω_A . The limit $\mu = 0$ obviously implies that all subtleties related to field strengths above $E = 1/\mu - 1$, which enrich the actual dispersion relation with cut-off frequencies that change their relative ordering (see section 2.1 and **Figure 1**) are completely left out of this approximation. Since we noted that beyond $E = 1/\mu$ both the ω_F and ω_A branch will demonstrate avoided crossings, this aspect can not be discussed on the basis of Appleton-Hartree. Finally, the assumption $\mu = 0$ is obviously completely inappropriate for a pair plasma, where $E = I$ and $\mu = 1$.

5.4. Faraday Rotation

A final well-known effect that is described by the (cold) ion-electron plasma dispersion relation is Faraday rotation (FR). The effect is usually described for purely parallel propagation, where as noted earlier, the dispersion relation in the form of Equation (19) is recognized as the product $(n^2 - L)(n^2 - R) = 0$. At fixed frequency ω , and paying special attention to the high-frequency electromagnetic branches (which are those for frequencies $\omega > \omega_u$ above the upper cutoff frequency from Equation 7), the solution from $n^2 - L$ is left circularly polarized (LCP), while $n^2 = R$ is right circularly polarized (RCP). The corresponding breaking n_L and n_R indices differ through their difference in wavenumber, as $n_L - n_R = c\Delta k/\omega$. A linearly

polarized wave (which can always be decomposed in a LCR and RCP wave) that travels along a magnetic field ends up with its plane of polarization rotated over a finite angle. We can quantify this in a variety of ways, e.g., by using the expressions (33, 34), we find that without any approximation, we can write

$$n_L^2 - n_R^2 = \frac{2(E - I)\omega}{(\omega^2 - E^2)(\omega^2 - I^2)}. \tag{42}$$

This expression demonstrates that for purely parallel propagation, the Faraday effect vanishes for an electron-positron plasma (see also Stewart and Laing, 1992), and that we can approximate the difference in breaking index using the Appleton-Hartree-type recipe ($\mu = 0$, high frequencies) to the textbook expression

$$n_L - n_R \approx \frac{E}{\omega^3} \left[\equiv \frac{\Omega_e \omega_{pe}^2}{\omega^3} \right], \tag{43}$$

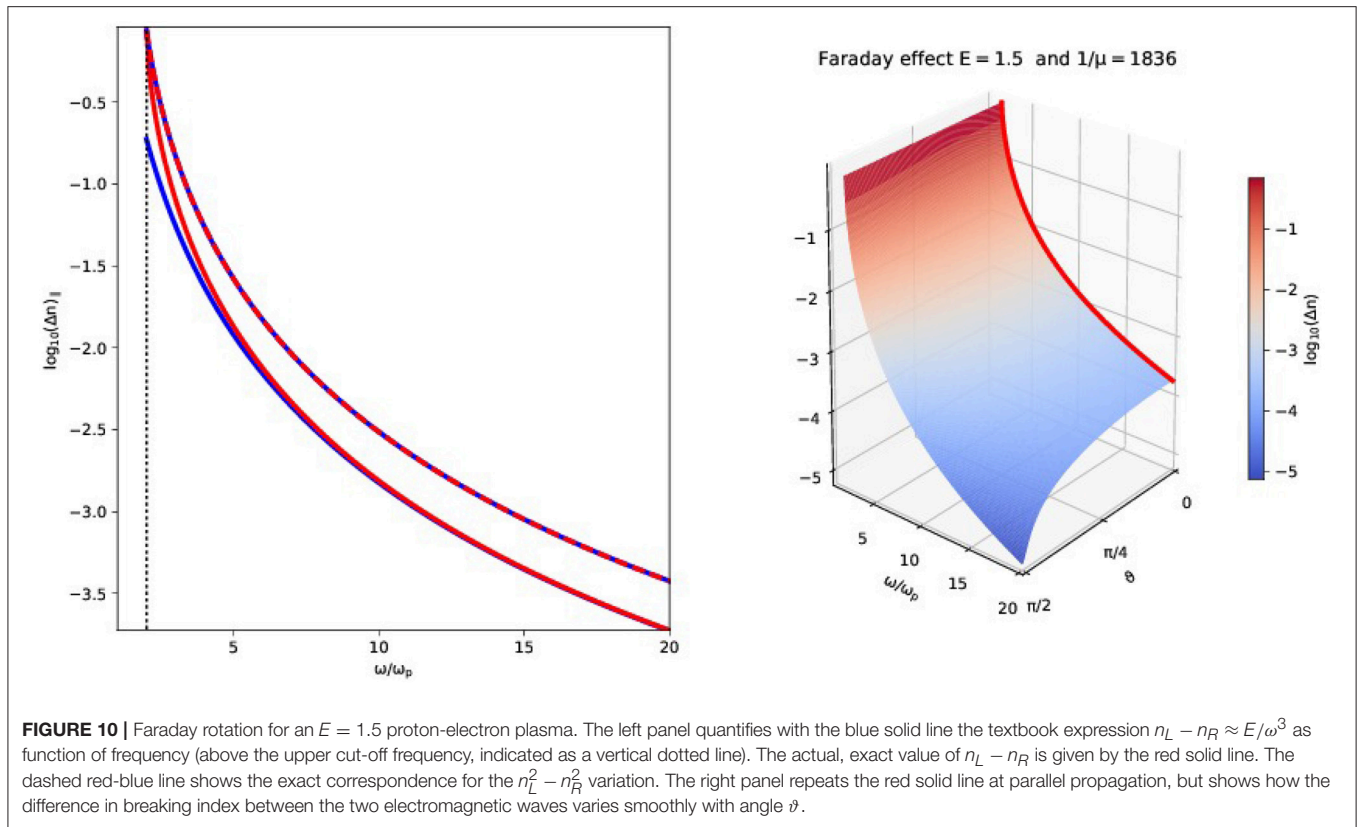
where the expression between brackets temporarily restores the dimensions. We can then use this latter expression to get the usual quantification for the change in angle χ for a linearly polarized wave, written with the so-called rotation measure RM as

$$\chi = \text{RM} \lambda^2 = \frac{e^3 \lambda^2}{8\pi^2 m_e^2 \epsilon_0 c^3} \int B_{\parallel} n_e dl, \tag{44}$$

where the integral is along the line of sight (LOS) and B_{\parallel} is the component of \mathbf{B} along it.

We note however that the proxy from Equation (43) is better written as $(E - I)/\omega^3$ to allow for the vanishing FR effect when $\mu = 1$ and propagation is along the magnetic field ($\vartheta = 0$). Moreover, we meanwhile recognize that this difference in phase speed between the high-frequency wave pair can actually be quantified readily for all angles ϑ . The FR effect for a specific $E = 1.5$ hydrogen plasma case is then shown in **Figure 10**, where the left panel shows the actual $\vartheta = 0$ variation in $n_L - n_R$ as function of frequency in red, the proxy from Equation (43) in blue, and the exact expression (42) in dashed red-blue. The right panel quantifies $\Delta n(\vartheta, \omega)$, where the parallel (red line) result is the same as in the left panel. It is clear from this plot that the variation is a smooth function of ϑ , and that one can meaningfully extend the FR quantification, as a corresponding phase speed difference at fixed frequency, for all angles (and for all local magnetic field and number density values incorporated in E). In that sense, the widely adopted RM quantification from Equation (44) is a (usually good) approximation only, since the integral quantifying the line of sight variation of the parallel magnetic field may just as well take the local $\Delta n(\vartheta, \omega; E)$ value into account when a field region with varying orientation in \mathbf{B} is traversed by a wave with fixed LOS-oriented wavevector \mathbf{k} .

This realization is important to revisit the claim that electron-positron plasmas do not show any Faraday rotation (Stewart and Laing, 1992): this statement is true for purely parallel propagation, but a quantification for all angles similar to what is shown in **Figure 10**, right panel, shows that finite to large $\Delta n(\vartheta, \omega; E)$ exist for cold electron-positron plasmas, especially near $\vartheta = \pi/2$ and at frequencies just above ω_u .



6. CONCLUSIONS

We have revisited the cold ion-electron plasma wave theory, and summarize our main findings as follows:

- The traditional means of using the plasma dielectric tensor for obtaining dispersion relations is equivalent to our polynomial-based treatment. This polynomial description emphasizes the 5 real solution pairs $\pm\omega_X$, $\pm\omega_O$, $\pm\omega_M$, $\pm\omega_F$, and $\pm\omega_A$ of forward-backward propagating waves, at fixed (real) wavenumber k . The only plasma parameters that matter are the dimensionless ratios μ and E .
- The traditional labeling of waves as fast/slow, L/R, O/E type, which rather fixes a real frequency ω and solves for (real or complex) k has some distinct disadvantages, since avoided crossings of the 5 branches actually alters the connectivity between small and large wavenumber solution branches, from those found at parallel (and sometimes also perpendicular, as in pair plasmas) behavior. The fast/slow terminology confuses the established ordering of MHD waves at low frequency, large wavenumber in slow-Alfvén-fast, which is the cornerstone of all MHD spectroscopy (Goedbloed and Poedts, 2004; Goedbloed et al., 2010, 2019). Our cold assumption has removed the slow MHD waves from the description, which will return in warm plasmas.
- We can use the polynomial representation to predict the wavenumbers for avoided crossings by Equation (20), and at fixed μ , the prevailing E value dictates which and how many branches cross. Our description is valid for all combinations

of (E, μ) , and covers especially also the high magnetization regime where $E > 1/\mu$. It covers all cold plasma waves, up to full relativistic magnetization.

- The polynomial dispersion relation form gives us a direct means to quantify and visualize the full 5-mode phase and group diagrams, for all wavenumbers k , which contain all relevant information on wave anisotropy and energy flow. These are very different from the wave normal surfaces at fixed frequencies, which are exploited in the CMA classification. In particular, they show intricate changeovers when the wavenumber crosses the special values from Equation (20).
- The Appleton-Hartree dispersion relation is a curious, unphysical limit setting $\mu = 0$, which gives satisfactory agreement on the high-frequency waves (and also contains some of the avoided crossings). It fails completely for pair plasmas, and misses all intricacy associated with $E > 1/\mu - 1$ regimes.
- Faraday rotation can be meaningfully extended to propagation angles different from $\vartheta = 0$, and the corresponding exact value of the refractive index difference $\Delta n(\vartheta, \omega; E)$ depending on angle, frequency and E can easily be quantified and used for rotation measure computations.

Further work should discuss the full variation of the wave polarizations, based on the 6×6 matrix formulation exploited in Goedbloed and Poedts (2004) and Goedbloed et al. (2019), which led to the polynomial dispersion relation. Also, the effects of a warm plasma can be easily incorporated, since then a 6th order polynomial in ω^2

enters, bringing in the slow MHD modes. This is left for future work.

DATA AVAILABILITY

The datasets generated for this study are available on request to the corresponding author.

AUTHOR CONTRIBUTIONS

RK performed the analysis of the dispersion relation and produced the figures for this article. It is based on the general

dispersion relation derived by HG, and simplified here to cold plasmas. Both authors iterated the paper and jointly improved the presentation.

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REFERENCES

- Bellan, P. M. (2012). Improved basis set for low frequency plasma waves. *J. Geophys. Res.* 117:A12219. doi: 10.1029/2012JA017856
- Bittencourt, J. A. (2004). *Fundamentals of Plasma Physics. 3rd Edn.* New York, NY: Springer-Verlag.
- Boyd, T. J. M., and Sanderson, J. J. (2003). *The Physics of Plasmas.*
- Chen, F. F. (2016). *Introduction to Plasma Physics and Controlled Fusion. 3rd Edn.* Springer.
- Damiano, P. A., Wright, A. N., and McKenzie, J. F. (2009). Properties of Hall magnetohydrodynamic waves modified by electron inertia and finite Larmor radius effects. *Phys. Plasmas* 16:062901. doi: 10.1063/1.3142479
- Denisse, J. F., and Delcroix, J. L. (1961). *Théorie des Ondes Dans Les Plasmas.*
- Goedbloed, J. P., Keppens, R., and Poedts, S. (eds.). (2010). *Advanced Magnetohydrodynamics.* Cambridge, UK: Cambridge University Press.
- Goedbloed, J. P., Keppens, R., and Poedts, S. (eds.). (2019). *Magnetohydrodynamics of Laboratory and Astrophysical Plasmas.* Cambridge, UK: Cambridge University Press.
- Goedbloed, J. P. and Poedts, S. (2004). *Principles of Magnetohydrodynamics.* Cambridge University Press.
- Ishida, A., Cheng, C. Z., and Peng, Y.-K. M. (2005). Properties of low and medium frequency modes in two-fluid plasma. *Phys. Plasmas* 12:052113. doi: 10.1063/1.1905604
- Keppens, R., and Demaerel, T. (2016). Stability of ideal MHD configurations. I. Realizing the generality of the G operator. *Phys. Plasmas* 23:122117. doi: 10.1063/1.4971811
- Keppens, R., and Goedbloed, H. (in press). Wave modes in a cold pair plasma: the complete phase and group diagram point-of-view. *J. Plasma Phys.*
- Stewart, G. A. and Laing, E. W. (1992). Wave propagation in equal-mass plasmas. *J. Plasma Phys.* 47, 295–319. doi: 10.1017/S002377800024247
- Stix, T. H. (1992). *Waves in Plasmas.* New York, NY: American Institute of Physics.
- Stringer, T. E. (1963). Low-frequency waves in an unbounded plasma. *J. Nuclear Ener.* 5, 89–107. doi: 10.1088/0368-3281/5/2/304
- Thorne, K. S., and Blandford, R. D. (2017). *Modern Classical Physics: Optics, Fluids, Plasmas, Elasticity, Relativity, and Statistical Physics.* Princeton, NJ: Princeton University Press.
- Zhao, J. (2015). Dispersion relations and polarizations of low-frequency waves in two-fluid plasmas. *Phys. Plasmas* 22:042115. doi: 10.1063/1.4919257
- Zhao, J. (2017). Properties of whistler waves in warm electron plasmas. *Astrophys. J.* 850:13. doi: 10.3847/1538-4357/aa906f

Conflict of Interest Statement: The authors declare that the research was conducted in the absence of any commercial or financial relationships that could be construed as a potential conflict of interest.

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